



KIRUNA GEOPHYSICAL DATA

DATA SUMMARY 00/4-6
APRIL-JUNE 2000

Collected at
SWEDISH INSTITUTE OF SPACE PHYSICS
KIRUNA, SWEDEN



Kiruna Geophysical Data

Data Summary 00/4-6
April - June 2000

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TABLE OF CONTENTS

Introduction.....page A1

Legend

Geomagnetic data.....A2
All-sky camera data.....A3
Ionospheric absorption data.....A3

Tables of April - June 2000

Geomagnetic K-indices.....B1
All-sky camera.....B4

Reproductions for April - June 2000

Magnetogram plots: <http://www.irf.se/mag>
Riometer recordings: <http://www.irf.se/~rio>

INTRODUCTION

"Kiruna Geophysical Data" is a collection of results of ground-based more or less continuous measurements of upper atmosphere variables carried out at the Swedish Institute of Space Physics. Our objective with this data summary is to present current geophysical data, related to polar upper atmospheric conditions, to those who require it for their re-search work.

Coordinates of the Swedish Institute of Space Physics are:

Geogr.		Geomagn. Dipole		Corr. Geomagn.	
Lat.	Long.	Lat.	Long.	Lat.	Long.
67.8°N	20.4°E	65.2°N	116.0°E	64.8°N	104.2°E

The following ground measurements of upper atmospheric parameters have been conducted at the Swedish Institute of Space Physics in the period for which data is presented in this report.

Geomagnetism	Optical aurora	Ionospheric ionization
1. Geomagnetic standard recording	1. All-sky camera	1. Ionosonde 0.5-16.0 MHz
2. Storm recording		2. Riometer 30.0 MHz
3. Geomagnetic digital recording		3. Riometer 38.2 MHz

Ionosonde data obtained at the Swedish Institute of Space Physics in Kiruna is published together with ionosonde data from Lycksele and Uppsala in a monthly report "Ionospheric Data Sweden" which is available from the Swedish Institute of Space Physics.

The Ionospheric Observatory at Lycksele (64.7°N, 18.8°E) is also operated by the Swedish Institute of Space Physics. Mainly magnetic, riometer, photometer, all-sky camera and ionosonde measurements are conducted there.

The digital magnetic recordings are reproduced in this report instead of the normal magnetograms from January 1980.

Universal time is used in all tables.

The most recent magnetometer and riometer data are now accessible via internet at <http://www.irf.se>. For more information contact the editor (victoria.barabash@irf.se).

GEOMAGNETIC DATA

Magnetograms

The magnetograms reproduced here are plots of data recorded by a three component digital fluxgate magnetometer with a sampling frequency of 1 sample/second.

The scaling is indicated by the tick mark values on the plots. Note that for large disturbances the normal scaling has been reduced. The positive direction of the components are: X north, Y east and Z down, in geographic coordinates.

K-indices

The K-indices give a measure of the amount of disturbance of the geomagnetic field.

A reading of maximum deflection has been made during each 3-hour interval (starting from 00 to 03 UT) of two components X and Y. The Sq + L variation has been compensated for. A detailed description of the determination of K-indices has been given by J. Bartels: The technique of scaling indices K and Q geomagnetic activity. IGY Annals 4, London 1957.

The following scale has been used:

K-indices	Deflection in nT
0	0 - 15
1	15 - 30
2	30 - 60
3	60 - 120
4	120 - 210
5	210 - 360
6	360 - 600
7	600 - 990
8	990 - 1500
9	1500 and more

ALL-SKY CAMERA DATA

Information about available records of all-sky camera films has been presented on page B4. The photographic records are made on 16 mm colour film normally at a rate of one frame per min. Data is recorded at higher frame rates (up to six per min) during selected periods. The start and stop time of camera operation and an average over the hour of auroral intensity and weather conditions is included in the tables.

Legend:

c	Cloudy
a	Aurora
r	Red aurora
A	Strong aurora
R	Strong red aurora
ac	Aurora and partly cloudy
rc	Red aurora and partly cloudy
Ac	Strong aurora and partly cloudy
Rc	Strong red aurora and partly cloudy
I	No aurora, clear weather
Ic	No aurora, partly cloudy
+	Camera fault
—	Camera off (camera not running due to daylight, snow or rain)

IONOSPHERIC ABSORPTION DATA

Ionospheric absorption at 30.0 MHz and 38.2 MHz is recorded by two La Jolla riometers. The riometer records of cosmic noise power are reproduced in this report. The zero power level is at bottom of each panel.

The absorption (in decibels) can be computed from the formula

where P is the received cosmic noise power, P_0 the cosmic noise power expected under ionospherically quiet conditions at the same sidereal time.

B1

GEOMAGNETIC K-INDICES, KIRUNA

April 2000

Lower limit of K=9 is 1500 nT

Time in UT

Day	00	- 03	- 06	- 09	- 12	- 15	- 18	- 21	- 24	K-sum
1	5	3	2	3	2	5	6	6	32	
2	6	5	3	4	4	2	3	5	32	
3	5	3	2	3	4	3	1	6	27	
4	6	5	3	2	5	4	7	5	37	
5	6	5	2	2	2	2	2	1	22	

6	3	5	2	2	4	6	6	7	35	
7	7	6	5	3	3	4	6	5	39	
8	4	2	3	4	2	3	5	5	28	
9	3	4	2	2	4	5	6	4	30	
10	6	5	3	3	4	6	5	6	38	

11	2	3	2	3	4	3	5	3	25	
12	3	2	2	2	2	1	1	3	16	
13	3	4	2	1	1	1	0	0	12	
14	0	1	1	2	2	2	1	0	9	
15	1	2	2	2	3	2	5	4	21	

16	5	7	4	4	5	4	4	3	36	
17	4	4	2	3	2	3	2	5	25	
18	4	2	2	1	1	2	4	3	19	
19	3	5	3	2	4	3	3	2	25	
20	4	4	3	2	3	4	3	3	26	

21	3	2	2	4	3	3	4	3	24	
22	1	1	2	2	2	2	3	1	14	
23	1	1	2	2	2	3	1	2	14	
24	1	3	3	4	6	3	2	4	26	
25	4	1	2	2	2	2	4	1	18	

26	0	1	2	2	2	1	0	0	8	
27	3	1	2	3	5	5	2	6	27	
28	6	3	2	3	4	3	5	4	30	
29	4	3	1	3	3	3	3	5	25	
30	5	5	2	1	3	3	5	4	28	

B2

GEOMAGNETIC K-INDICES, KIRUNA

May 2000

Lower limit of K=9 is 1500 nT

Time in UT

Day	00	-	03	-	06	-	09	-	12	-	15	-	18	-	21	-	24	K-sum
1		2		3		3		3		4		3		4		5		27
2		5		2		2		3		4		3		5		6		30
3		2		2		2		2		3		4		6		5		26
4		3		2		2		3		3		3		4		3		23
5		4		2		3		3		4		3		2		4		25

6		5		5		2		3		3		4		2		5		29
7		4		2		2		2		2		1		1		1		15
8		1		1		2		2		2		2		1		3		14
9		1		2		2		2		4		5		3		5		24
10		4		4		2		2		4		3		4		3		26

11		2		3		1		2		3		0		2		3		16
12		3		3		3		3		2		2		4		5		25
13		6		5		2		3		4		4		5		3		32
14		5		4		2		3		3		1		5		3		26
15		6		4		2		3		3		4		2		2		26

16		6		7		3		3		3		5		5		6		38
17		8		7		4		3		2		2		4		3		33
18		5		3		2		2		1		1		2		2		18
19		3		2		3		3		3		2		1		2		19
20		2		1		2		1		3		1		3		4		17

21		3		2		3		2		2		2		2		4		20
22		3		3		2		3		3		2		3		2		21
23		2		2		2		2		3		3		5		8		27
24		8		9		5		5		5		4		7		4		47
25		7		6		4		4		4		5		5		6		41

26		5		6		3		3		5		3		4		1		30
27		1		2		1		3		4		4		4		5		24
28		3		3		2		3		4		4		5		5		29
29		5		3		3		2		5		5		5		6		34
30		5		4		5		5		5		4		4		5		37

31		5		5		3		2		2		1		3		4		25

B3

GEOMAGNETIC K-INDICES, KIRUNA

June 2000

Lower limit of K=9 is 1500 nT

Time in UT

Day	00	- 03	- 06	- 09	- 12	- 15	- 18	- 21	- 24	K-sum
1	5	3	3	2	2	2	1	2	20	
2	2	2	1	2	2	1	2	3	15	
3	5	4	3	4	4	3	3	2	28	
4	1	2	2	2	2	3	4	6	22	
5	6	4	4	4	5	2	3	4	32	

6	5	3	3	3	4	4	1	1	24	
7	2	2	3	3	3	3	5	6	27	
8	4	5	3	5	7	7	5	5	41	
9	2	2	2	3	2	2	0	2	15	
10	5	5	5	5	5	5	4	4	38	

11	5	4	3	5	5	5	5	3	35	
12	3	3	3	2	5	4	4	3	27	
13	1	3	2	3	5	3	2	6	25	
14	(6)	2	3	3	4	4	5	4	(31)	
15	4	3	3	3	4	4	6	4	31	

16	5	2	2	1	2	2	1	3	18	
17	3	0	2	2	3	3	1	2	16	
18	1	1	2	3	4	4	5	5	25	
19	4	4	2	2	3	2	2	2	21	
20	4	2	3	3	4	2	1	1	20	

21	1	2	2	2	2	2	1	5	17	
22	6	3	3	3	3	3	3	2	26	
23	3	3	3	3	6	4	6	6	34	
24	6	4	4	3	2	2	2	4	27	
25	2	2	2	2	2	2	2	3	17	

26	4	5	4	5	5	6	7	5	41	
27	6	4	3	3	4	4	5	6	35	
28	7	6	3	2	2	3	4	5	32	
29	6	3	2	3	2	2	1	3	22	
30	2	1	2	1	1	2	3	3	15	

ALL-SKY CAMERA, KIRUNA
April 2000

	13	14	15	16	17	18	19	20	21	22	23	24	01	02	03	04	05	UT
1-2	-	-	-	-	-	-	-	-	A	Ac	ac	Rc	Ac	Rc	-	-	-	-
2-3	-	-	-	-	-	-	-	-	a	A	a	a	a	a	-	-	-	-
3-4	-	-	-	-	-	-	-	-	c	c	c	c	+	+	-	-	-	-
4-5	-	-	-	-	-	-	-	-	c	c	c	c	Ac	c	-	-	-	-
5-6	-	-	-	-	-	-	-	-	c	c	c	c	c	c	-	-	-	-
6-7	-	-	-	-	-	-	-	-	ac	Rc	Rc	R	R	r	-	-	-	-
7-8	-	-	-	-	-	-	-	-	A	a	a	a	a	I	-	-	-	-
8-9	-	-	-	-	-	-	-	-	Ic	Ic	c	Ic	ac	Ic	-	-	-	-
9-10	-	-	-	-	-	-	-	-	ac	a	a	a	a	a	-	-	-	-
10-11	-	-	-	-	-	-	-	-	A	A	ac	a	ac	c	-	-	-	-
11-12	-	-	-	-	-	-	-	-	Ic	Ic	c	c	c	c	-	-	-	-
12-13	-	-	-	-	-	-	-	-	Ic	c	c	c	c	c	-	-	-	-
13-14	-	-	-	-	-	-	-	-	c	-	c	-	-	-	-	-	-	-
14-15	-	-	-	-	-	-	-	-	c	c	c	c	c	c	-	-	-	-
15-16	-	-	-	-	-	-	-	-	c	c	c	c	c	c	-	-	-	-
16-17	-	-	-	-	-	-	-	-	c	-	-	-	-	-	-	-	-	-

No records after this date due to sunlight

KIRUNA MAGNETOGRAM PLOTS

April - June 2000

<http://www.irf.se/mag>

KIRUNA VERTICAL RIOMETER

April - June 2000

<http://www.irf.se/~rio>

(No records during 000417 – 000630 due to riometer fault)