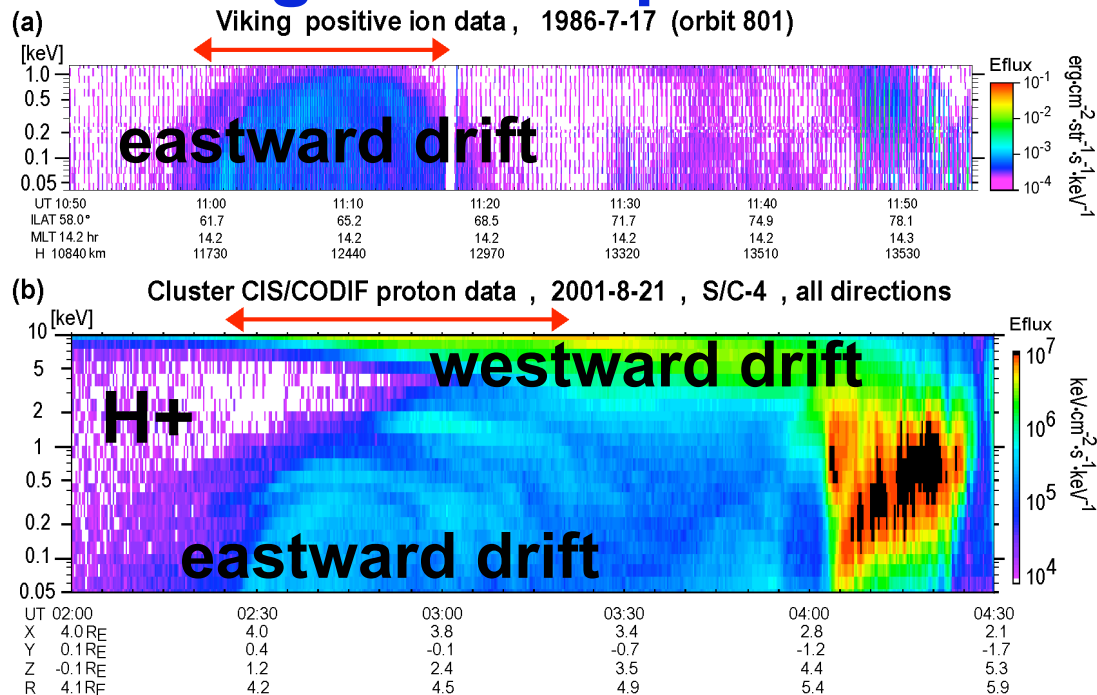


# Sub-keV Ring Current Ions: Source, Transport, and O<sup>+</sup>/H<sup>+</sup> difference

*M. Yamauchi, R. Lundin, H. Nilsson, S. Arvelius (IRF-Kiruna),  
Y. Ebihara (NIPR),  
and **Cluster-CIS** team*

## "Wedge-like dispersion"



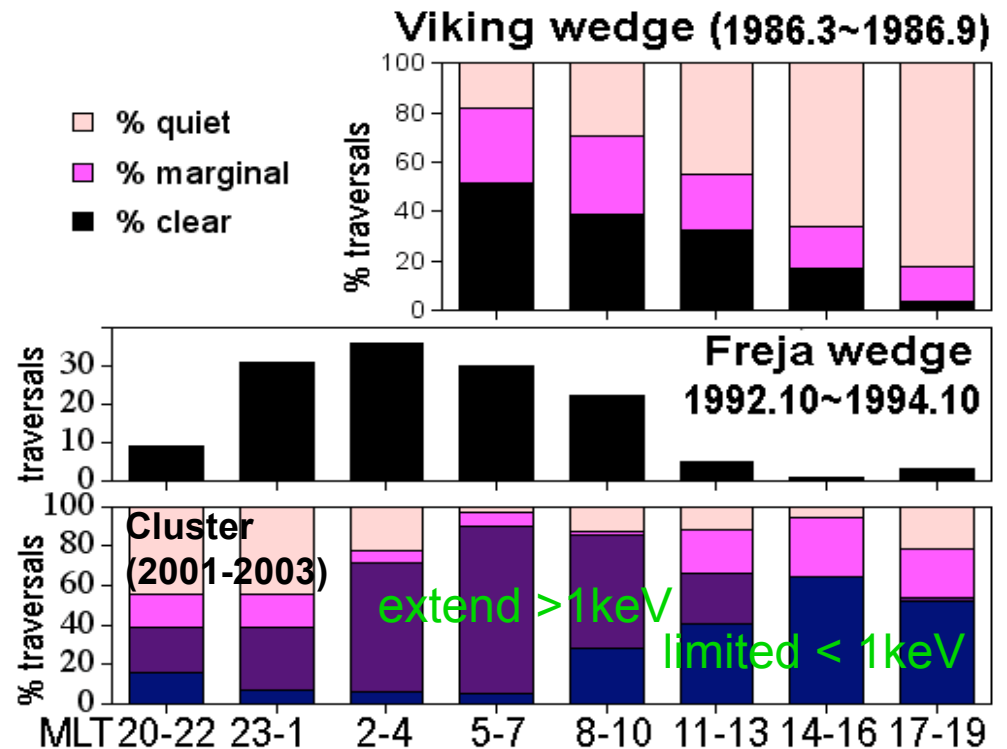
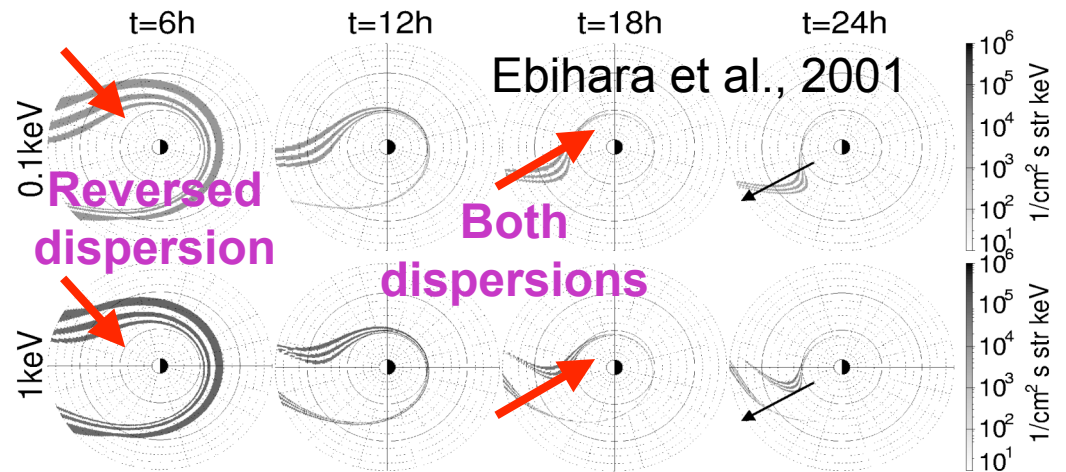
Sub-keV trapped ions are seen almost all satellites at around L=4-6. They are **wedge-like** energy-latitude dispersed as shown in both Viking data (mid-altitude) and Cluster data (equatorial plane). They are trapped ions drifting eastward, i.e., the ExB drift (including corotation) is stronger than the magnetic drift ( $\nabla B$  and curvature) at this energy range.

We show (1) **Viking statistics**  
(2) **Cluster event studies.**

# Previous Works

*sub-keV ion precipitation @ subauroral region):*

- Aureol 1 (400~2500km): Sauvaud et al., 1980  
00-06 MLT: increases after substorms.
- \* DMSP F6/F7 (800 km): Newell & Meng, 1986  
0830 MLT: correlated with Kp with some hours delay, and event may last a day.
- \* Viking (2~3  $R_E$ ): Yamauchi et al., 1996a,b  
**"Wedge-like dispersed structures"**  
modulation by pc-5 pulsation.
- \* Simulation: Ebihara et al., 2001  
**drift** model ( $\text{ExB}$ ,  $\nabla|B|$ , and curvature)  
**many hours after nightside injection**  
dispersion patterns + MLT dependence.
- \* Freja/Viking/Cluster: Yamauchi et al., 2005  
**morning peak**  
**O+ at low-altitude / H+ at high-altitude**
- \* others: Shelley et al., 1972; Chappel et al., 1982



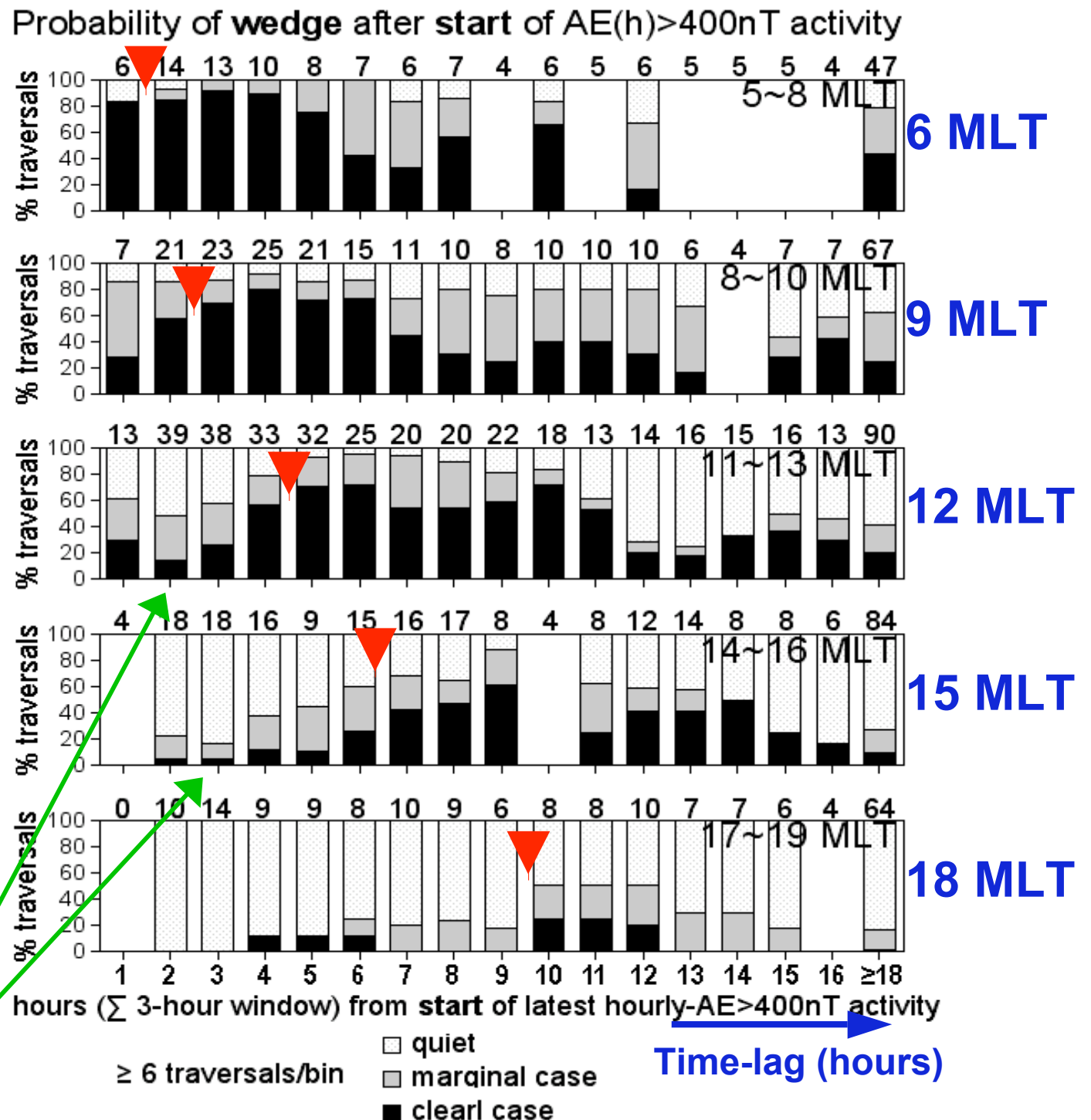
# (1) Viking

## backward superposed epoch analyses

Probabilities of observing the **wedge-like structure** after the start of AE activity. Probability is calculated from numbers of traversals with/without the structure for each 3-hr bin (3-hr running sum) for each 3-hr MLT bin.

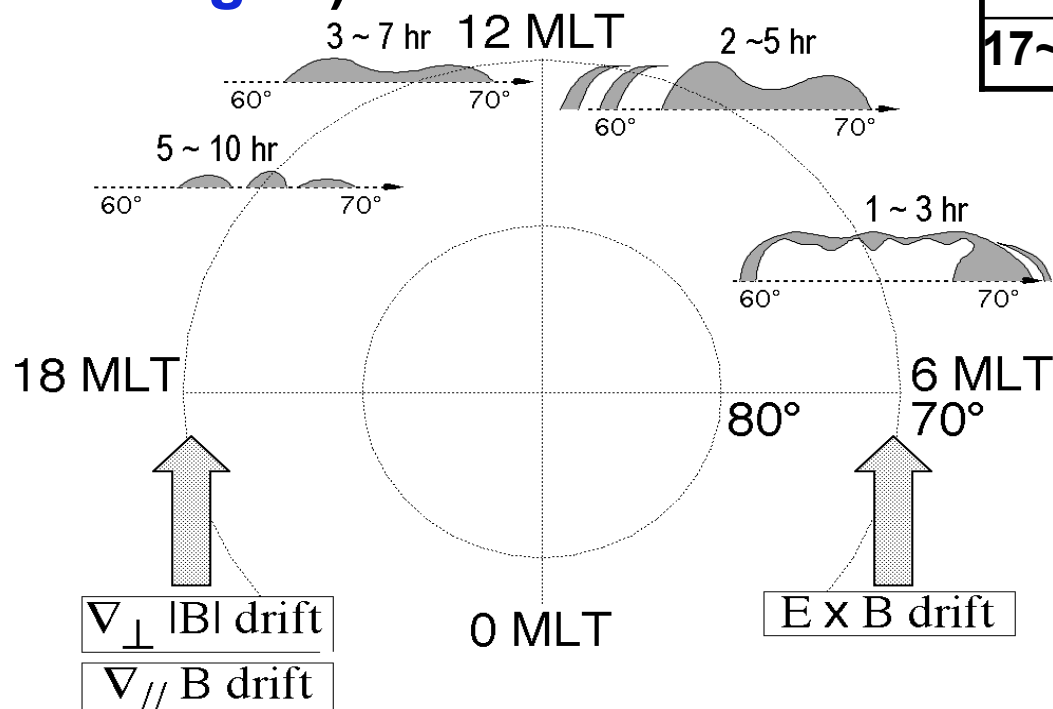
**The peak probability moves eastward**, while the peak value of the probability decrease as the peak moves eastward.

Evacuation is seen (the probability is even lower than asymptotic one)



# Viking Summary

The wedge-like structure drifts eastward, and is a fossil of substorm activity (**model is right!**). Decay time is several hours (**charge exchange model is right!**).



| MLT   | Minimum<br>Quiet case<br>After end<br>of 300 nT<br>activity | Asymptotic<br>Quiet case<br>After end of<br>300 nT activity | Maximum<br>Clear case<br>After start<br>of 400 nT<br>activity |
|-------|-------------------------------------------------------------|-------------------------------------------------------------|---------------------------------------------------------------|
| 5~7   | 1~3h (0%)                                                   | 8~9h (30%)                                                  | 0~3h (85%)                                                    |
| 8~10  | 2~3h (5%)                                                   | 9~10h (50%)                                                 | 2~4h (75%)                                                    |
| 11~13 | 3~5h (10%)                                                  | 10~11h (70%)                                                | 4~6h (70%)                                                    |
| 14~16 | 4~6h (35%)                                                  | 12~13h (80%)                                                | 6~7h (50%)                                                    |
| 17~19 | 6~8h (50%)                                                  | 14~16h (100%)                                               | 10h (25%)                                                     |

**However, it appears much earlier than prediction,** suggesting that a substantial amount of "wedge" might be formed in the morning sector during substorms. **We need to identify the source location from event study.**

## We have several possibilities

| Ion source | dispersion | scenario    |
|------------|------------|-------------|
| night      | night      | No !        |
| night      | morning    | (A)         |
| morning    | morning    | (B) (C) (D) |

- (A) Strong electric field push ions quickly.
- (B) Scattering of <10 keV ions
- (C) Energetic ions precipitate and sputter ionospheric ions into the space.
- (D) Unknown local energization process.

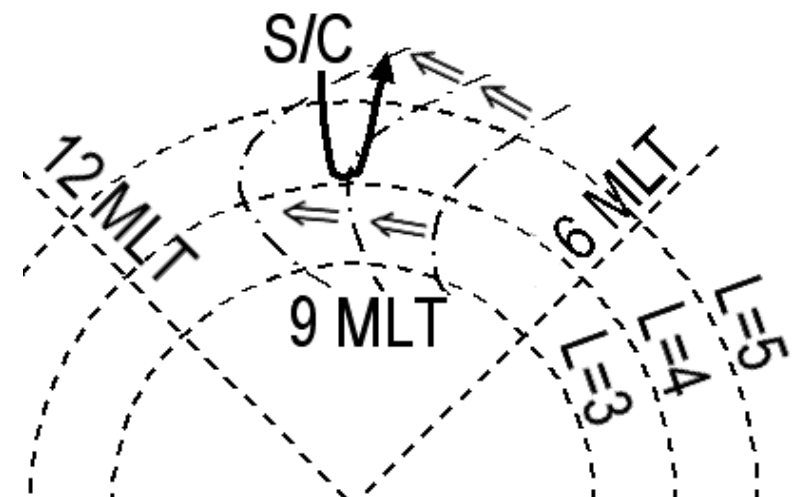
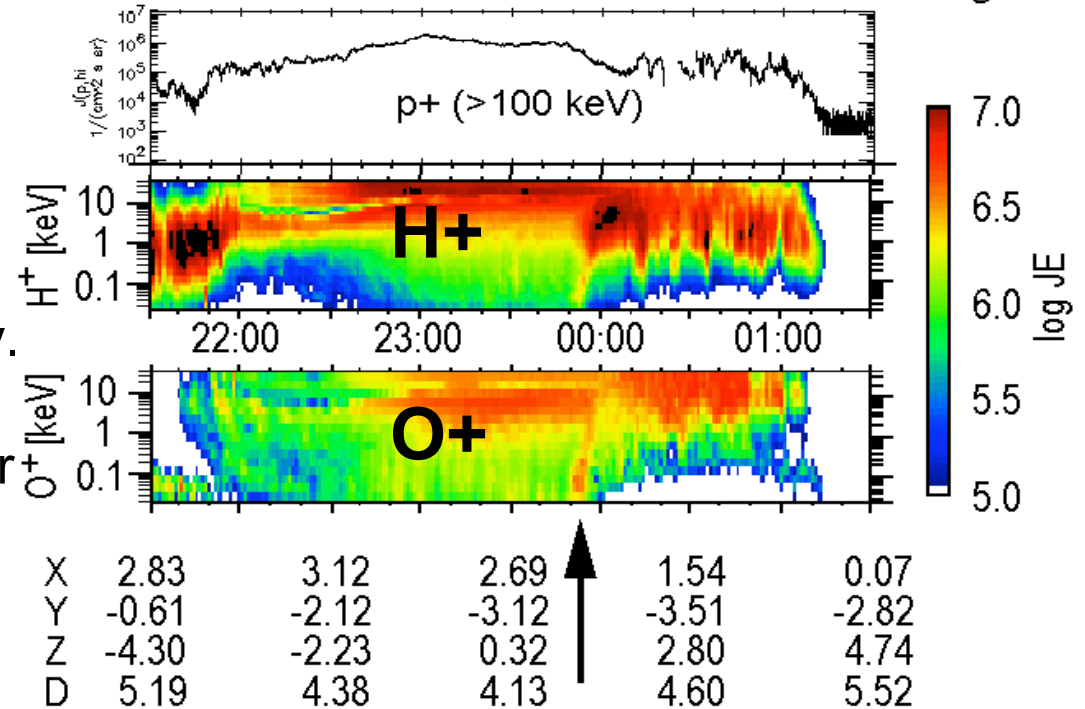
⇒ Need to find events when the wedge is formed during a substorm.

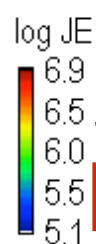
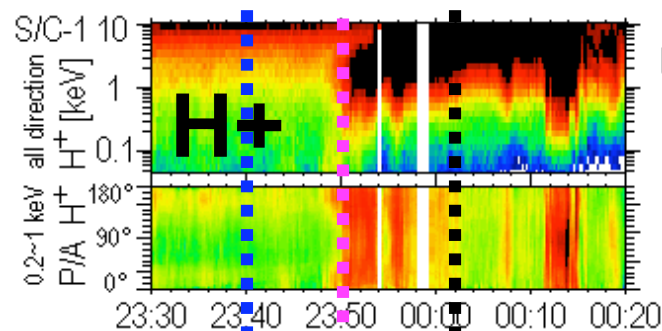
⇒ We found one case. **Wedge is seen only at outbound.**

⇒ **Case study!**

## (2) Cluster

Cluster CIS/CODIF (energy flux) & RAPID (flux)  
SC-4 21-22 October 2001 .  $4\pi$  average





S/C-1

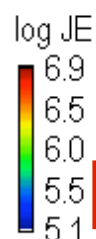
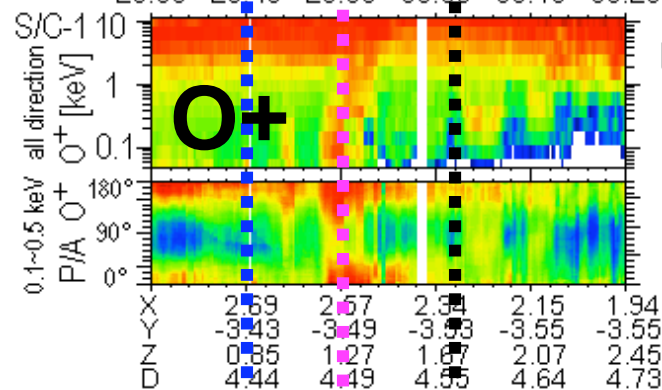
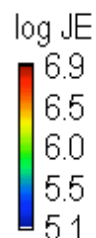
No

Yes

No

Yes

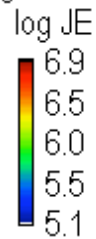
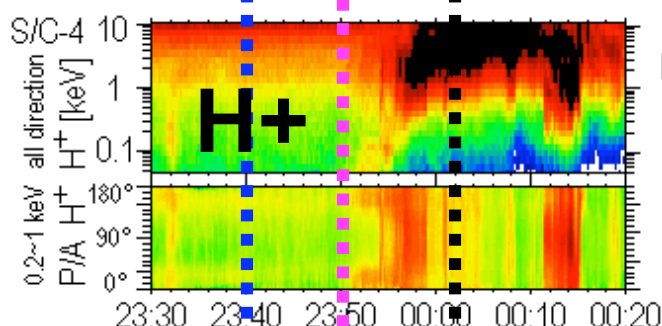
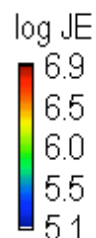
S/C-4



No

Yes

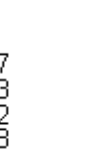
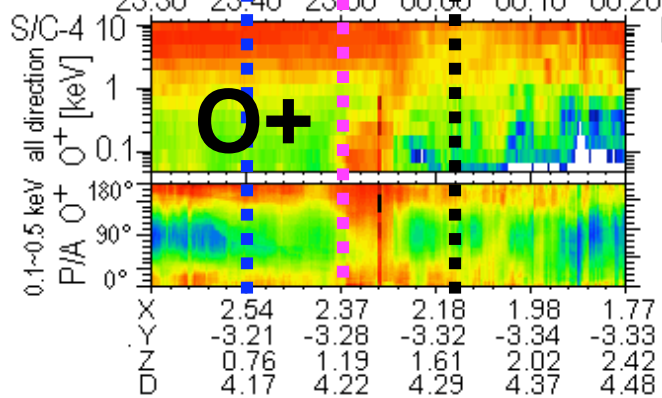
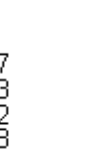
S/C-4



No

Yes

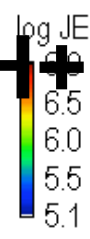
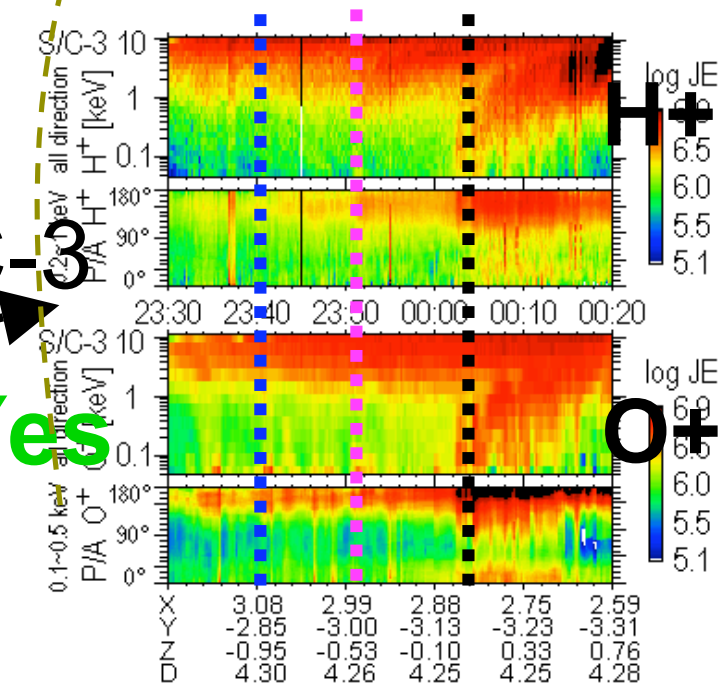
S/C-3



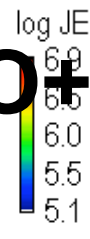
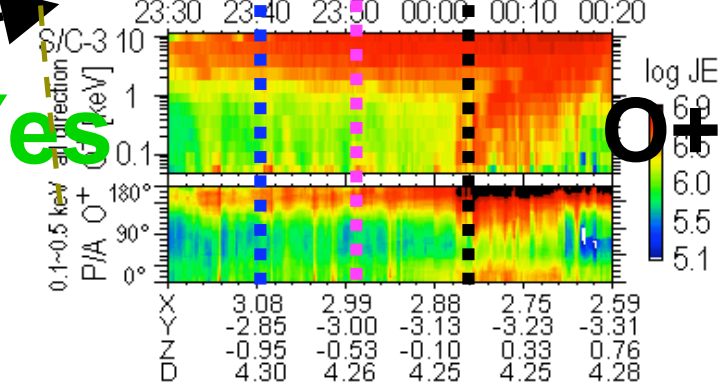
No

Yes

S/C-3



H+



O+

2001-10-21

23:40-24:00 UT

Relative S/C position:  
all at  $9.0 \pm 0.1$  MLT

S/C-3

?

Yes

No

No

No

No

No

No

# Observation summary

S/C-1 (23:45 UT), S/C-4 (23:50 UT), and S/C-3 (23:40 UT) passed through the same magnetic flux tube at 9 MLT ( $L \approx 4$ ).

**Wedge-like dispersion at 23:50 UT.**  
**No low-energy signature at 23:40 UT.**

Butterfly-trapped distribution

⇒ Bounce inside the geomagnetic bottle.

⇒ **Difference between 23:40 UT and 23:50 UT in the same flux tube means an temporal variation** although observation is made in the opposite hemisphere.

**VE**: eastward  $\mathbf{E} \times \mathbf{B}$  drift speed = energy independent, MLT dependent

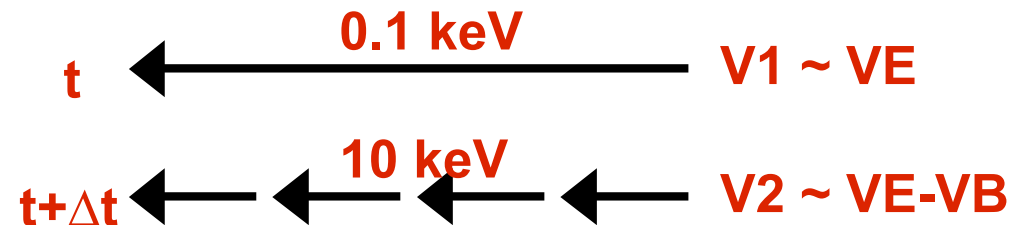
**VB**: westward magnetic ( $\nabla|B|$ +curvature) drift velocity = energy dependent

**VE  $\gg$  VB at low energy (<100 eV)** and **VE  $\sim$  VB at high energy** (value depends on E-field strength). From dispersion curve, the last-coming ions are 10-20 keV. Therefore, **VE  $\sim$  VB at 20 keV in the present case.**

# time-of-flight principal

$$V1 = VE - VB \sim VE @ 0.1 \text{ keV}$$

$$V2 = VE - VB \ll VE @ 10 \text{ keV}$$



$$V1 \cdot t = V2 \cdot (t + \Delta t) \quad \text{or}$$

$$(t + \Delta t) / \Delta t = V1 / (V1 - V2)$$

$$\sim VE / VB \quad (\text{note : } VB @ 10 \text{ keV})$$

$$= (E/B) \cdot (q \cdot R \cdot B / 3 \cdot W \cdot g)$$

$$\sim E [\text{mV/m}] / g \quad \text{or}$$

$$t = \Delta t \cdot E [\text{mV/m}] / g - \Delta t$$

for observation near equatorial plane, where

E and B are the field strengths,

q is the charge,

R = 4 RE is the geocentric distance,

W = 10 keV is the ion energy, and

g  $\sim$  1, 0.9 & 0.7 for 90°, 40° & 0° pitch angles

# Dispersion analysis

Pitch angle of the "wedge" is about  $40\sim 90^\circ$   
 $(g=0.9\sim 1.0) \Rightarrow t \leq (1.1 \cdot E[\text{mV/s}] - 1) \cdot \Delta t$

Electric field is  $1\sim 3 \text{ mV/m}$  for half an hour  
 $\Rightarrow t = 0.1\sim 2.3 \cdot \Delta t$  &  $VE = 3\sim 10 \text{ km/s}$

(a)  $0.1 \text{ keV}$  @  $23:50 \text{ UT}$ , S/C-1  $\Leftrightarrow$

Nothing @  $23:40 \text{ UT}$ , S/C-3 : temporal change

(b)  $10 \text{ keV}$  @  $23:53 \text{ UT}$ , S/C-1  $\Leftrightarrow$

Nothing @  $23:40 \text{ UT}$ , S/C-3 : temporal change

(c)  $0.1 \text{ keV}$  @  $23:50 \text{ UT}$ , S/C-1  $\Leftrightarrow$

$10 \text{ keV}$  @  $23:53 \text{ UT}$ , S/C-1 : temporal or spatial

Combination of (a)+(b) : it is temporal change

$\Rightarrow \Delta t < 13 \text{ min} \Rightarrow t < 30 \text{ min before } 23:40 \text{ UT}$

$\Rightarrow \text{drift distance} = VE \cdot t < 20000 \text{ km}$

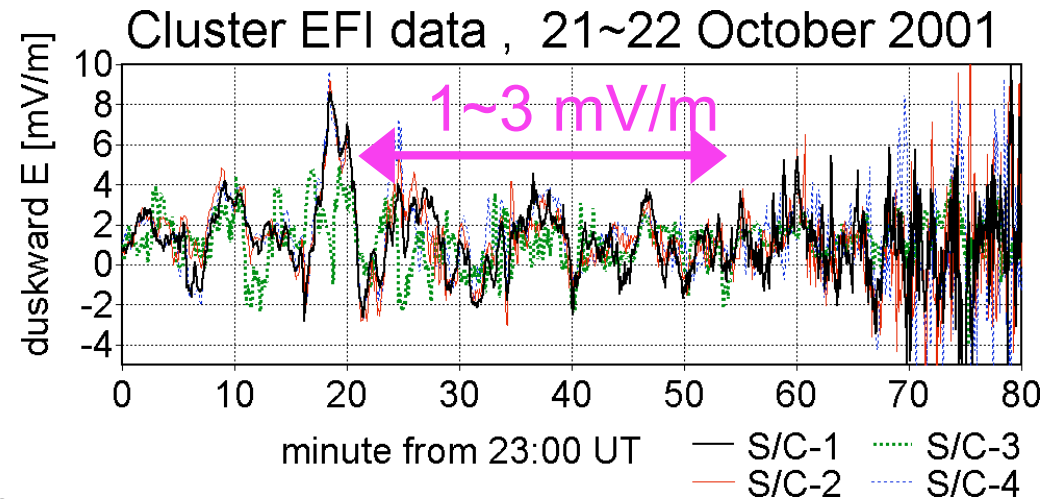
$\Rightarrow \text{dispersion started at } 7\sim 9 \text{ MLT.}$

Combination of (b)+(c) : if temporal

$\Rightarrow \Delta t \sim 3 \text{ min} \Rightarrow t = 0.5\sim 8 \text{ min before } 23:50 \text{ UT}$

$\Rightarrow \text{drift distance} = VE \cdot t = 100\sim 5000 \text{ km}$

$\Rightarrow \text{dispersion started at } 8\sim 9 \text{ MLT.}$



On the other hand, we observed O+ "wedge" at  $0.05\text{--}0.3 \text{ keV}$  ( $20 \text{ km/s} \sim 50 \text{ km/s}$ ). The  $0.05 \text{ keV}$  O+ takes  $20\sim 30 \text{ min}$  to travel from the ionosphere to the Cluster location along B in best case. From this:

(1) **Source timing is about 20~30 min before, i.e., at 23:20~23:30 UT.**

(2) **The combination (b)+(c) cannot be true**, i.e., the observed dispersion is mostly the spatial structure.

(3) O+ pitch angle is uni-direction, i.e., should **not have been mirror-bounced, endorsing point (1).**

# H<sup>+</sup>/O<sup>+</sup> differences

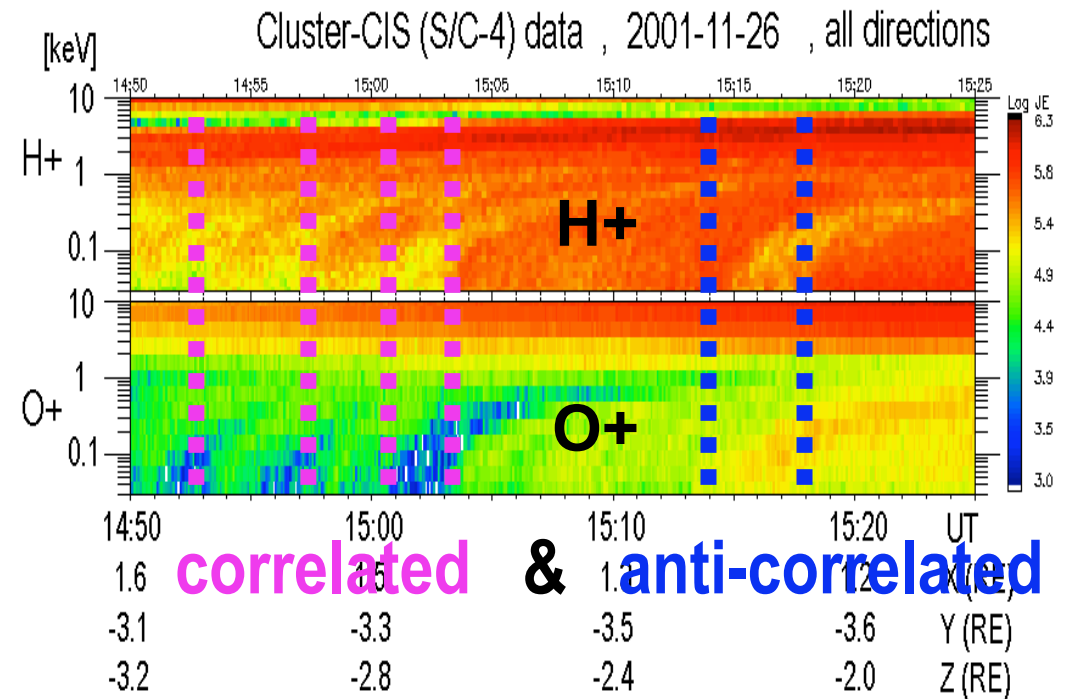
## O<sup>+</sup> motion $\neq$ H<sup>+</sup> motion

The 2001-10-21 event showed a clear O<sup>+</sup>/H<sup>+</sup> difference inside the wedge, with H<sup>+</sup> bounce-averaged feature (with butterfly pitch-angle distribution), whereas O<sup>+</sup> is not bounce-averaged.

Statistically the wedge-like structure is O<sup>+</sup> rich at low-altitudes (Freja) whereas it is H<sup>+</sup> rich as high-altitudes (Cluster).

These fact suggests that **O<sup>+</sup> source could be different from H<sup>+</sup> source.** We found couple of good Cluster examples that endorse this idea

## O<sup>+</sup> source $\neq$ H<sup>+</sup> source



Correlation part means that H<sup>+</sup> and O<sup>+</sup> has the same bounce-average drift motion. Then, **how can we understand the anti-correlation part just 15 minutes later?**

# Summary and conclusions

- (1) The dispersion might start in the morning for a substantial numbers of the wedge-like structure. This is suggested by the local time distribution, superposed epoch analyses, and a case study on the 2001-10-21 event (**source <30 min, <3 RE distance**).
- (2) Pitch-angle distribution, particularly for O<sup>+</sup>, suggest ionospheric source (consistent with morning source).
- (3) In addition to the altitude dependence of the O<sup>+</sup>/H<sup>+</sup> ratio, O<sup>+</sup> are sometimes behaving in a different way from H<sup>+</sup>.
  - \* non-bounce-average feature (2001-10-21 event).
  - \* correlation and anti-correlation in a single traversal. ⇒ ???

**Future task : understand the source of the wedge-like structure for both O<sup>+</sup> and H<sup>+</sup>. This final target is still far away.**