

KIRUNA GEOPHYSICAL DATA

DATA SUMMARY 00/7-9 JULY-SEPTEMBER 2000

Collected at SWEDISH INSTITUTE OF SPACE PHYSICS KIRUNA, SWEDEN



Kiruna Geophysical Data

Data Summary 00/7-9 July - September 2000

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INTRODUCTION

"Kiruna Geophysical Data" is a collection of results of ground-based more or less continuous measurements of upper atmosphere variables carried out at the Swedish Institute of Space Physics. Our objective with this data summary is to present current geophysical data, related to polar upper atmospheric conditions, to those who require it for their re-search work.

Coordinates of the Swedish Institute of Space Physics are:

Geogr.		Geomag	gn. Dipole	Corr. Ge	Corr. Geomagn.		
Lat.	Long.	Lat.	Long.	Lat.	Long.		
67.8°N	20.4°E	65.2°N	116.0°E	64.8°N	104.2°E		

The following ground measurements of upper atmospheric parameters have been conducted at the Swedish Institute of Space Physics in the period for which data is presented in this report.

Geomagnetism	Optical aurora	Ionospheric ionization
Geomagnetic standard recording	1. All-sky camera	1. Ionosonde 0.5-16.0 MHz
2. Storm recording		2. Riometer 30.0 MHz
Geomagnetic digital recording		3. Riometer 38.2 MHz

Ionosonde data obtained at the Swedish Institute of Space Physics in Kiruna is published together with ionosonde data from Lycksele and Uppsala in a monthly report "Ionospheric Data Sweden" which is available from the Swedish Institute of Space Physics.

The Ionospheric Observatory at Lycksele (64.7°N, 18.8°E) is also operated by the Swedish Institute of Space Physics. Mainly magnetic, riometer, photometer, all-sky cam-era and ionosonde measurements are conducted there.

The digital magnetic recordings are reproduced in this report instead of the normal magnetograms from January 1980.

Universal time is used in all tables.

The most recent magnetometer and riometer data are now accessible via internet at http://www.irf.se. For more information contact the editor (victoria.barabash@irf.se).

GEOMAGNETIC DATA

Magnetograms

The magnetograms reproduced here are plots of data recorded by a three component digi-tal fluxgate magnetometer with a sampling frequency of 1 sample/second.

The scaling is indicated by the tick mark values on the plots. Note that for large disturb-ances the normal scaling has been reduced. The positive direction of the components are: X north, Y east and Z down, in geographic coordinates.

K-indices

The K-indices give a measure of the amount of disturbance of the geomagnetic field.

A reading of maximum deflection has been made during each 3-hour interval (starting from 00 to 03 UT) of two components X and Y. The Sq + L variation has been compen-sated for. A detailed description of the determination of K-indices has been given by J. Bartels: The technique of scaling indices K and Q geomagnetic activity. IGY Annals 4, London 1957.

The following scale has been used:

K-indices	Deflection in	Deflection in nT				
0	0 - 1:	5				
1	15 - 30	0				
2	30 - 60	0				
3	60 - 120	0				
4	120 - 210	0				
5	210 - 360	0				
6	360 - 600	0				
7	600 - 990	0				
8	990 - 1500	0				
9	1500 and m	ore				

ALL-SKY CAMERA DATA

Information about available records of all-sky camera films has been presented on page B4. The photographic records are made on 16 mm colour film normally at a rate of one frame per min. Data is recorded at higher frame rates (up to six per min) during selected periods. The start and stop time of camera operation and an average over the hour of auroral intensity and weather conditions is included in the tables.

Legend:

- cCloudy Aurora a r Red aurora Α Strong aurora R Strong red aurora Aurora and partly cloudy ac Red aurora and partly cloudy rc Strong aurora and partly cloudy Ac Rc Strong red aurora and partly cloudy
- I No aurora, clear weather Ic No aurora, partly cloudy
- + Camera fault
- Camera off (camera not running due to daylight, snow or rain)

IONOSPHERIC ABSORPTION DATA

Ionospheric absorption at 30.0 MHz and 38.2 MHz is recorded by two La Jolla riometers. The riometer records of cosmic noise power are reproduced in this report. The zero power level is at bottom of each panel.

The absorption (in decibels) can be computed from the formula

where P is the received cosmic noise power, P_o the cosmic noise power expected under ionospherically quiet conditions at the same sidereal time.

GEOMAGNETIC K-INDICES, KIRUNA July 2000 Lower limit of K=9 is 1500 nT Time in UT

Day 1 2 3 4 5	00 - 0 4 1 3 4 2	03 - 06 2 2 1 2 2	- 0 2 2 2 2 1 3	1 1 2 2 5	12 - 1 1 3 3 5	15 - 2 2 2 2 3 4	18 - 2 2 2 2 2 2 3	21 - 24 1 3 5 2 2	K-sum 15 14 20 19 26
6 7 8 9 10	2 2 3 1 4	2 2 2 2 2 2	2 2 2 2 2 4	2 2 2 3 3	3 2 4 2 4	3 1 3 2 4	3 2 2 2 5	2 2 1 4 4	19 15 19 18 30
11 12 13 14 15	3 7 1 6 5	3 3 2 4 4	3 1 3 4 4	4 3 5 5 6	6 1 6 4 7	5 2 7 8 8	5 1 5 6 9	6 2 5 6 7	35 20 34 43 50
16 17 18 19 20	6 5 2 5 6	5 3 3 2 7	5 3 3 2 5	4 3 2 2 6	5 3 3 3 6	2 2 4 4 2	1 1 4 3 3	1 2 4 5 5	29 22 25 26 40
21 22 23 24 25	6 3 4 4 1	2 3 4 3 2	2 2 2 2 2 2	2 4 2 2 2	3 6 5 2 3	3 6 5 1 3	2 5 7 1 3	4 3 4 3 2	24 32 33 18 18
26 27 28 29 30	4 5 5 6 5	3 5 4 4 6	3 2 5 4 2	3 2 4 5 3	6 1 7 6 4	4 2 6 4 3	5 1 5 4 2	7 3 3 4 2	35 21 39 37 27
31	3	3	2	2	4	3	4	7	28

GEOMAGNETIC K-INDICES, KIRUNA August 2000 Lower limit of K=9 is 1500 nT Time in UT

Day 1 2 3 4 5	00 - 03 2 6 6 6 4	3 - 06 2 5 3 5 3	- 09 2 2 3 3 4	3 3 4 4	12 - 15 4 3 3 4 6	5 - 18 5 4 4 6 5	8 - 21 4 6 3 5 5	- 24 6 5 2 6 6	K-sum 28 34 27 39 37
6 7 8 9 10	5 2 3 4 2	6 2 2 2 3	3 2 2 2 4	3 3 2 2 4	5 4 3 2 4	5 3 3 2 4	4 4 2 2 2 6	5 3 4 3 7	36 23 21 19 34
11 12 13 14 15	7 5 6 1 3	6 8 6 2 3	6 7 4 2 2	5 6 3 2 2	5 7 3 4 2	5 7 2 4 2	6 8 2 4 3	5 4 3 6 2	45 52 29 25 19
16 17 18 19 20	3 2 2 2 1	2 2 1 1	2 2 2 2 2 2	1 1 2 2 2	3 2 1 1 2	2 2 2 2 2 2	3 2 0 1 1	4 2 0 1 2	20 15 10 12 13
21 22 23 24 25	2 3 1 2	3 1 1 2 1	2 1 1 3 1	4 2 2 3 2	3 1 3 2	3 1 3 2 2	5 1 2 1 2	6 0 2 1 1	28 10 15 16 11
26 27 28 29 30	2 2 3 7 6	2 1 2 5 5	2 1 2 5 2	2 2 3 4 3	2 3 5 4 3	1 1 5 5 3	1 3 5 6 5	1 3 6 5 5	13 16 31 41 32
31	5	3	2	4	4	3	4	7	32

GEOMAGNETIC K-INDICES, KIRUNA September 2000 Lower limit of K=9 is 1500 nT Time in UT

Day 1 2 3 4 5	00 - 0 5 6 1 2 4	03 - 06 3 5 2 2 3	- 09 4 4 2 2 2	3 5 2 2 2 2	12 - 4 5 2 3 2	15 - 18 3 5 2 4 4	8 - 21 7 1 5 7 4	1 - 24 6 3 5 6 4	K-sum 35 34 21 28 25
6	2	2	2	2	3	3	4	4	22
7	4	3	2	3	4	3	6	6	31
8	5	2	2	3	6	5	5	5	33
9	4	3	2	2	2	1	1	0	15
10	0	2	1	3	1	2	3	2	14
11	0	2	1	2	2	3	2	2	14
12	1	3	2	4	5	4	7	7	33
13	5	5	2	2	2	2	3	5	26
14	2	2	2	2	2	2	4	3	19
15	0	2	1	2	1	1	2	6	15
16	5	4	2	2	5	6	6	7	37
17	6	6	3	3	6	4	5	9	42
18	7	5	5	6	5	6	4	5	43
19	4	3	3	3	5	4	5	5	32
20	6	4	2	3	3	3	2	3	26
21 22 23 24 25	3 4 2 0 3	2 2 2 2 2 2	2 2 1 1 2	3 2 2 2 2 3	3 2 2 2 4	2 1 2 3 5	4 2 3 6 7	3 1 2 5 5	22 16 16 21 31
26	6	6	4	3	5	4	5	6	39
27	5	2	3	2	3	3	3	5	26
28	4	4	2	3	3	4	2	2	24
29	2	3	2	3	4	3	5	3	25
30	6	4	5	5	5	6	7	8	46

ALL-SKY CAMERA, KIRUNA September 2000

Start of recording season

13 - 14 - 15 - 16 - 17 - 18 - 19 - 20 - 21 - 22 - 23 - 24 - 01 - 02 - 03 - 04 - 05 - UT 1-2 2-3 3-4 4-5 5-6 ac c ccIc ac 6-7 a r a a a a 7-8 ccc c cc8-9 a A A Ac ac c 9-10 c c ac Ic Ic c10-11 a ac c cac 11-12 ac ac Ic I Ic Ic 12-13 Α ac ac ac ac ac 13-14 I I a ac ac ac 14-15 I a a a a +15-16 Ic Ic ac ac 16-17 ac ac ac ac ac ac 17-18 c c c c c c18-19 a Α a a a a 19-20 ac ac a a ac ac 20-21 cc a a a a 21-22 a a a a a a 22-23 c c ca a a 23-24 a ac cc cc 24-25 R a ac ac ac r 25-26 a a a Α a a 26-27 a a +a +27-28 I a a a a a 28-29 c c c c c c29-30 a a a a A A 30-1/10 R a Α A c ac

KIRUNA MAGNETOGRAM PLOTS

July - September 2000

http://www.irf.se/mag

KIRUNA VERTICAL RIOMETER

July - September 2000

(No records during July – August due to riometer fault)

http://www.irf.se/~rio