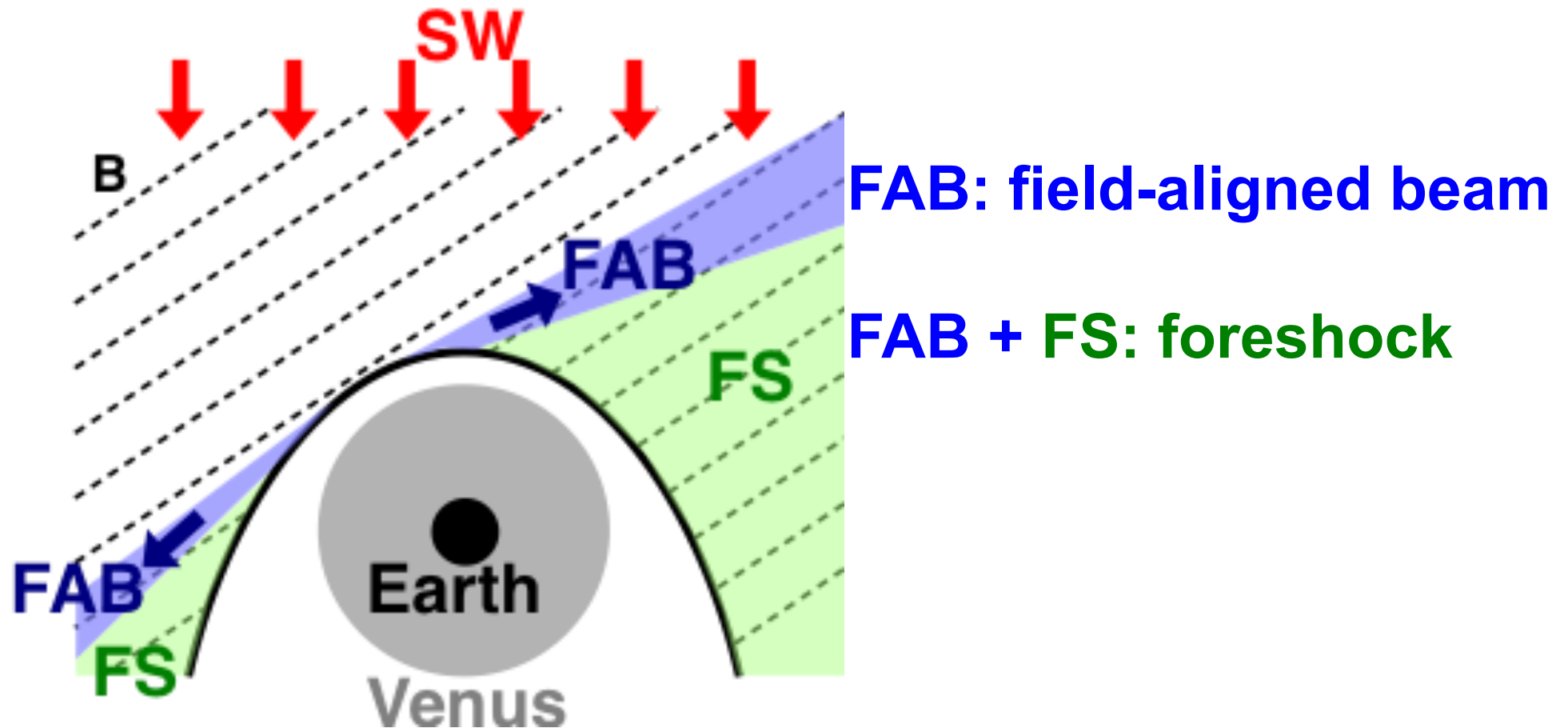


Foreshock studies by MEX and VEX

M. Yamauchi et al.



Shock = Fluid nature

Foreshock = Particle nature

Important parameters

(1) Alfvén Mach number (M_A)

$x \Rightarrow$ (2) Gyroradius (r_g) / Bow-shock radius (R_S)

(3) Inertia length (c/ω_{pi}) / Bow-shock size (R_S)

SW parameter	R_S (BS radius)	M_A ($\propto n^{1/2}V/B$)	c/ω_{pi} ($\propto n^{-1/2}$) & $c/\omega_{pi}R_S$	r_g ($\propto V/B$) & r_g/R_S
Venus	1	1	1 & 1	1 & 1
Earth	~ 5	~ 1.2	~ 1.7 & ~ 0.3	~ 2 & ~ 0.4
Mars	~ 0.5	~ 1.4	~ 3 & ~ 5	~ 4 & ~ 8

For Mars: $R_S \sim 5000$ km for Martian Subsolar

2 keV H^+ under 6 nT $\Rightarrow r_g = 1000$ km

5/cm³ H^+ $\Rightarrow c/\omega_{pi} = 100$ km

Venus-Mars difference: (2) cold H⁺

(1) Gravity: **Venus > Mars**



(2) Exosphere: **Venus < Mars**



(3) newly born H⁺: **Venus << Mars**

(This is clear from the difference in “ring distribution”)



(4) cold H⁺ at Bow shock: **Venus << Mars**

(High density cold H⁺ is observed only for Mars)

Venus - Mars difference (summary)

- (1) Alfvén Mach number (M_A)
- (2) Gyroradius (r_g) / Bow-shock radius (R_S)
- (3) Inertia length (c/ω_{pi}) / Bow-shock size (R_S)
- (4) Cold ion inside Bow-shock

parameter	R_S	M_A	$c/\omega_{pi}R_S$	r_g/R_S	cold H+ at BS
Venus	1	1	1	1	very little
Mars	~ 0.5	~ 1.4	~ 5	~ 8	a lot

↑
?

↑
?

Ending (add Earth)

- (1) Alfvén Mach number (M_A)
- (2) Gyroradius (r_g) / Bow-shock radius (R_S)
- (3) Inertia length (c/ω_{pi}) / Bow-shock size (R_S)
- (4) Cold ion inside Bow-shock

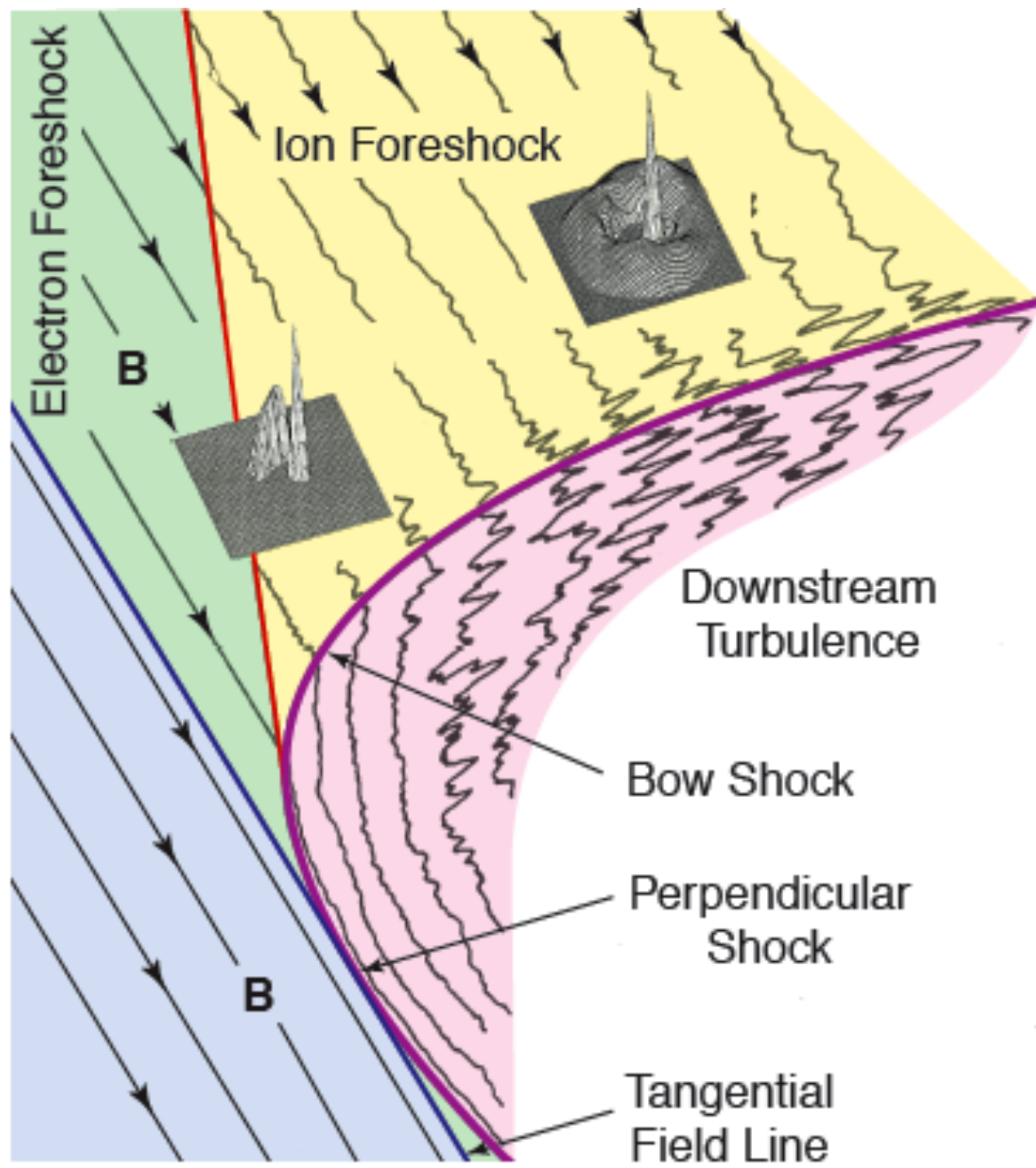
parameter	R_S	M_A	$c/\omega_{pi}R_S$	r_g/R_S	cold H+ at BS
Earth	5	~ 1.2	~ 0.3	~ 0.4	no
Venus	1	1	1	1	very little
Mars	~ 0.5	~ 1.4	~ 5	~ 8	a lot

↑
?

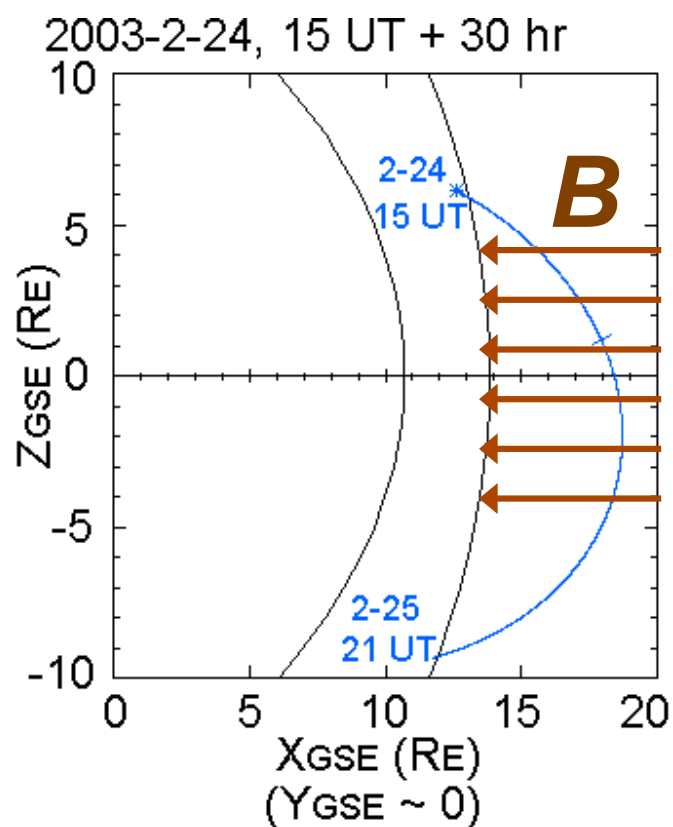
↑
?

Outline

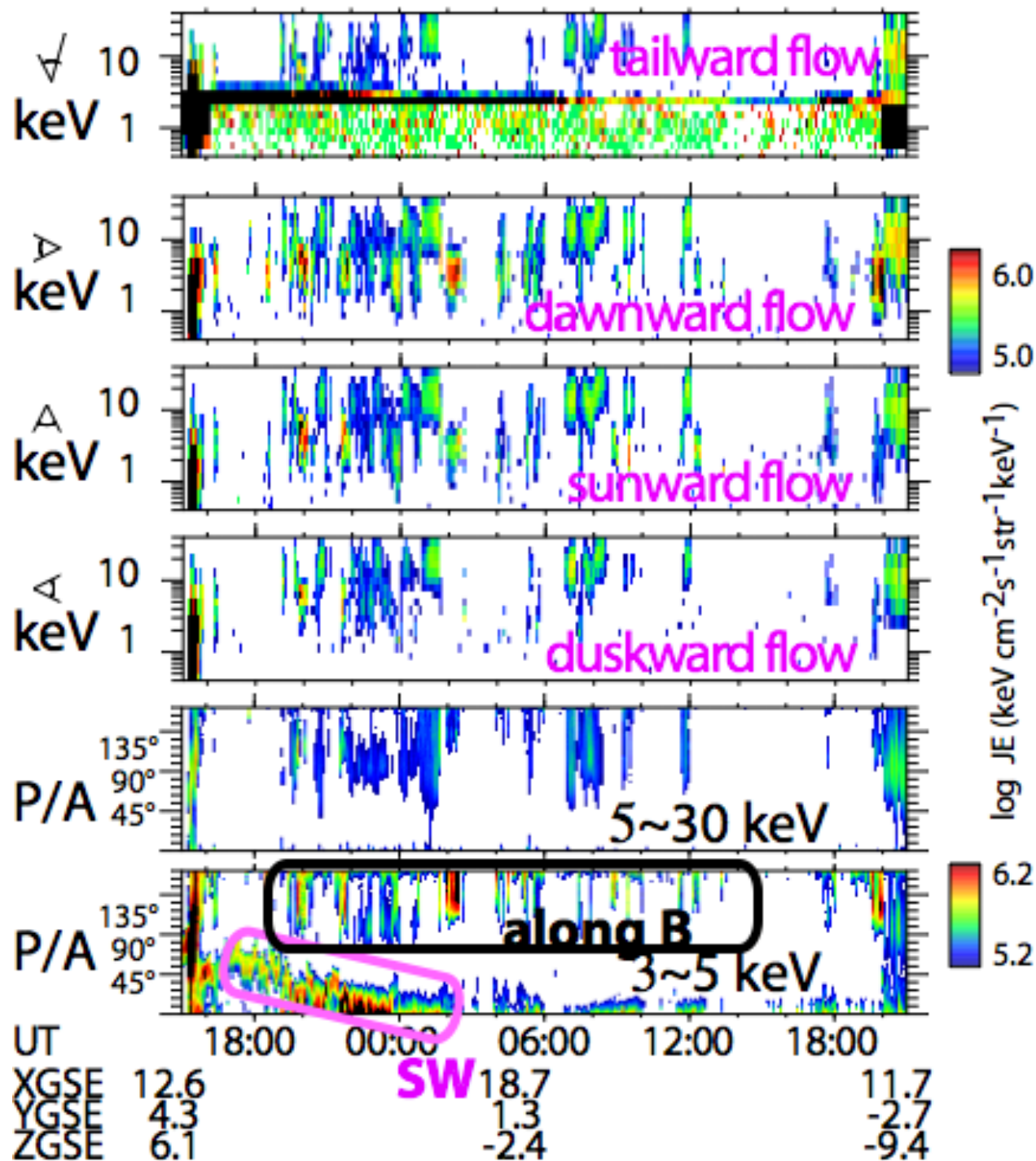
- 1. Introduction:** ion motion
- 2. Introduction:** Earth's knowledge
- 3. Venus** (similar to Earth)
- 4. Mars** (Different from Venus/Earth)



Earth's case
= since 1970's



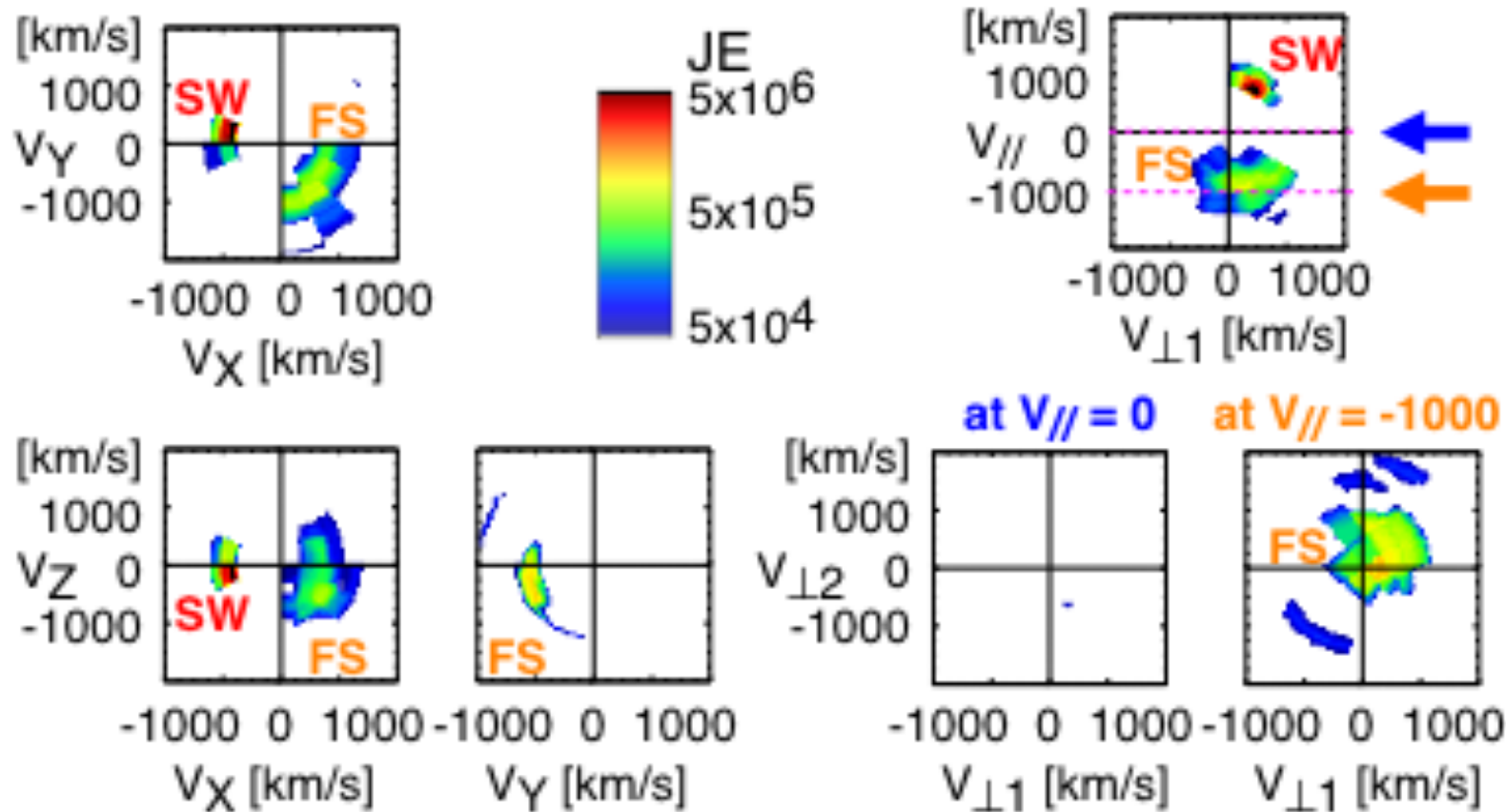
Cluster CIS (SC-1) H⁺ data , 24~25 Feb 2003



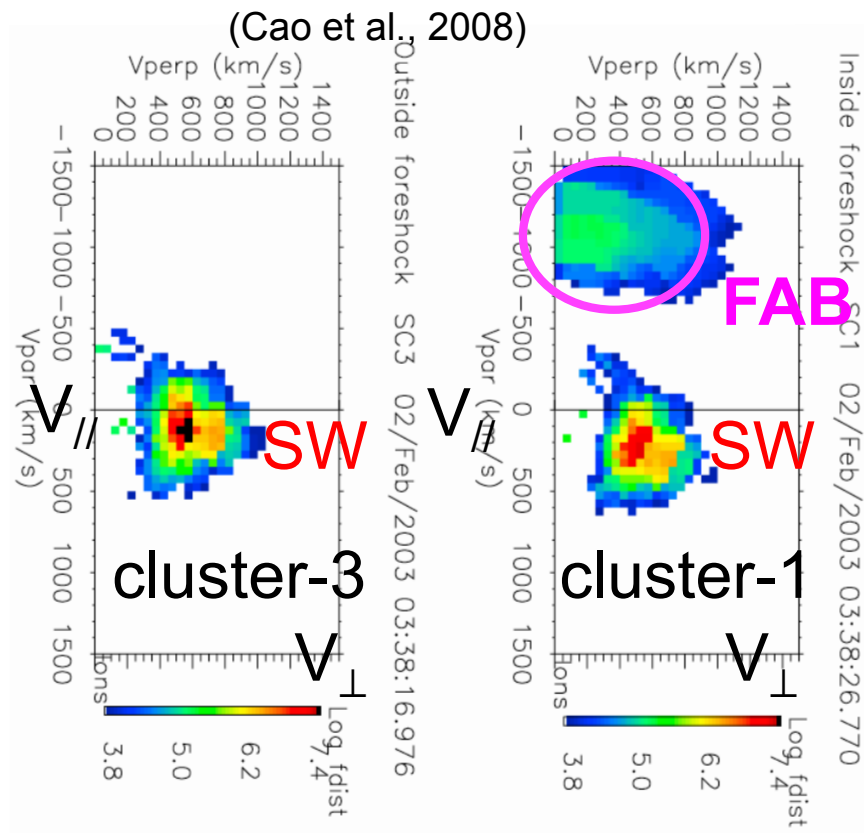
- * Upstream region
- * Large $V_{//}$ (& sunward)
- * Energized ($> E_{sw}$)

SC-1 CIS/HIA

2007-1-18 , 08:23:51



* Localized (< few 100km)



Ion motion in B (and E)

Lorenz transform

$$B' = B$$

$$\mathbf{E}' = \mathbf{E} + \mathbf{V} \times \mathbf{B}$$

Lorenz force on ion

$$\mathbf{F} = q(\mathbf{E}' + q\mathbf{v} \times \mathbf{B}')$$

where

V: velocity of frame

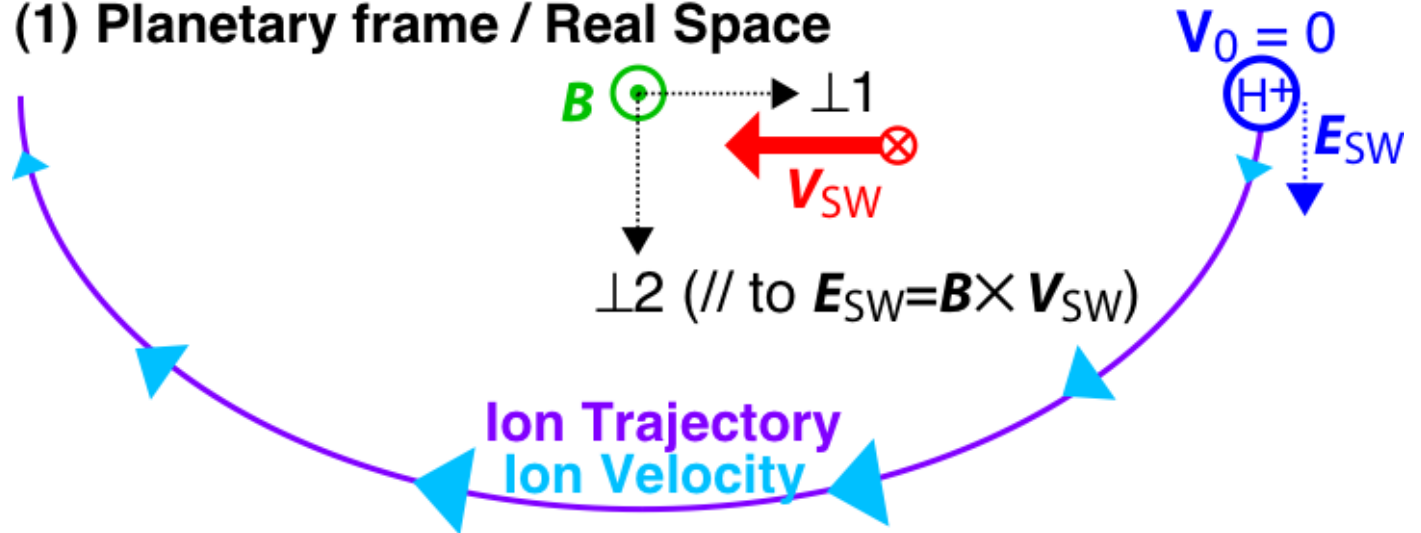
v: velocity of ion

q: charge of ion

and

$$R_G = mv_{\perp}/qB$$

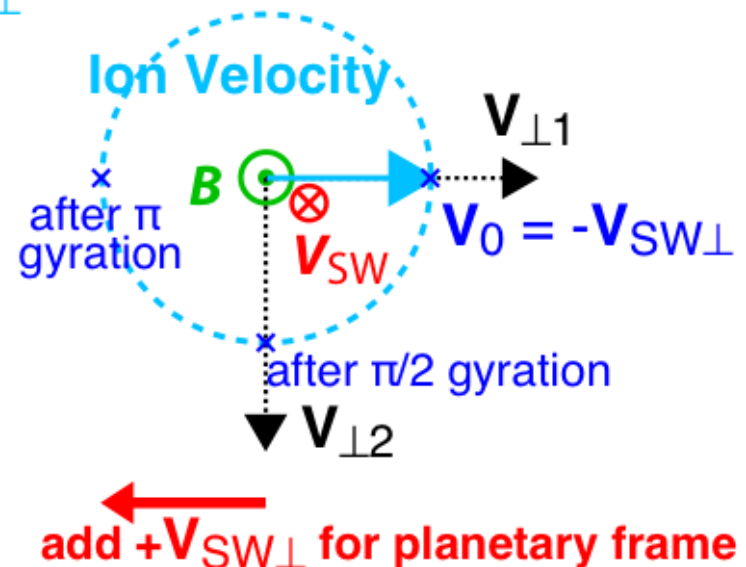
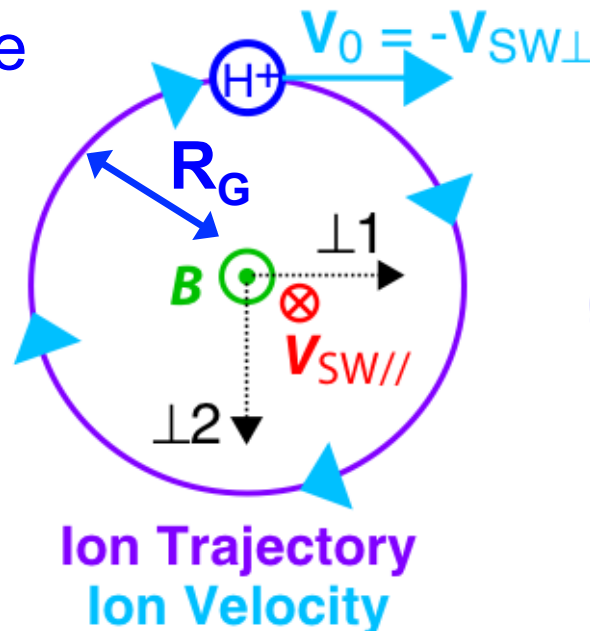
(1) Planetary frame / Real Space

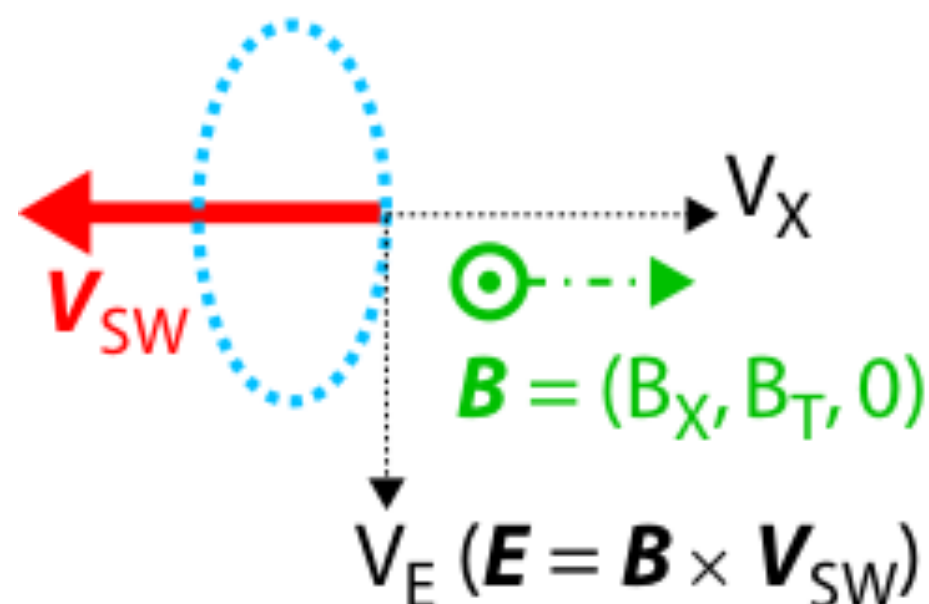
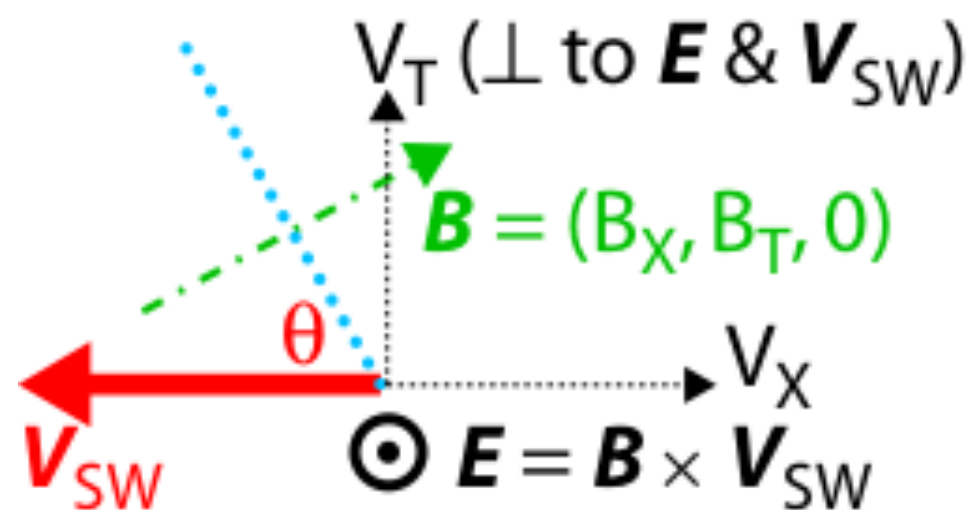


(2) Solar Wind (shifted by $-\mathbf{V}_{\text{SW}\perp}$) frame

(2a) Real Space

(2b) Velocity Space



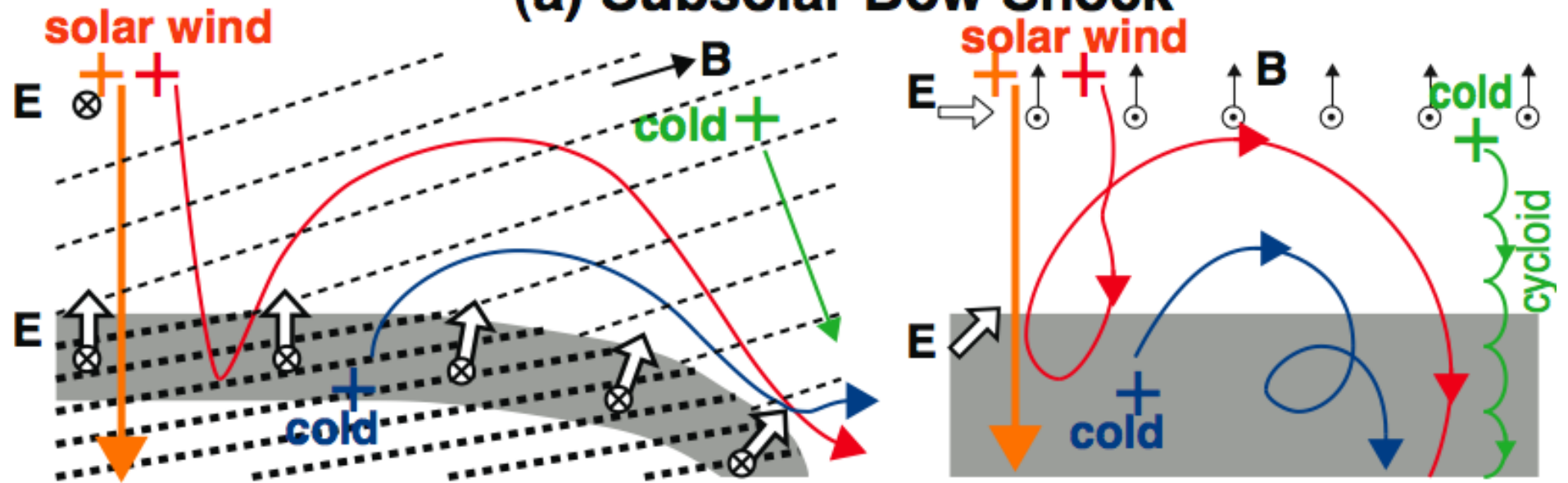


Long Radius = $|\mathbf{V}_{SW} \times \mathbf{b}|$

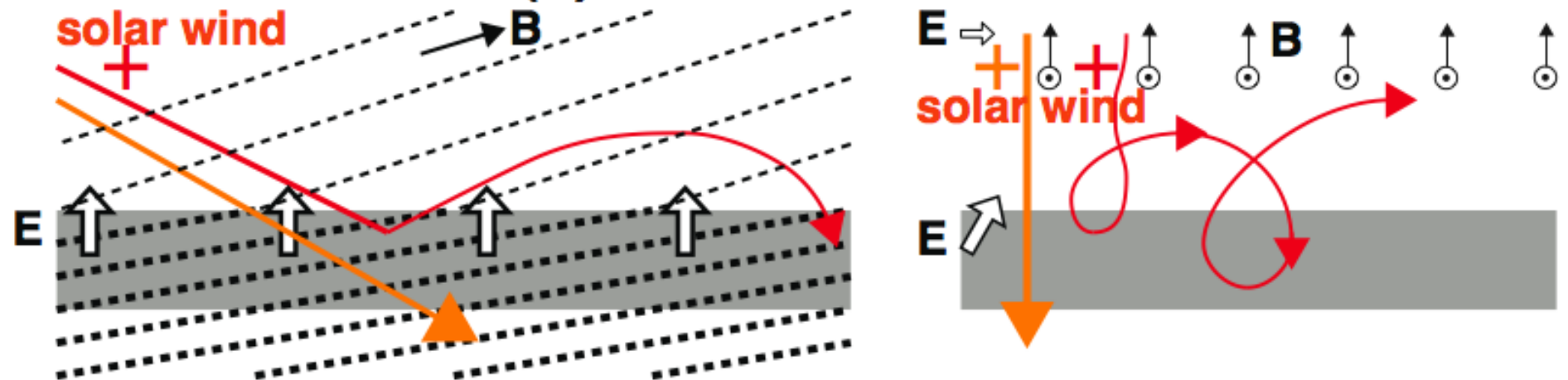
Short radius = $|\mathbf{V}_{SW} \times \mathbf{b}| \cos(\theta)$

1. solar wind, 2. newly born ion, 3. bow-shock cold ion

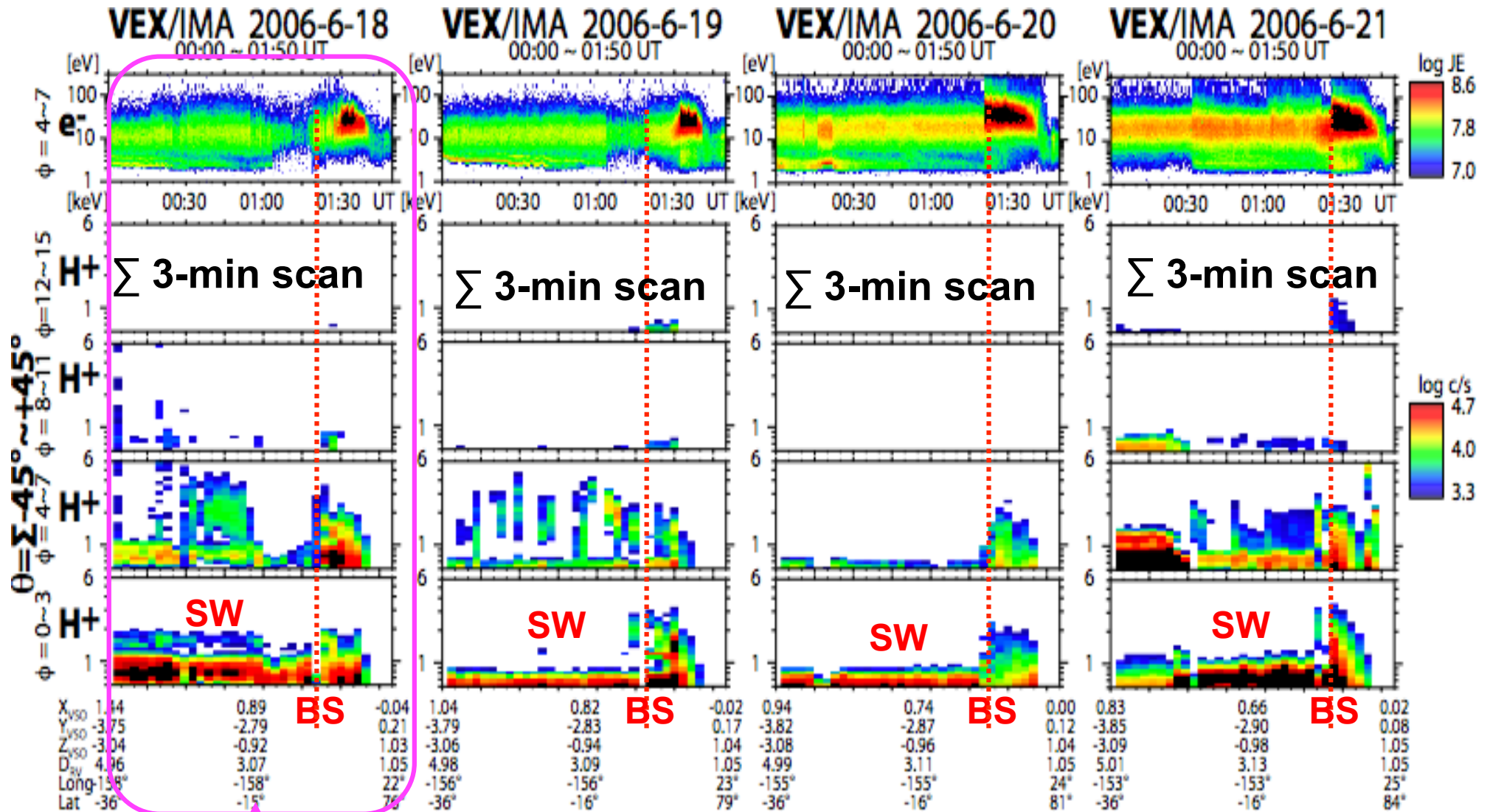
(a) Subsolar Bow Shock



(b) Flank Bow Shock



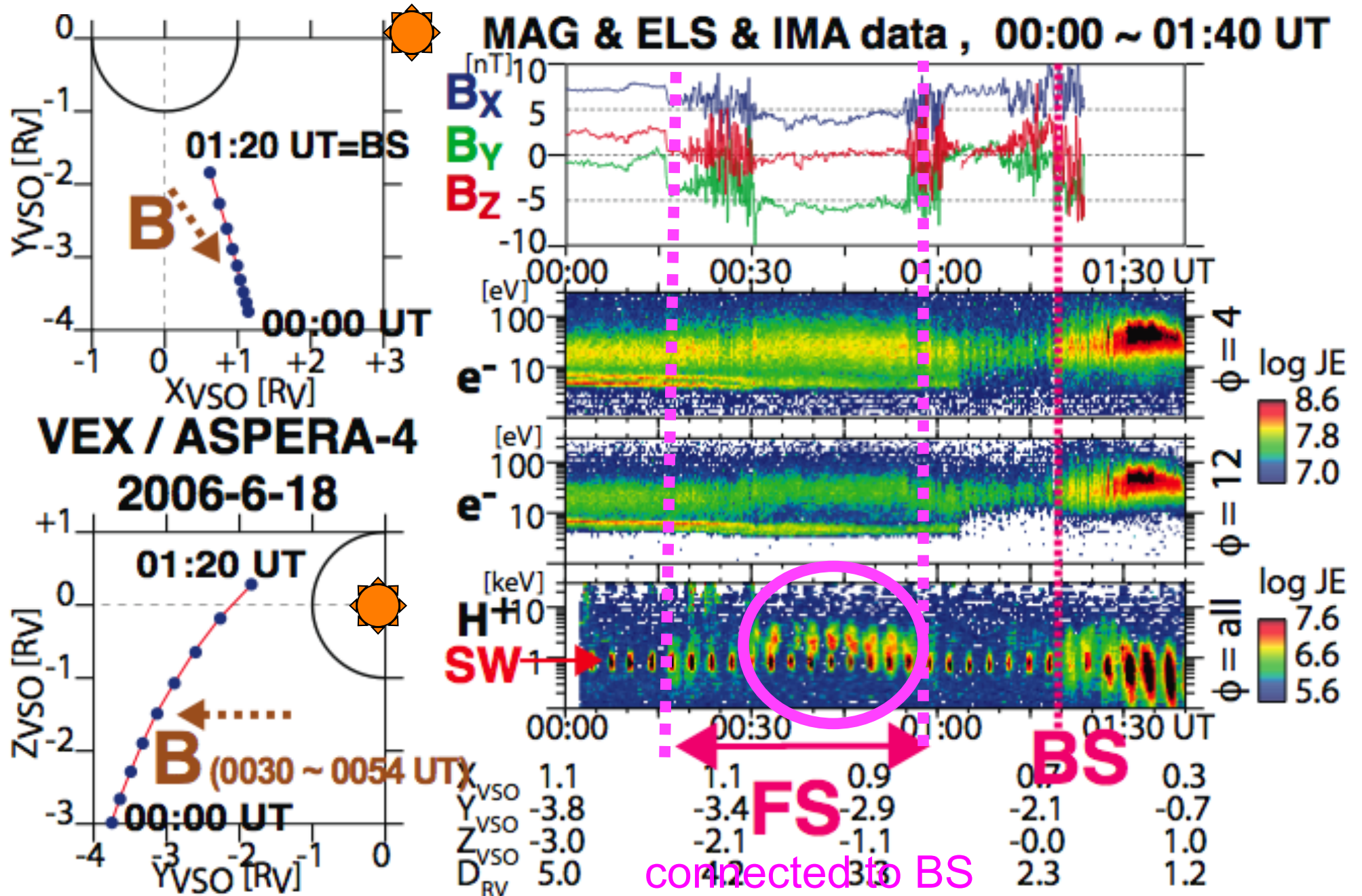
Venus (Venus Express)



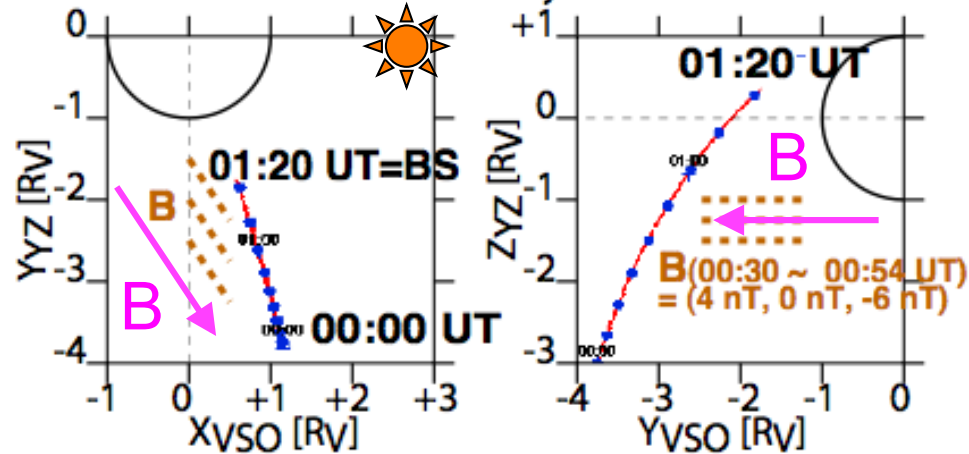
e- (top) & H+ (rest) at different angle

we show this

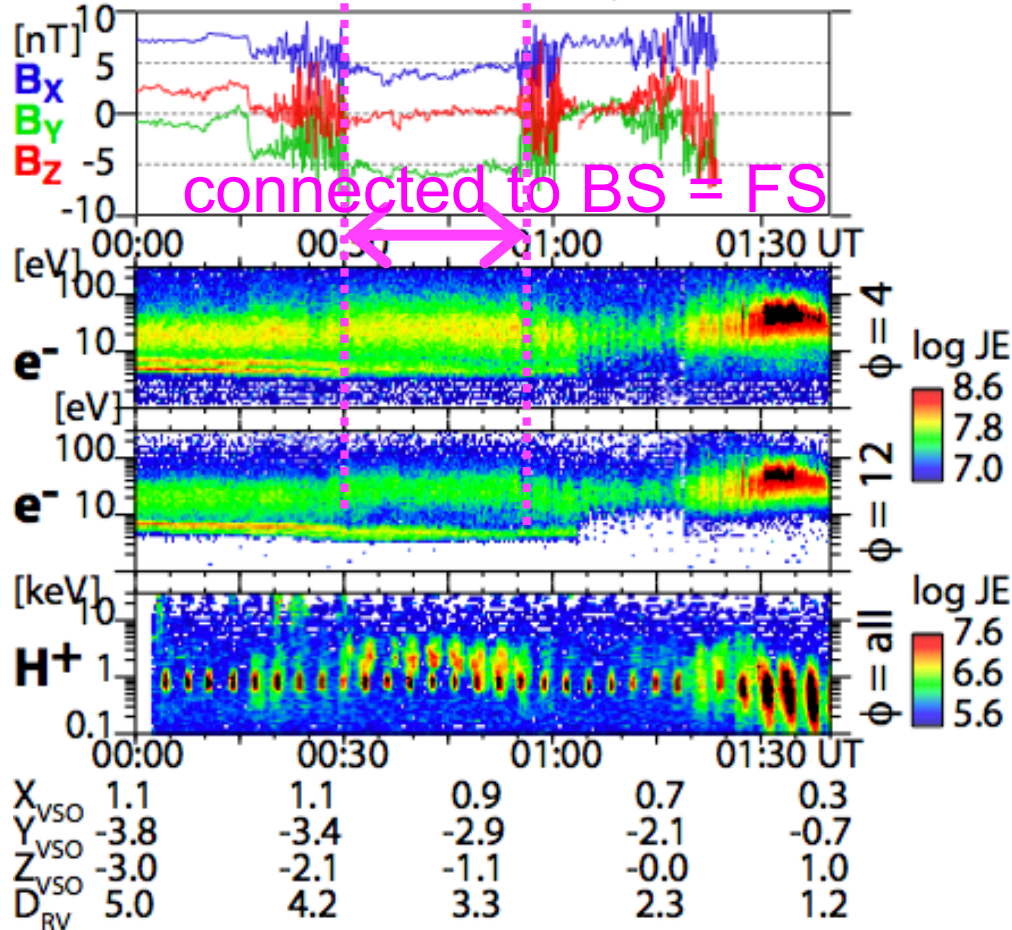
Venus (VEX)



VEX/ASPERA-4, 2006-6-18

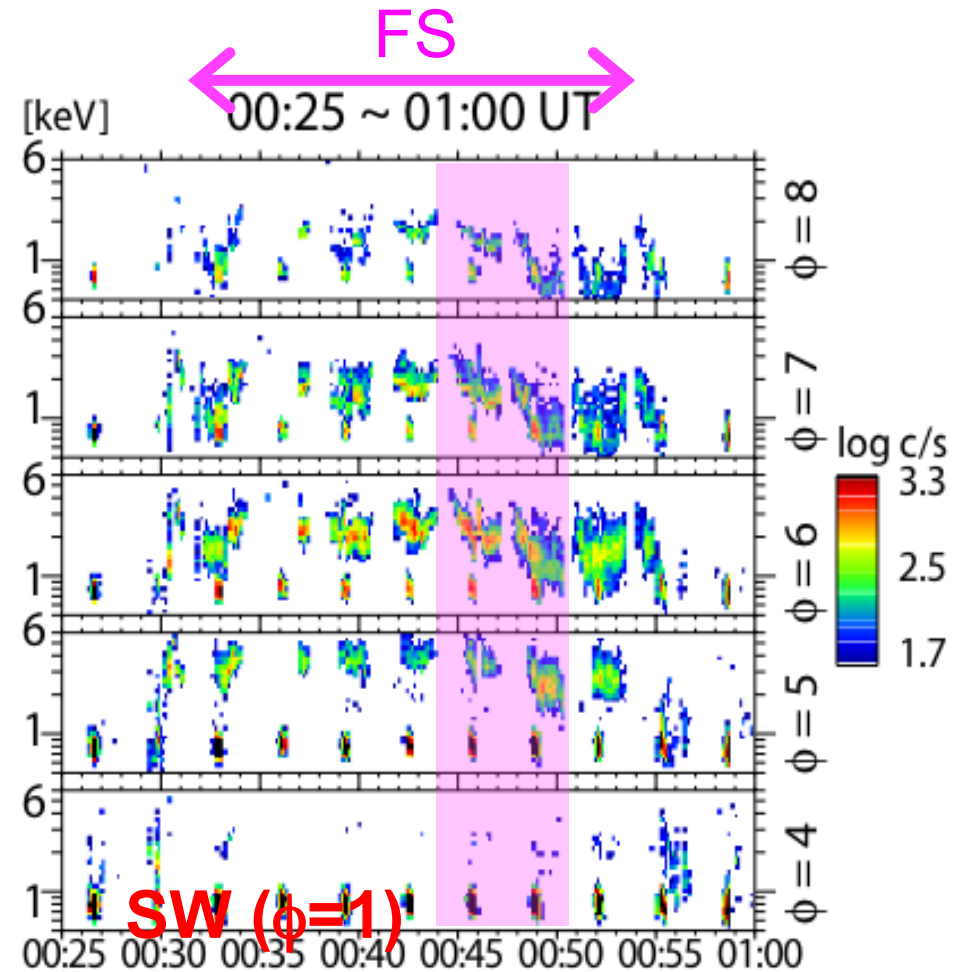


MAG & ELS & IMA data, 00:00 ~ 01:40 UT



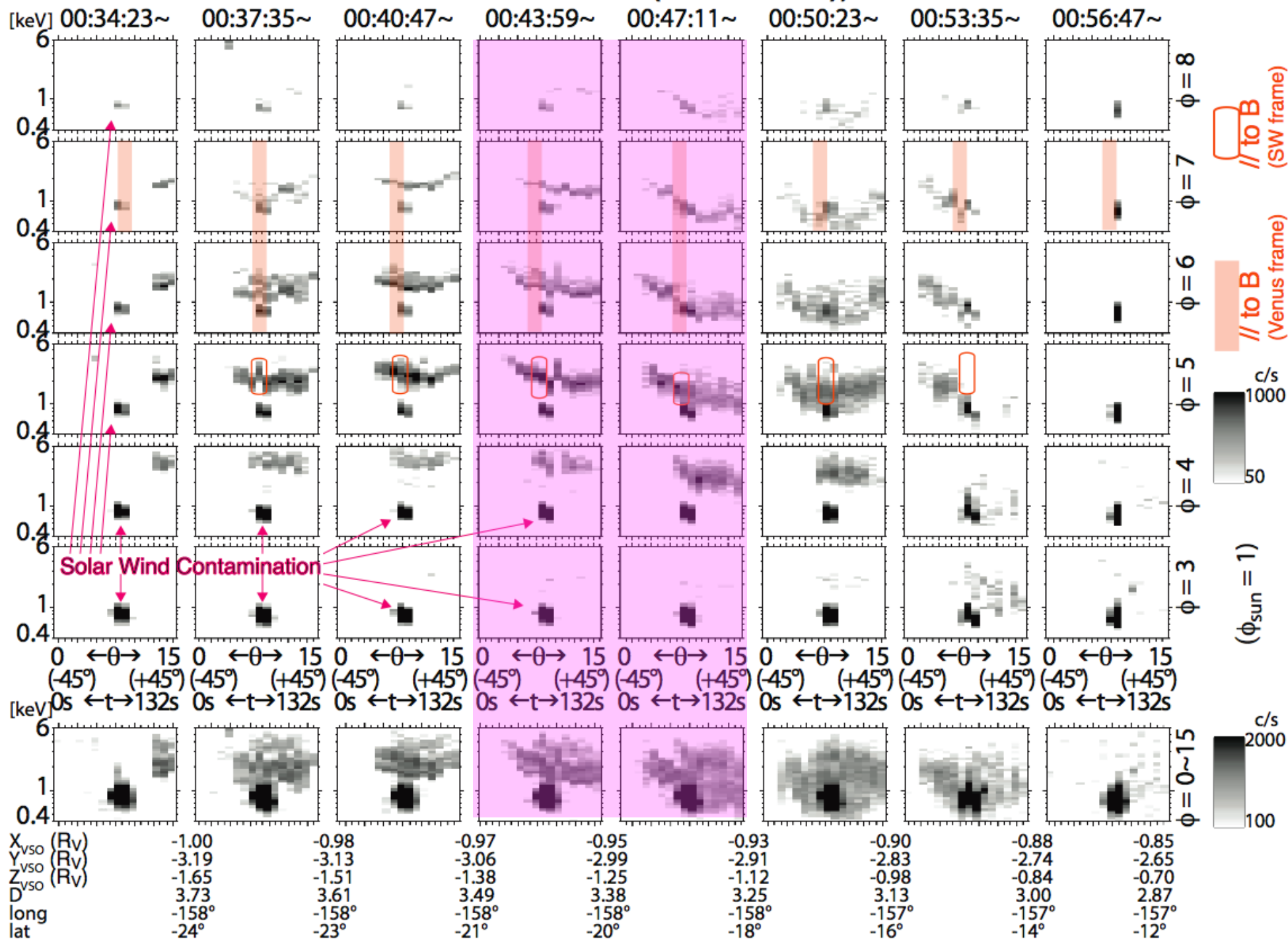
Venus $\approx$ Earth

2006-6-18

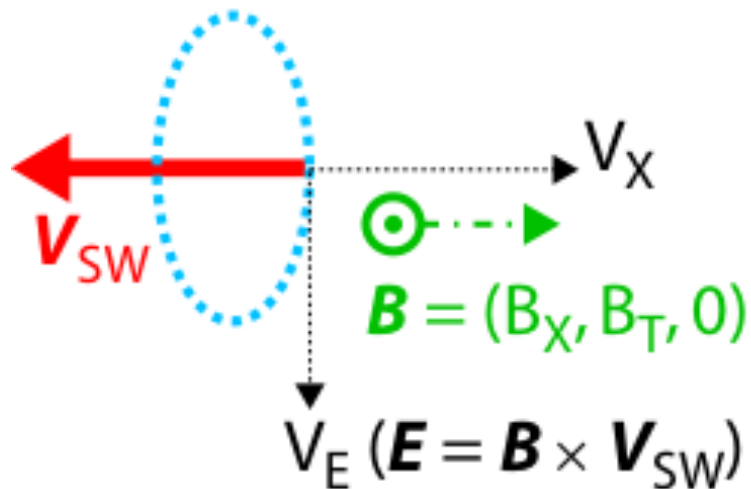
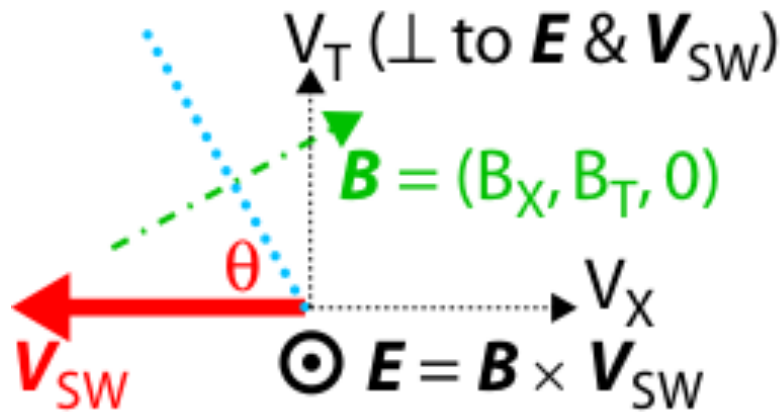


Scanning over $-45^\circ \sim +45^\circ$

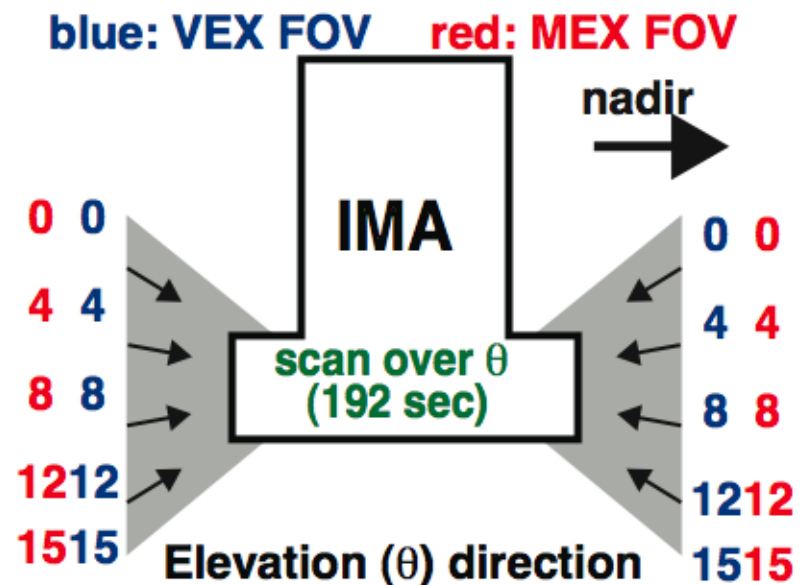
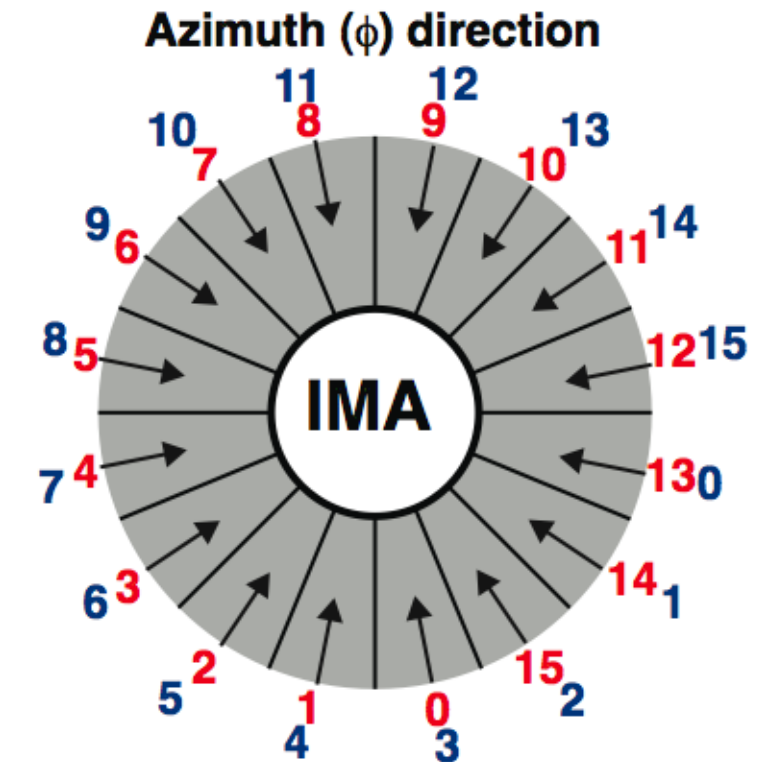
VEX/IMA H⁺ data (0.4~6 keV), 2006-6-18

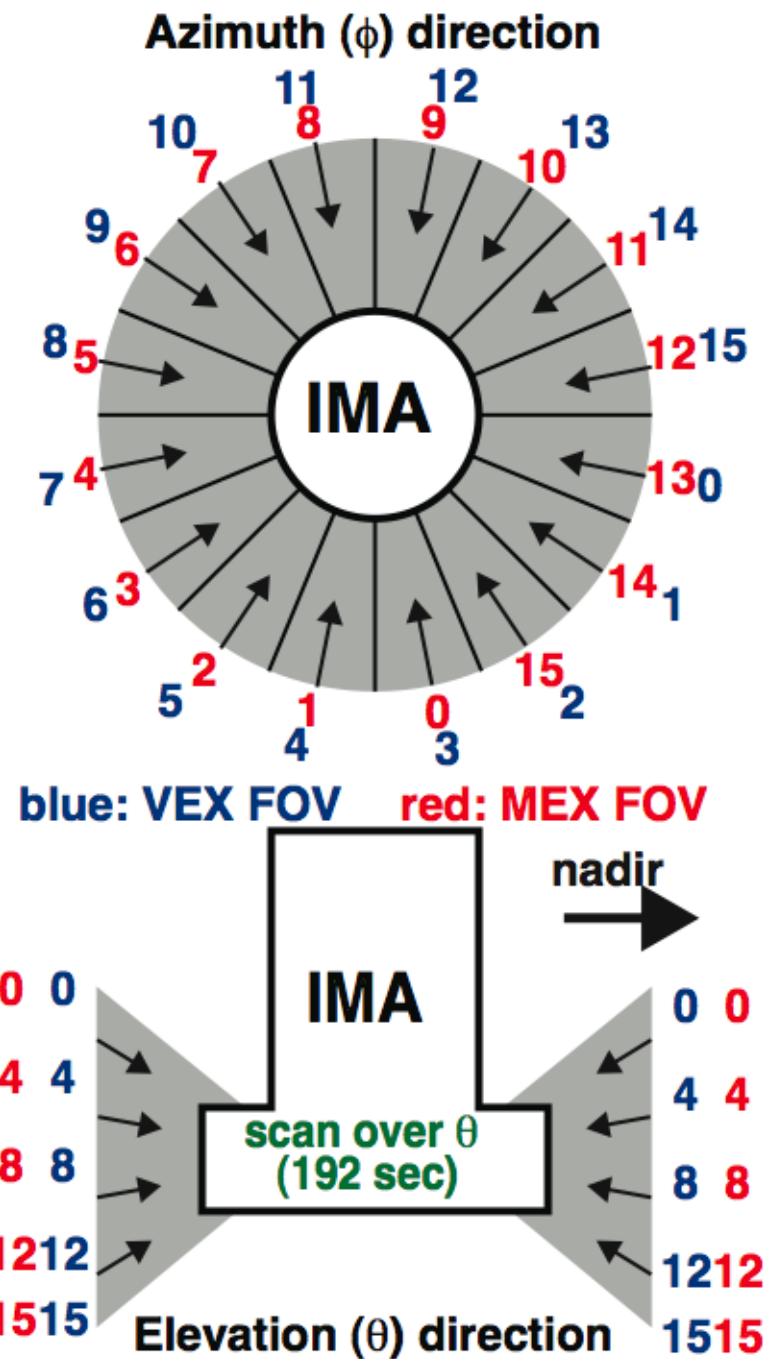
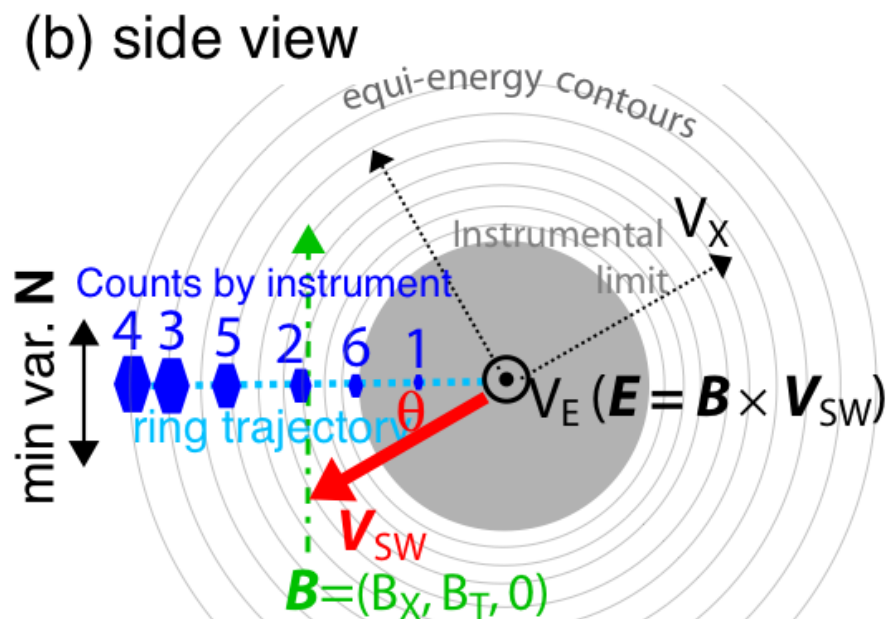
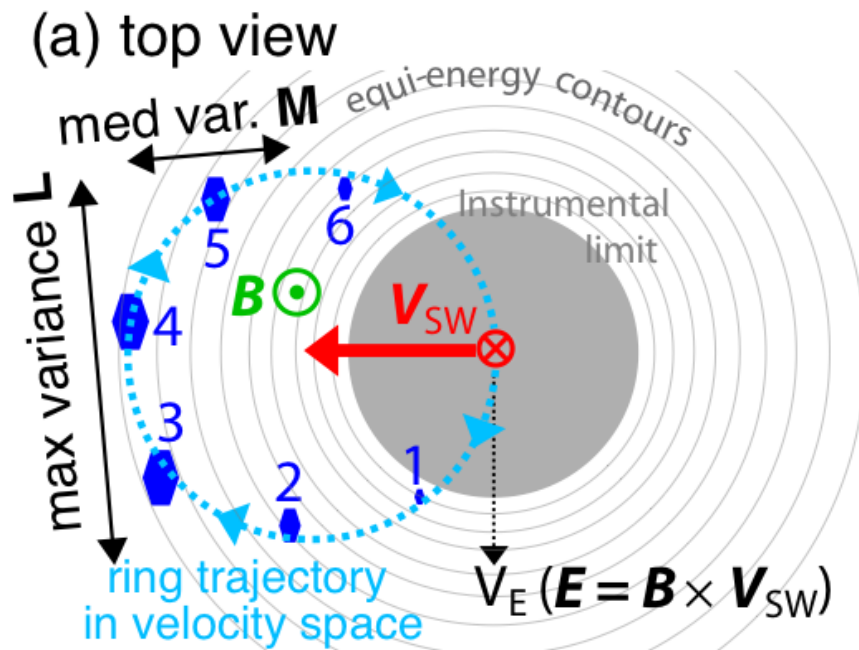


gyrotropic ion @ IMA Instrument



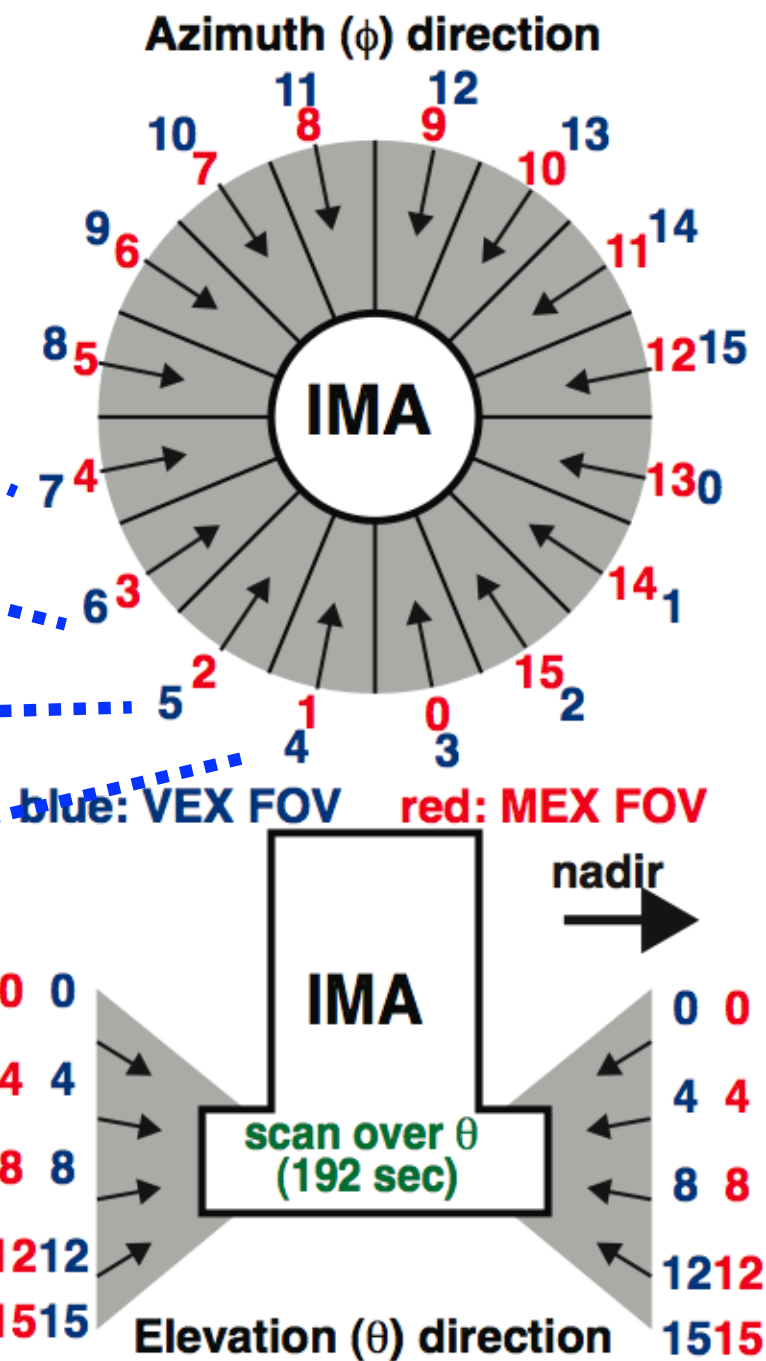
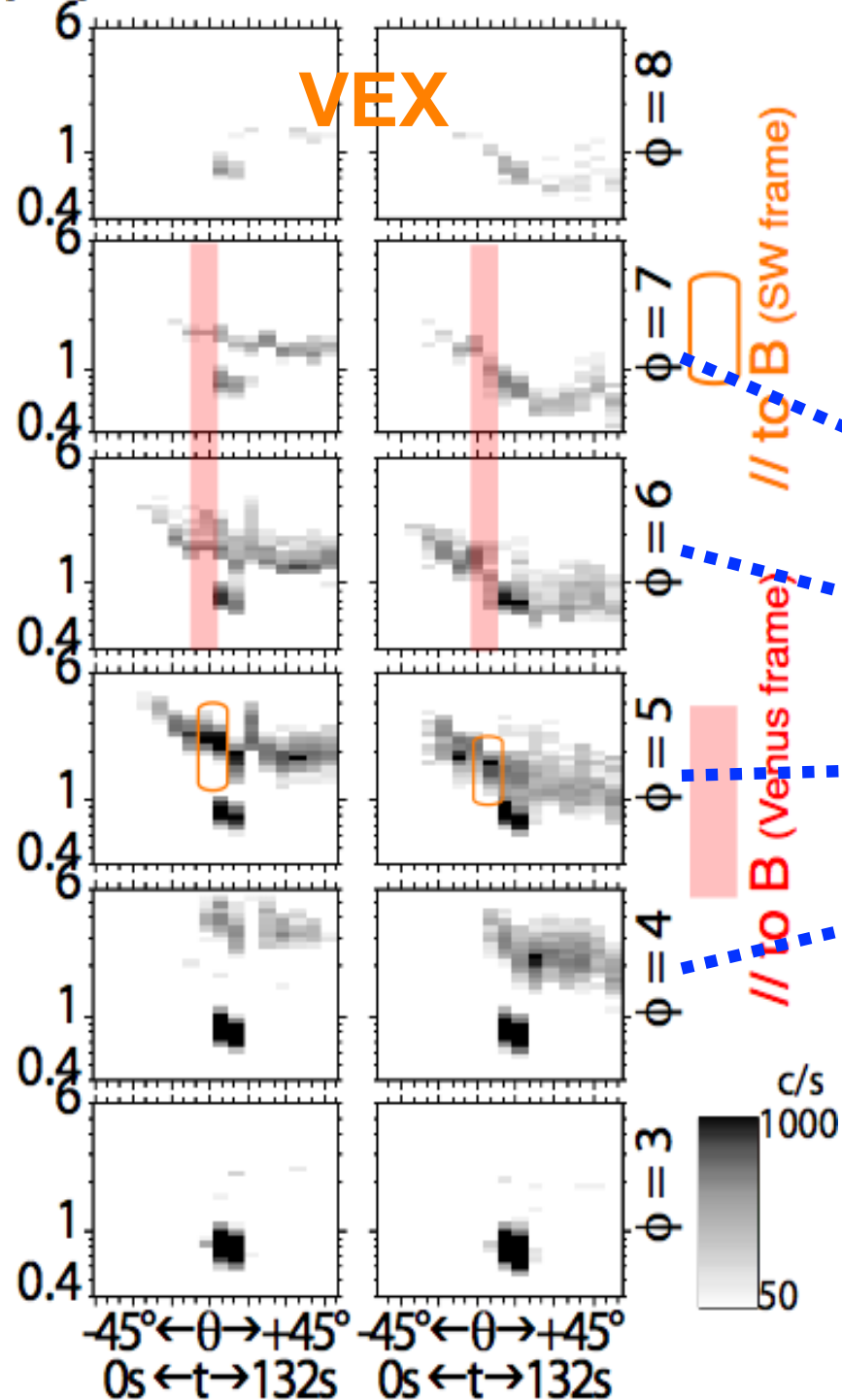
Long Radius = $|\mathbf{V}_{SW} \times \mathbf{b}|$
 Short radius = $|\mathbf{V}_{SW} \times \mathbf{b}| \cos(\theta)$

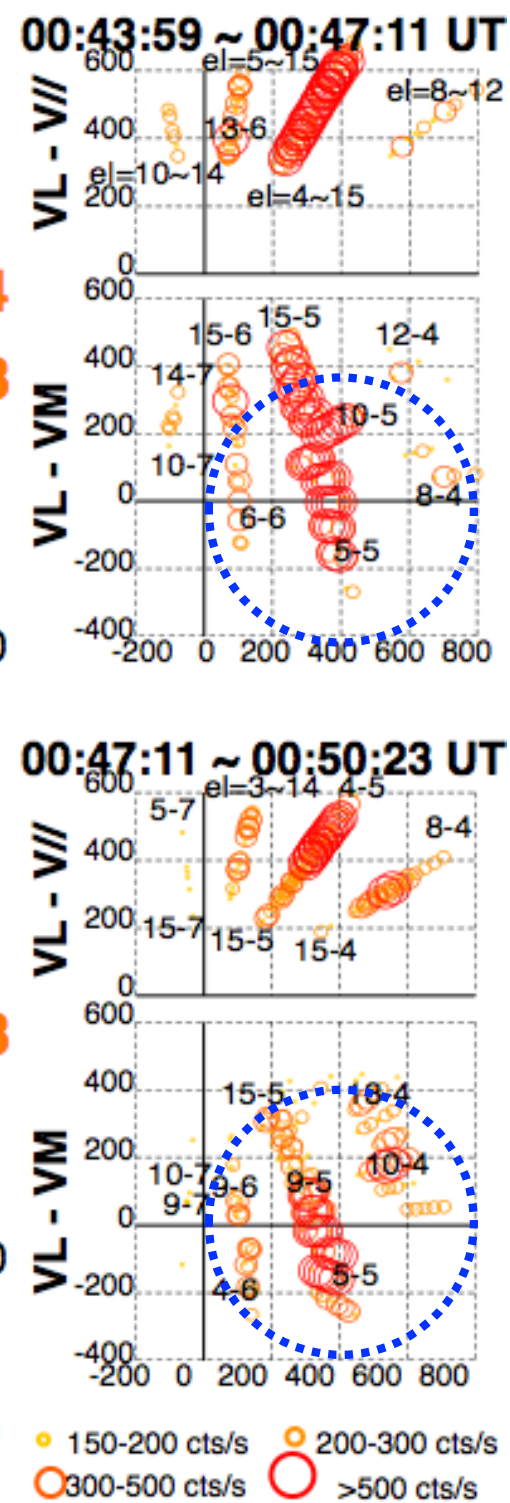
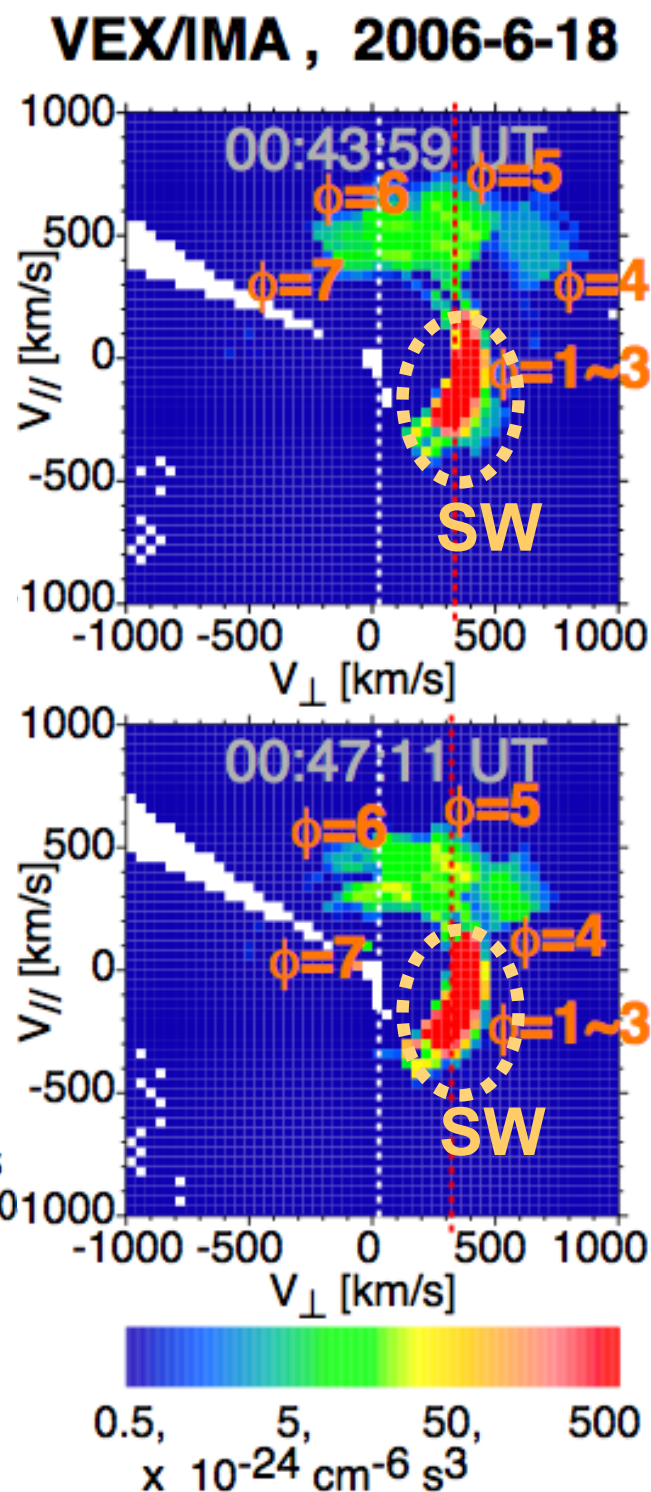
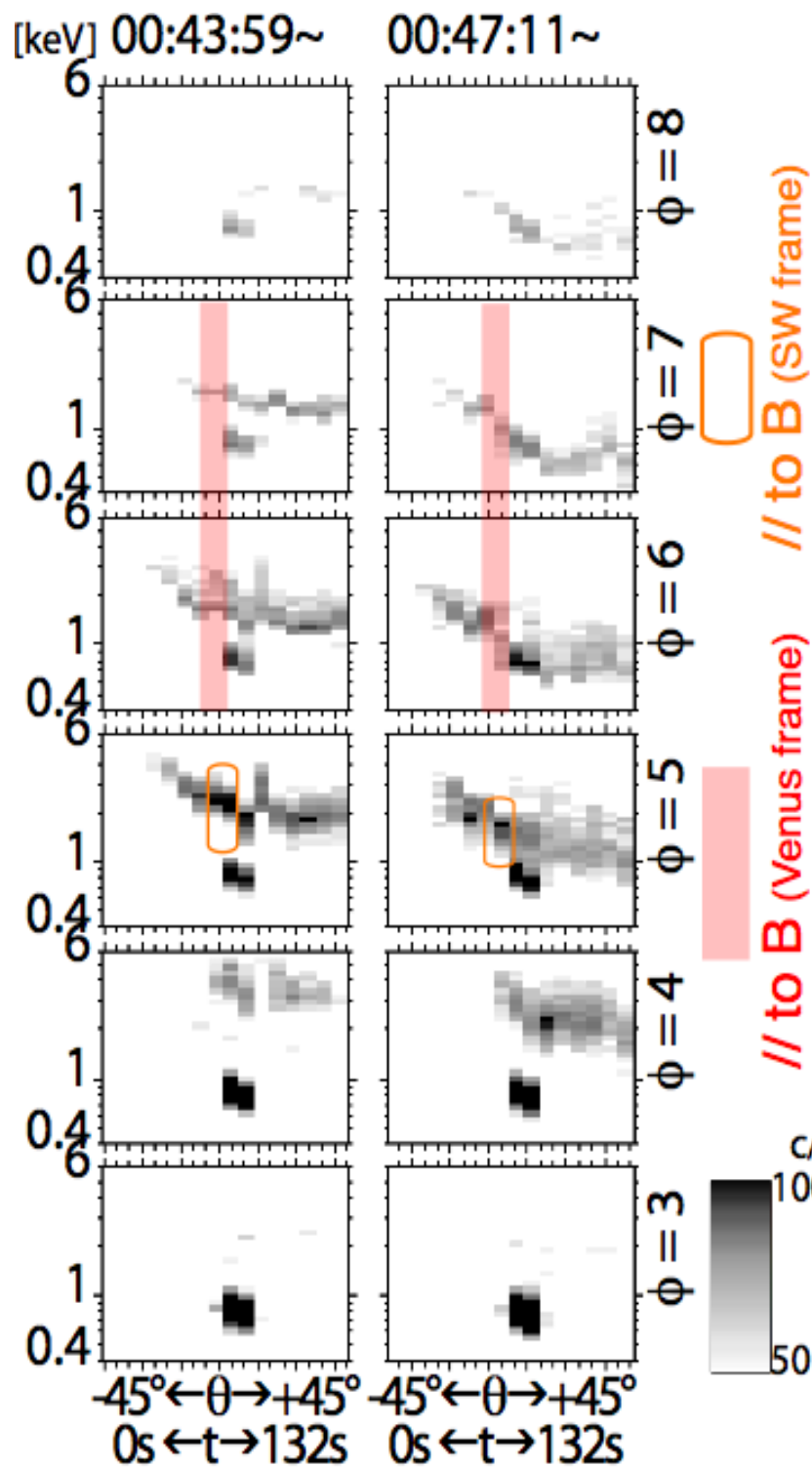




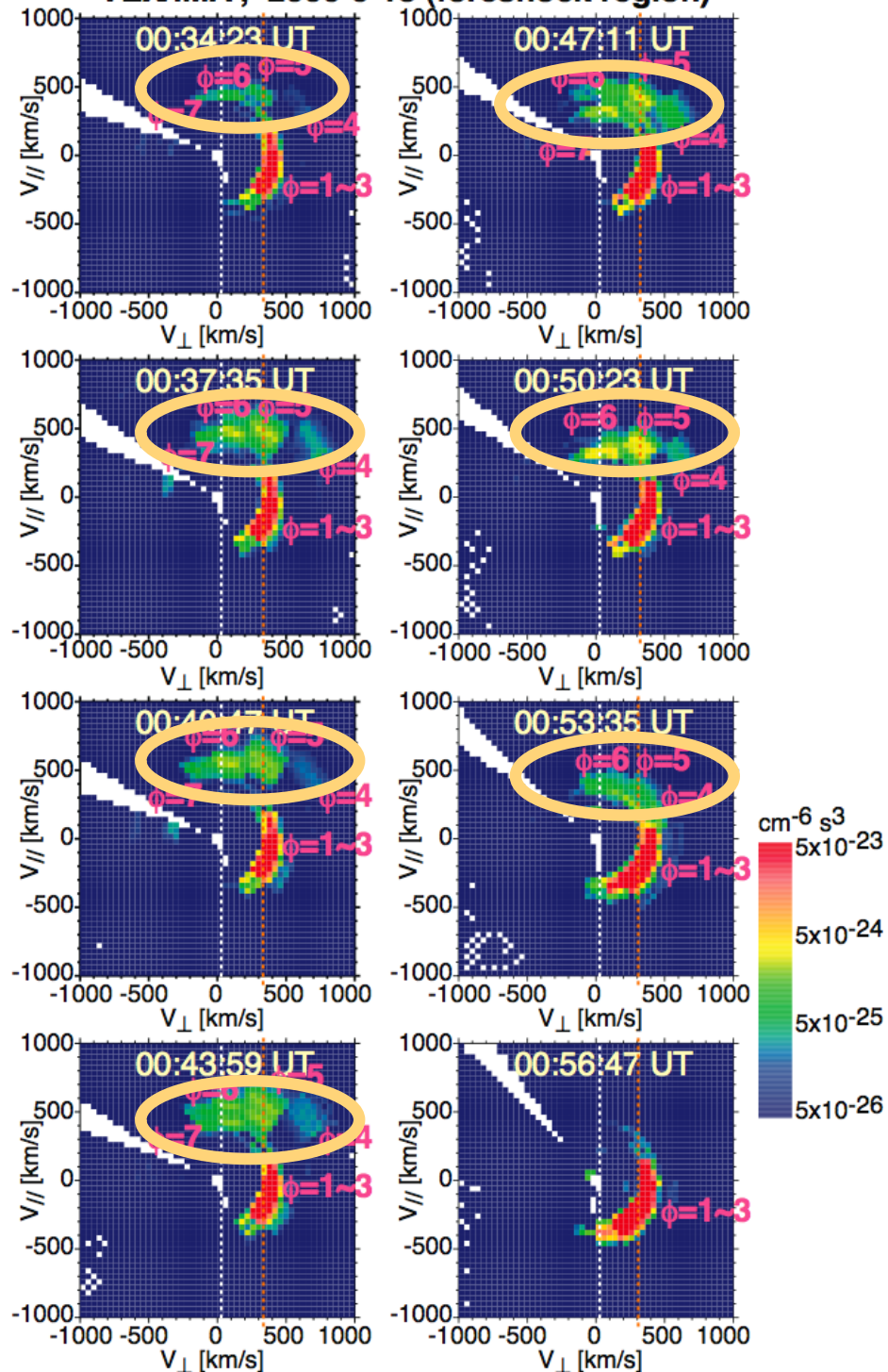
[keV] 00:43:59~ 00:47:11~

IMA looking direction





VEX/IMA, 2006-6-18 (foreshock region)



Two populations (same as Earth)

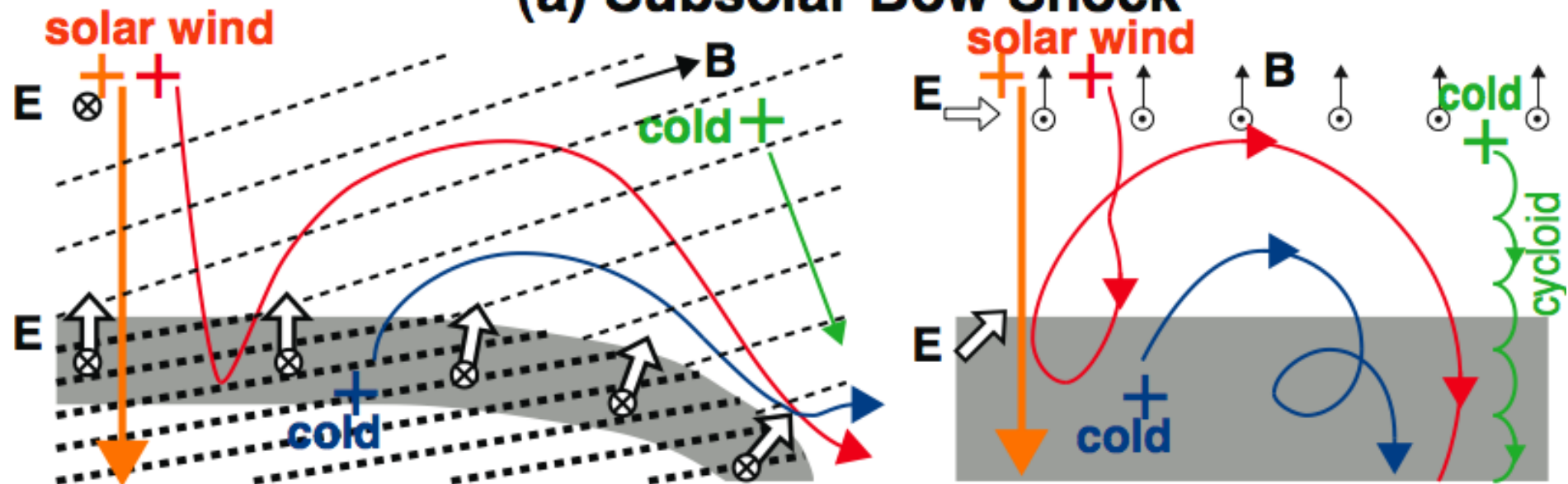
1. Field-aligned H^+ .

2. Gyration H^+ with large $V_{||}$.

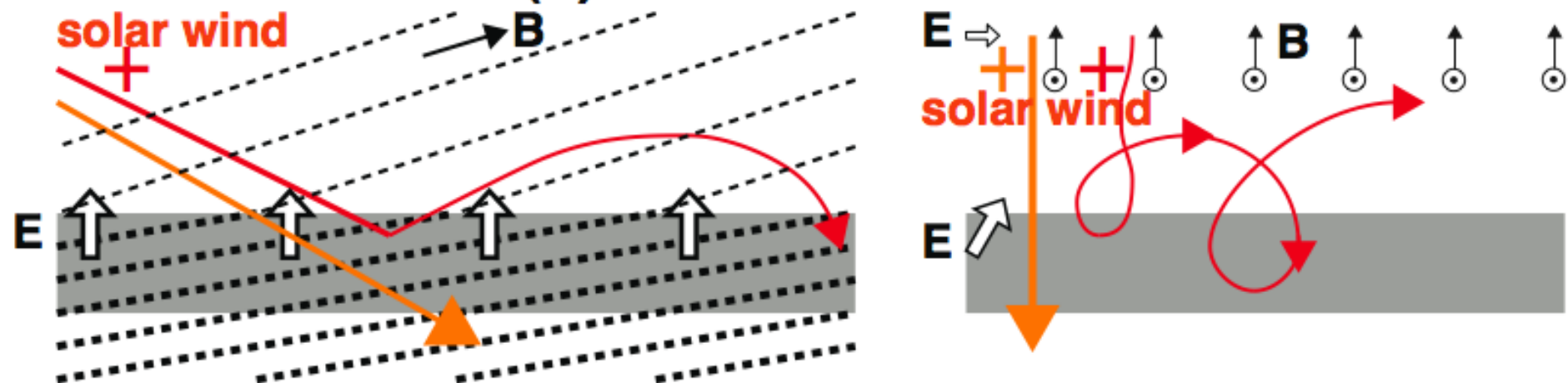
Both types are $\Delta V_{||} \ll V_{||}$,

1. solar wind, 2&3. bow-shock cold ion, 4. sneak out
 $\Delta V_{\parallel} \ll V_{\parallel}$ (yes), $\Delta V_{\parallel} \ll V_{\parallel}$ (yes), $\Delta V_{\parallel} \sim V_{\parallel}$ (no),

(a) Subsolar Bow Shock

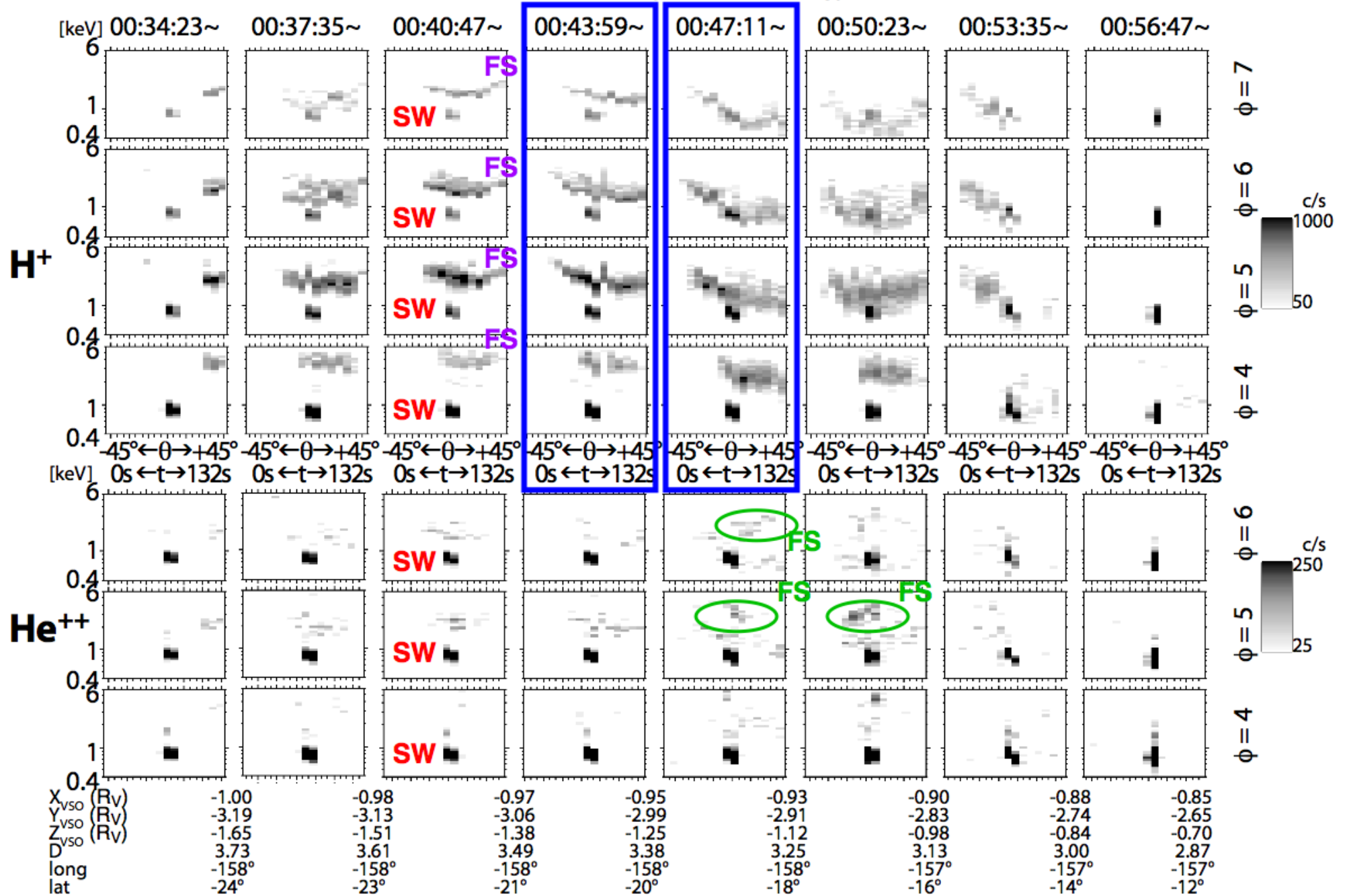


(b) Flank Bow Shock



He++ and H+ show different behavior (future work)

VEX/IMA data (0.4~6 keV), 2006-6-18



Venus $\approx$ Earth

No internal magnetic field.

Planet is the same size as the Earth



Smaller bow shock size than the Earth, yet MHD regime.

Effect of cod ions in the bow shock can be ignored.

Mars $\neq$ Earth

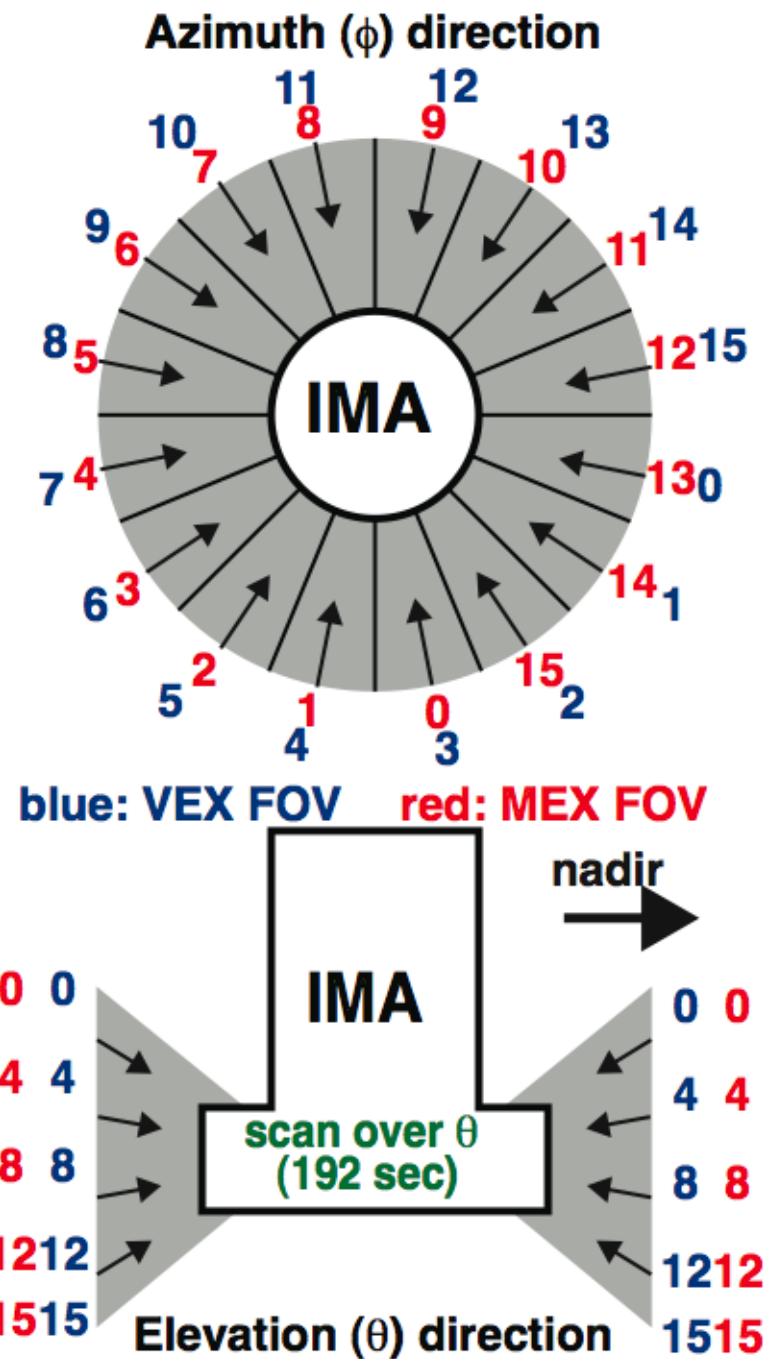
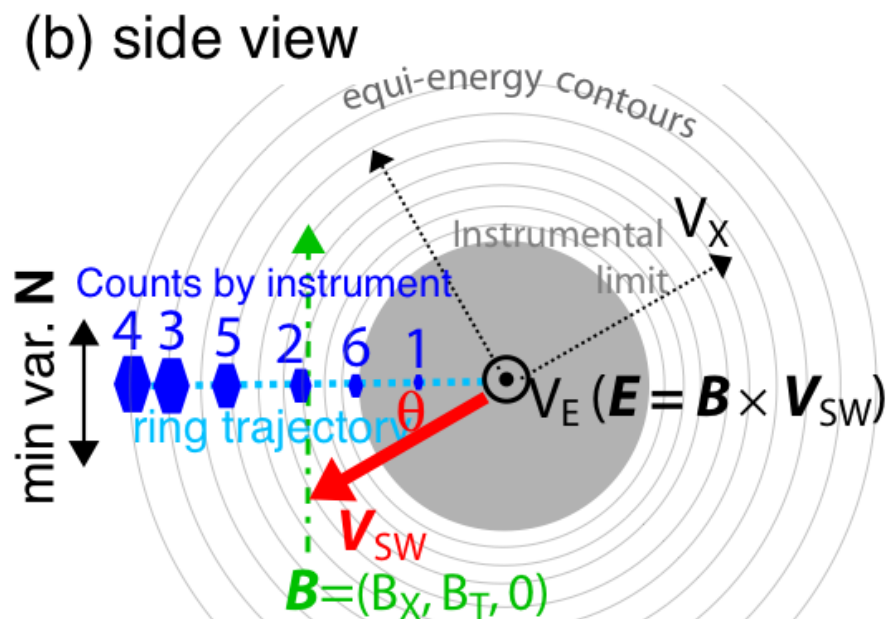
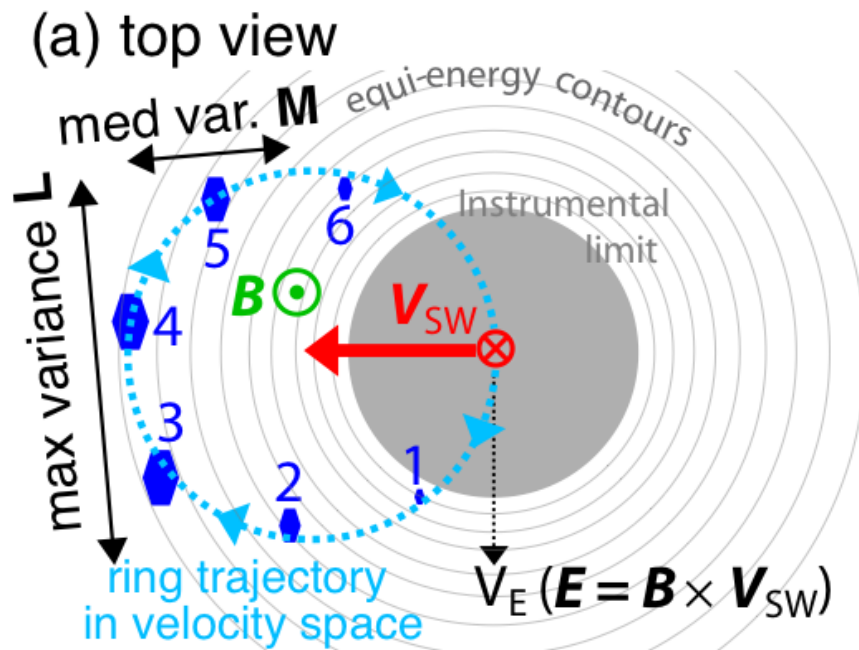
No internal magnetic field.

Planet is smaller than the Earth.

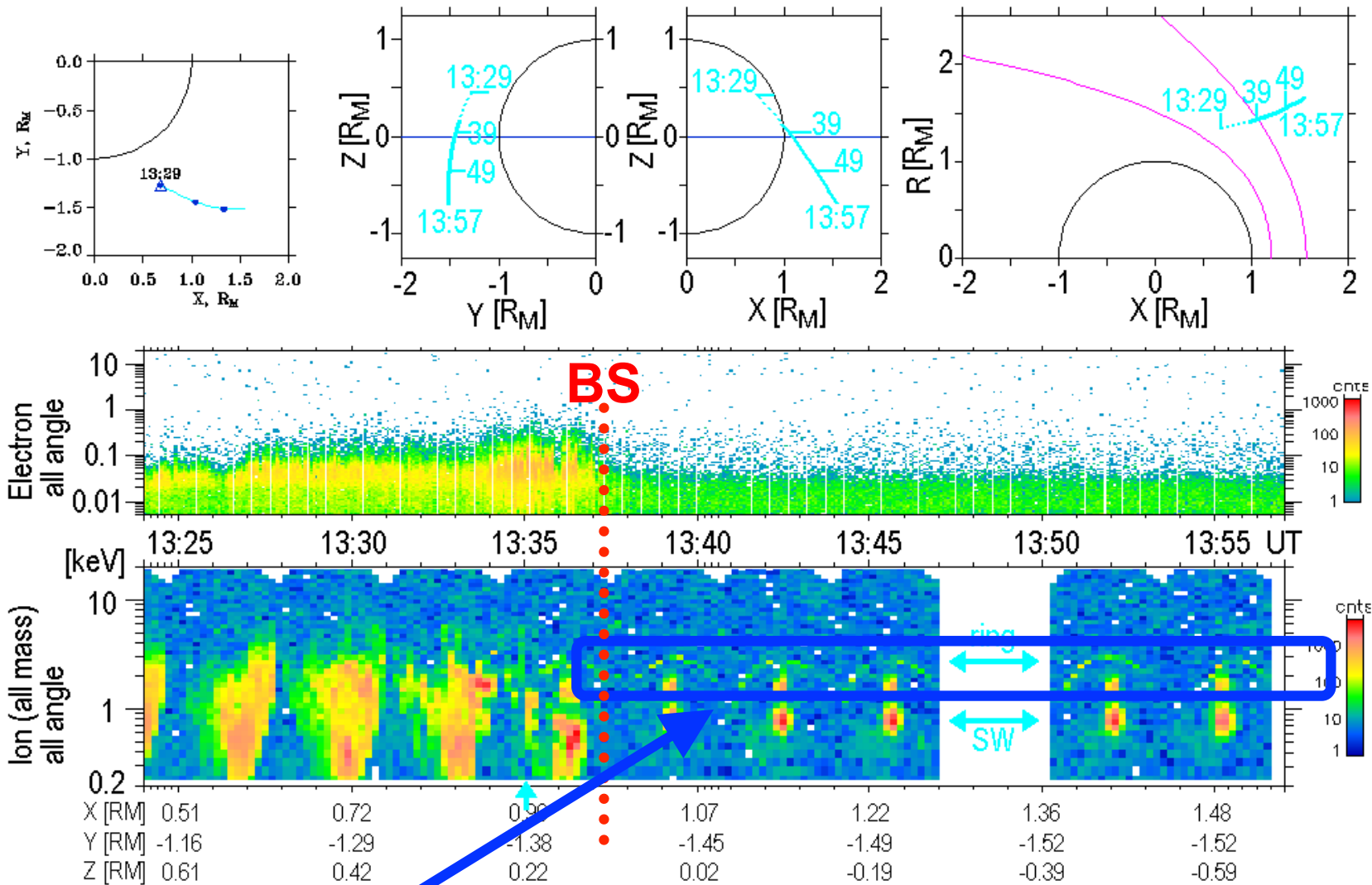


The bow shock size is too small to treat with MHD.

Effect of cod ions in the bow shock cannot be ignored.



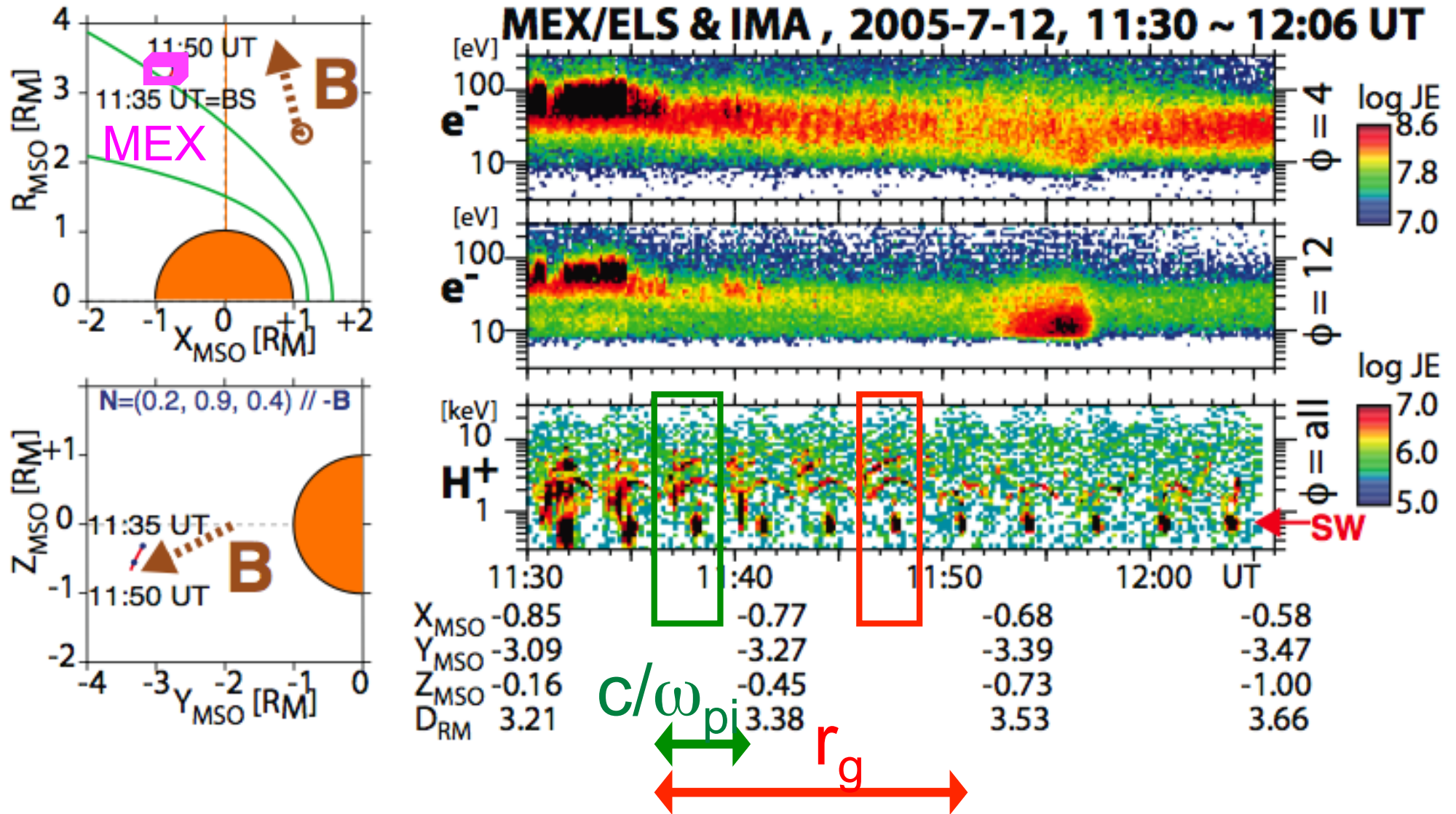
Quite different from Venus:



(1) only "ring" distribution

(2) no "foreshock" signature (examined ~ 500 traversals)

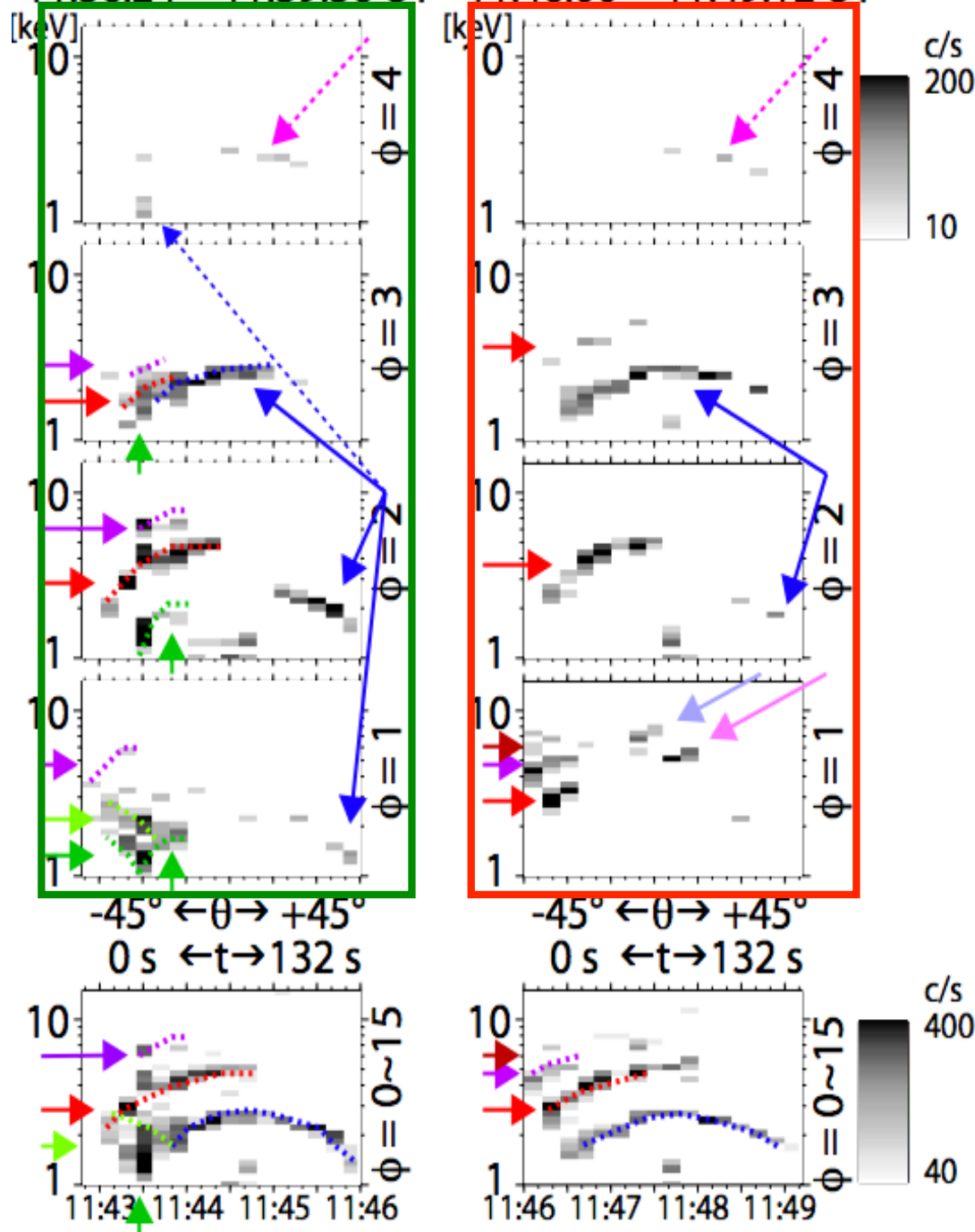
Examine close to the Bow Shock



We sometimes observed “multiple-ring” structure.

MEX/IMA H⁺, 2005-7-12

11:36:24 ~ 11:39:36 UT 11:46:00 ~ 11:49:12 UT



Three types of accelerated ions

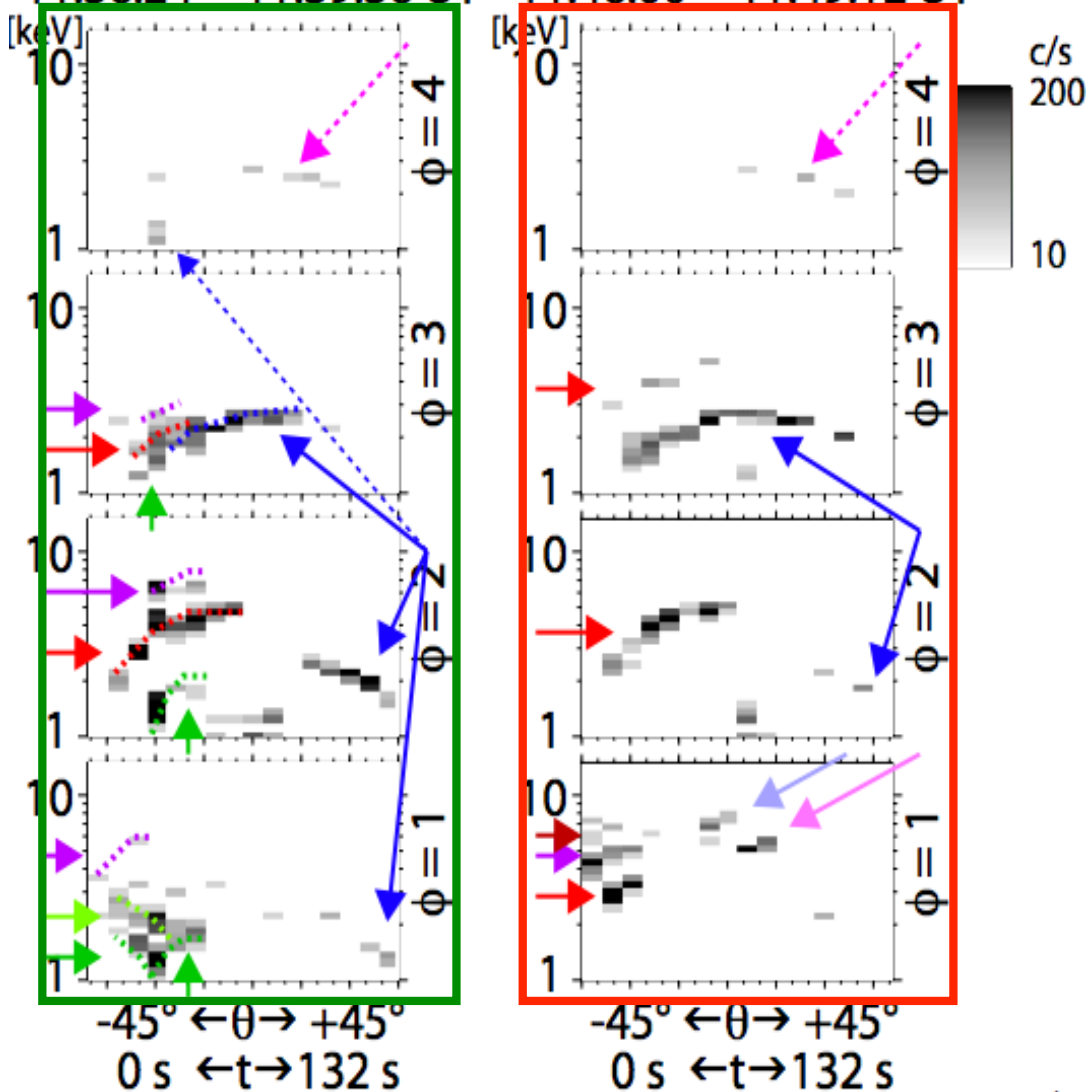
→ beyond r_g
= pickup ions
⇒ obtain B direction

→ beyond c/ω_{pi}
→ within r_g
= reflected ions

→ within c/ω_{pi}
= foot ions

MEX/IMA H⁺, 2005-7-12

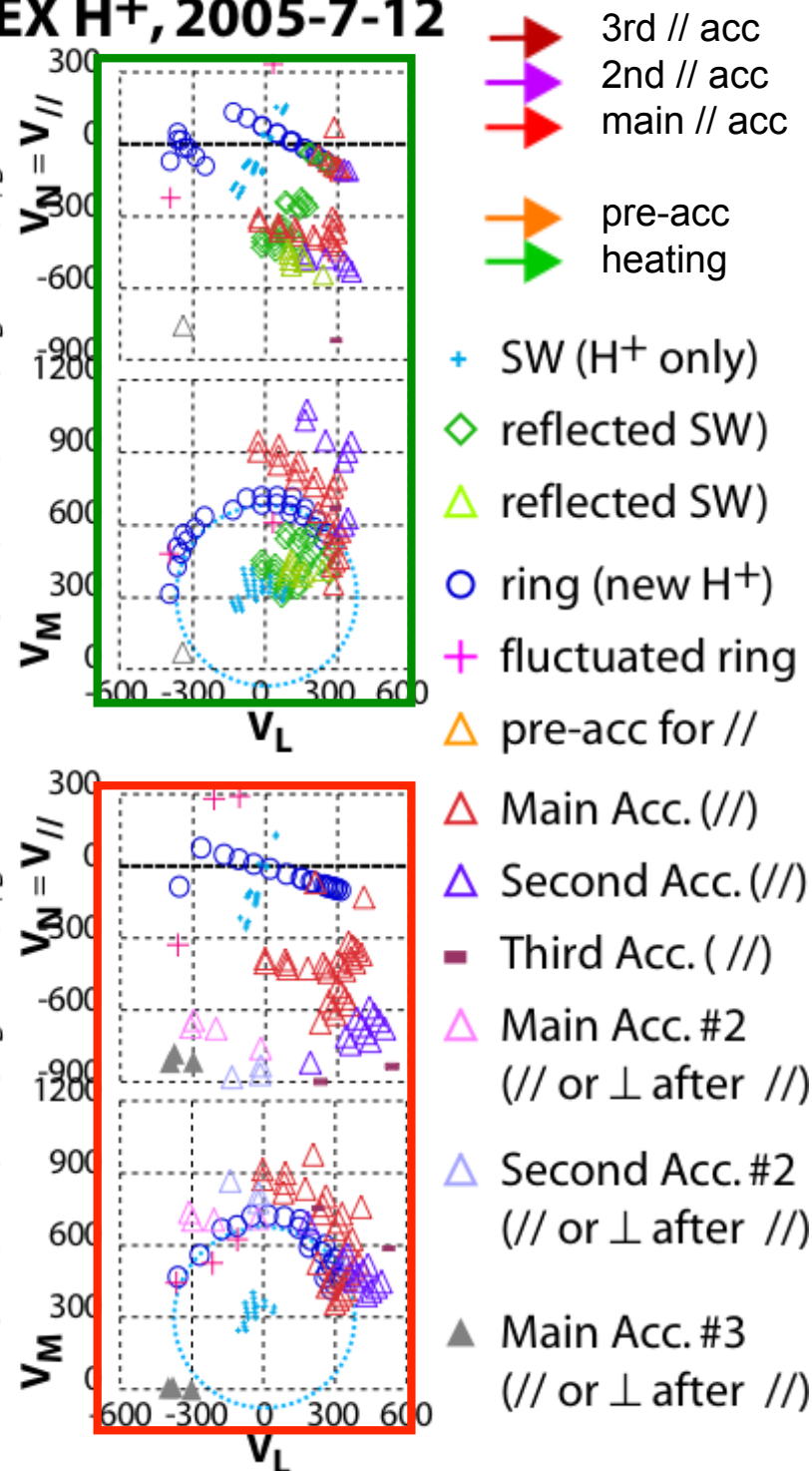
11:36:24 ~ 11:39:36 UT 11:46:00 ~ 11:49:12 UT



MEX H⁺, 2005-7-12

11:36:24 ~ 11:39:36
N = (0.17, 0.85, 0.49) $\theta_B = 10^\circ$, $\phi_B = 60^\circ$

11:46:00 ~ 11:49:12
N = (0.26, 0.91, 0.33) $\theta_B = 15^\circ$, $\phi_B = 70^\circ$



- 3rd // acc
- 2nd // acc
- main // acc
- pre-acc heating

- + SW (H⁺ only)
- ◇ reflected SW
- △ reflected SW
- ring (new H⁺)
- + fluctuated ring
- △ pre-acc for //
- △ Main Acc. (//)
- △ Second Acc. (//)
- Third Acc. (//)
- △ Main Acc. #2 (// or ⊥ after //)
- △ Second Acc. #2 (// or ⊥ after //)
- ▲ Main Acc. #3 (// or ⊥ after //)

Multiple acceleration

green: foot

blue: primary ring

red: 1st branch

purple: 2nd branch

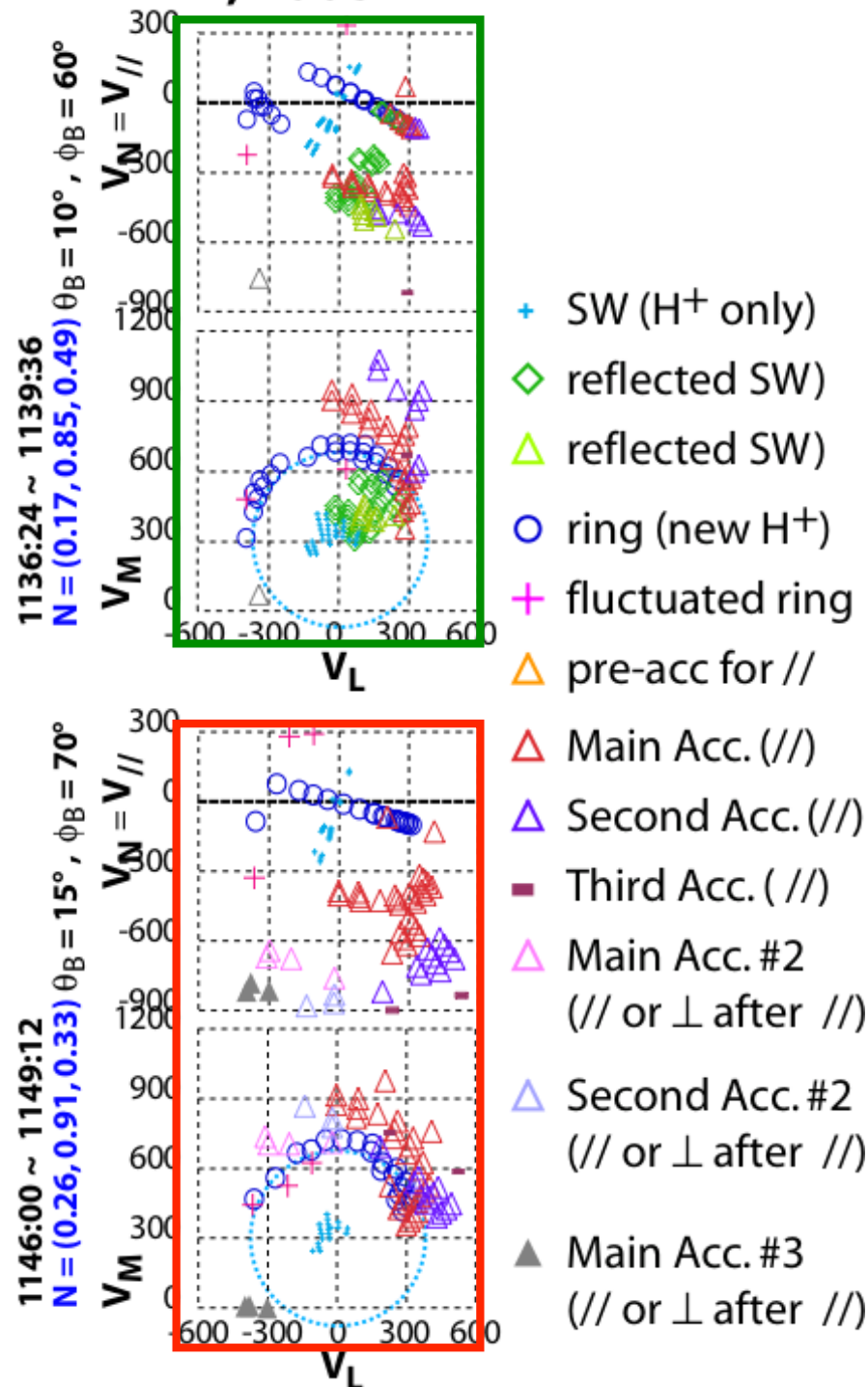
brown: 3rd branch

Gyro-phase bunching

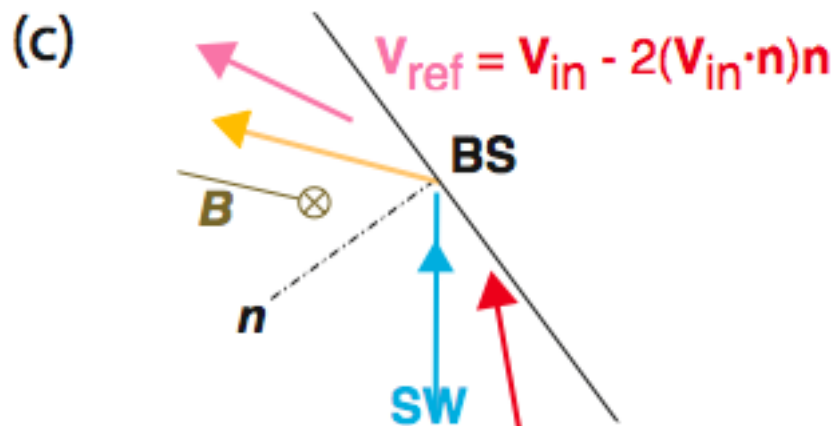
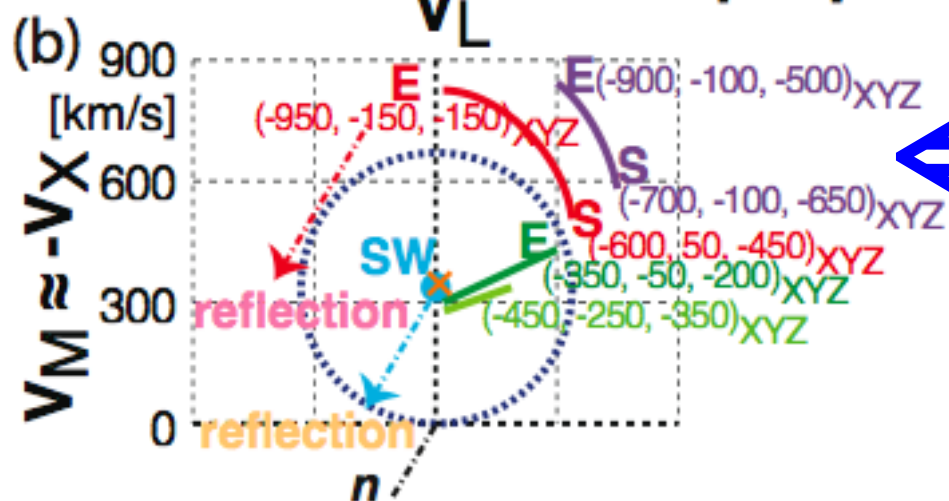
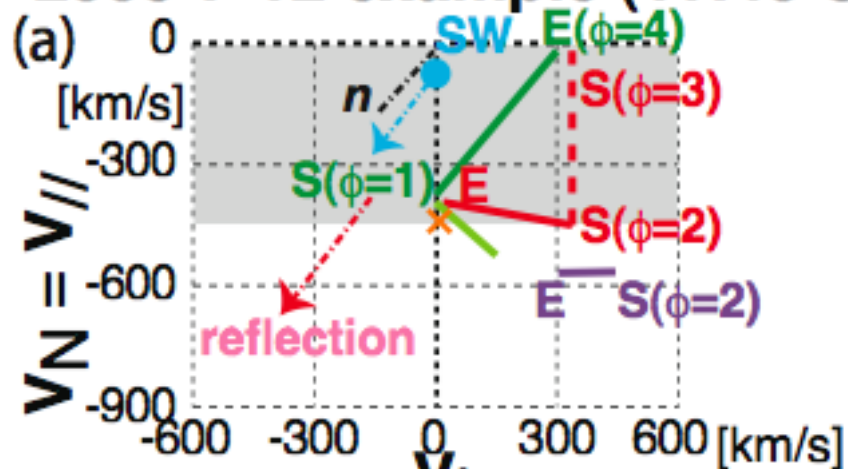
red: half gyro

purple: one third gyro

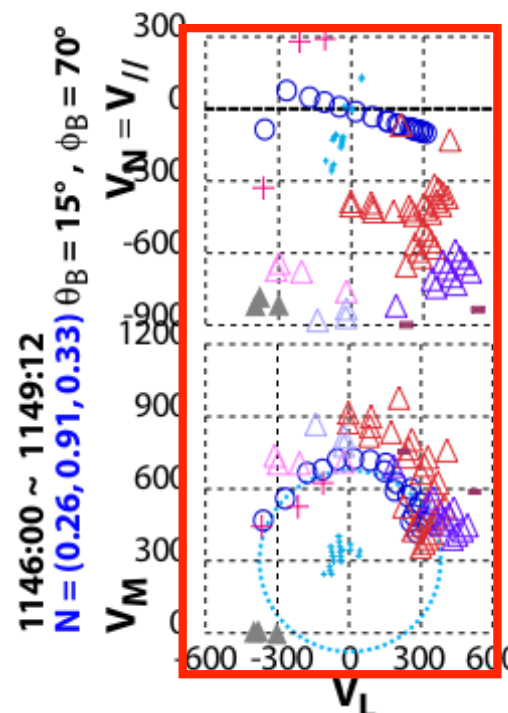
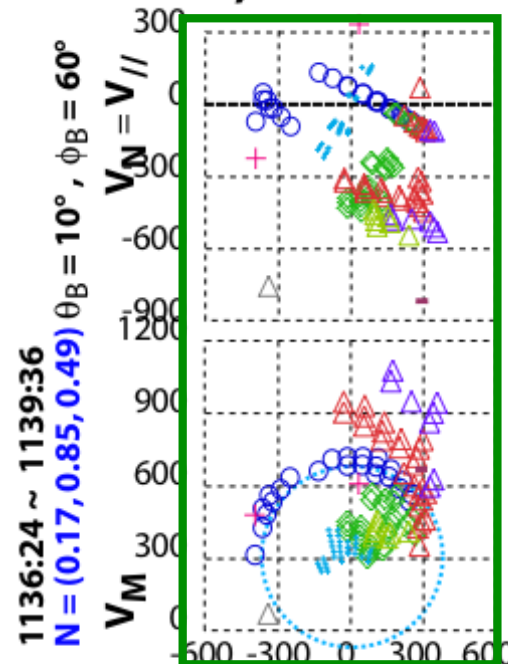
MEX H⁺, 2005-7-12



2005-7-12 example (11:40 UT)



MEX H⁺, 2005-7-12



- + SW (H⁺ only)
- ◇ reflected SW
- △ reflected SW
- ring (new H⁺)
- + fluctuated ring
- △ pre-acc for //
- △ Main Acc. (//)
- △ Second Acc. (//)
- Third Acc. (//)
- △ Main Acc. #2 (// or ⊥ after //)
- △ Second Acc. #2 (// or ⊥ after //)
- ▲ Main Acc. #3 (// or ⊥ after //)

2005-7-12 example (11:40 UT)

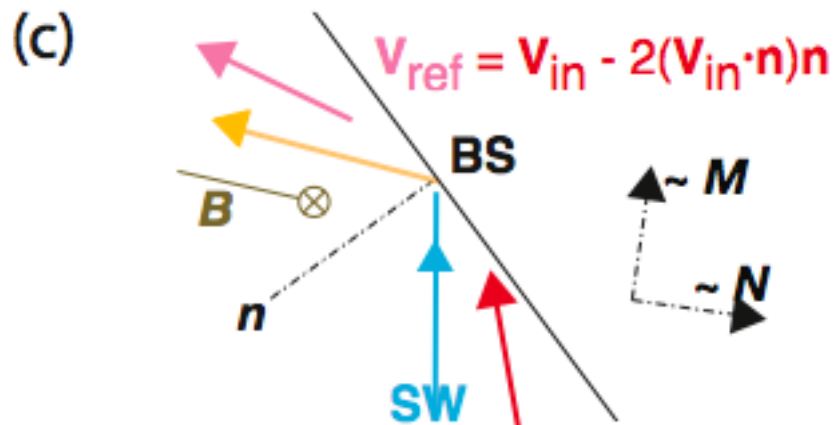
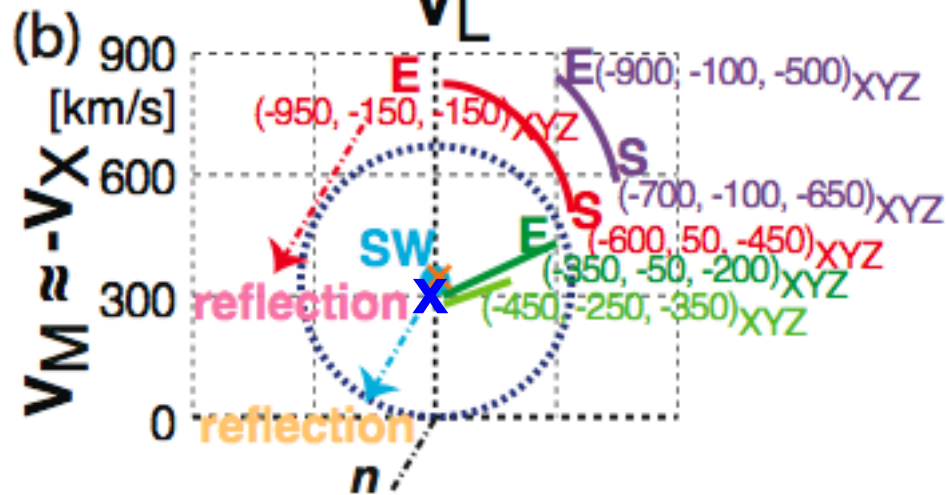
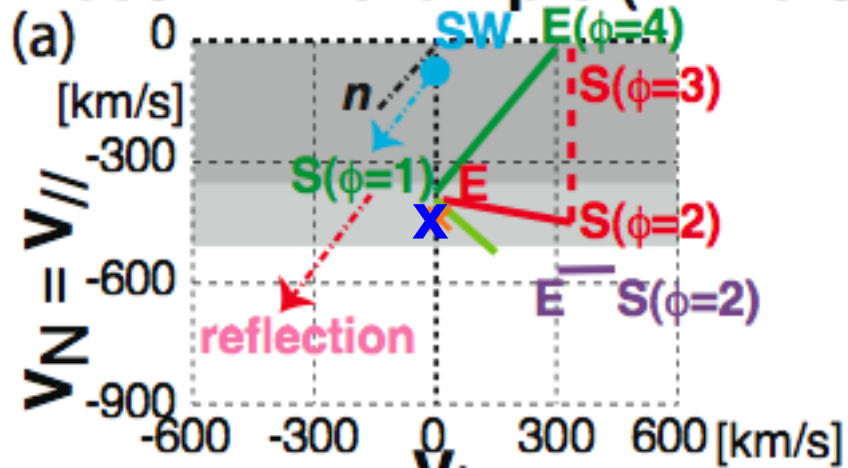
Multiple-Reflection

n : shock normal

V_R : specularly reflected SW

$\times$ V_{HT} : de Hoffman Teller ($V'_{SW} \parallel B$)

e.g., 11:40 UT

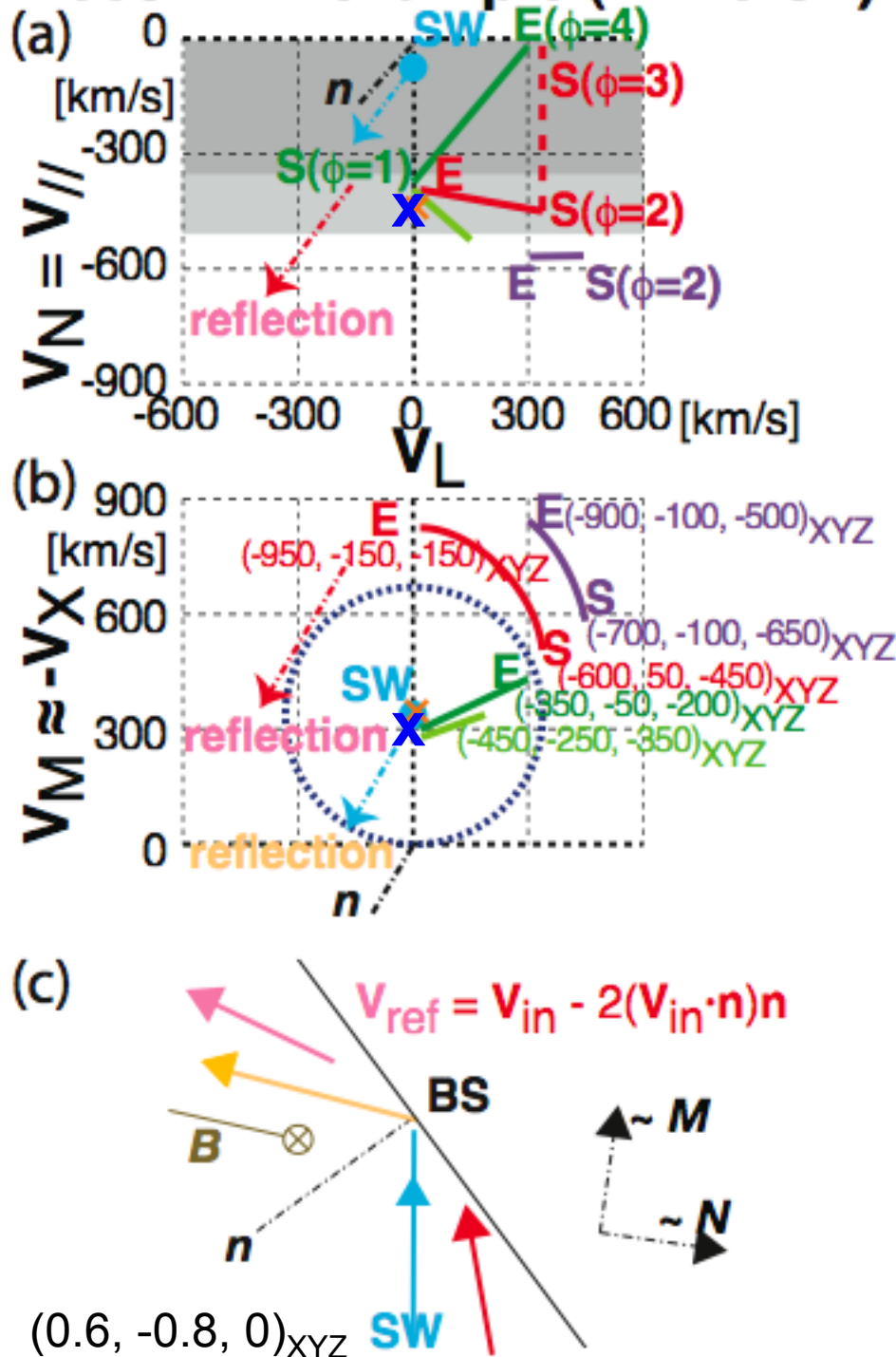


L	$\sim (0, 0.6, -0.8)_{XYZ}$
M	$(-1.0, 0.2, 0.1)_{XYZ}$
$N = -B/B$	$(0.2, 0.8, 0.6)_{XYZ}$

n	$(-0.5, -0.7, -0.5)_{LMN}$ $\pm(0.05, 0.05, 0.05)$
V_{SW}/V_{SW}	$(0.0, 1.0, 0.2)_{LMN}$
V_R/V_{SW}	$(-0.6, 0.1, -0.8)_{LMN}$ $\pm(0, 0.2, 0)$
V_{HT}/V_{SW}	$(0., 1.0, -1.4)_{LMN}$ $\pm(0, 0, 0.3)$

2005-7-12 example (11:40 UT)

Multiple-Reflection



S: toward BS from left
E: toward BS from right



S&E: toward BS from left



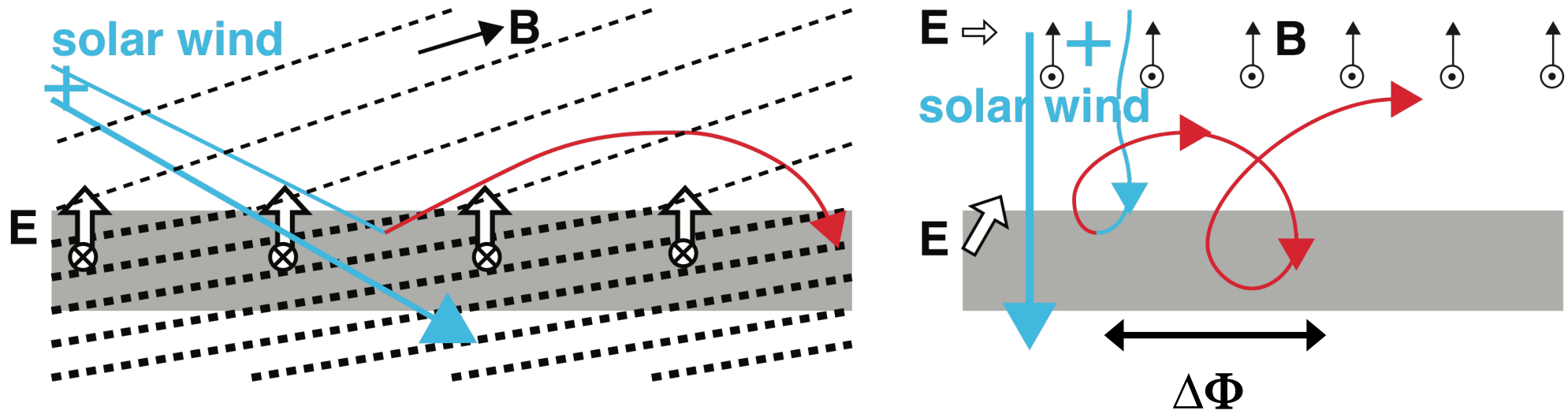
$S \sim V_{HT} =$ along BS
E: along BS



S: along BS
E: toward BS

SW Reflection $\Rightarrow$ convert $V_{\perp}$ to $V_{\parallel}$ in SW frame

Flank Bow Shock



The observed multiple ring structure is well explained by multiple specular reflection.

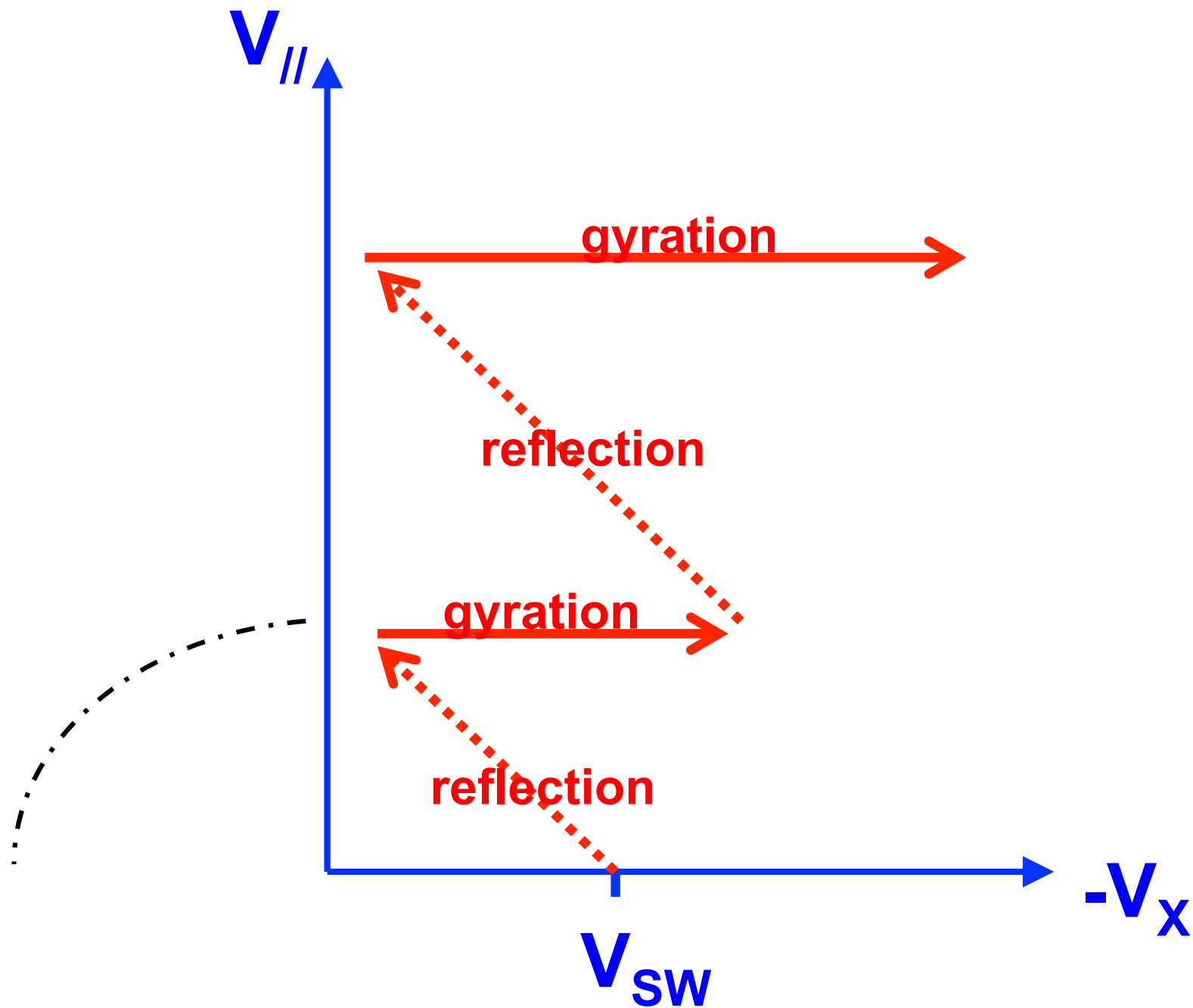
But, **why is it observed outside the foot region?**



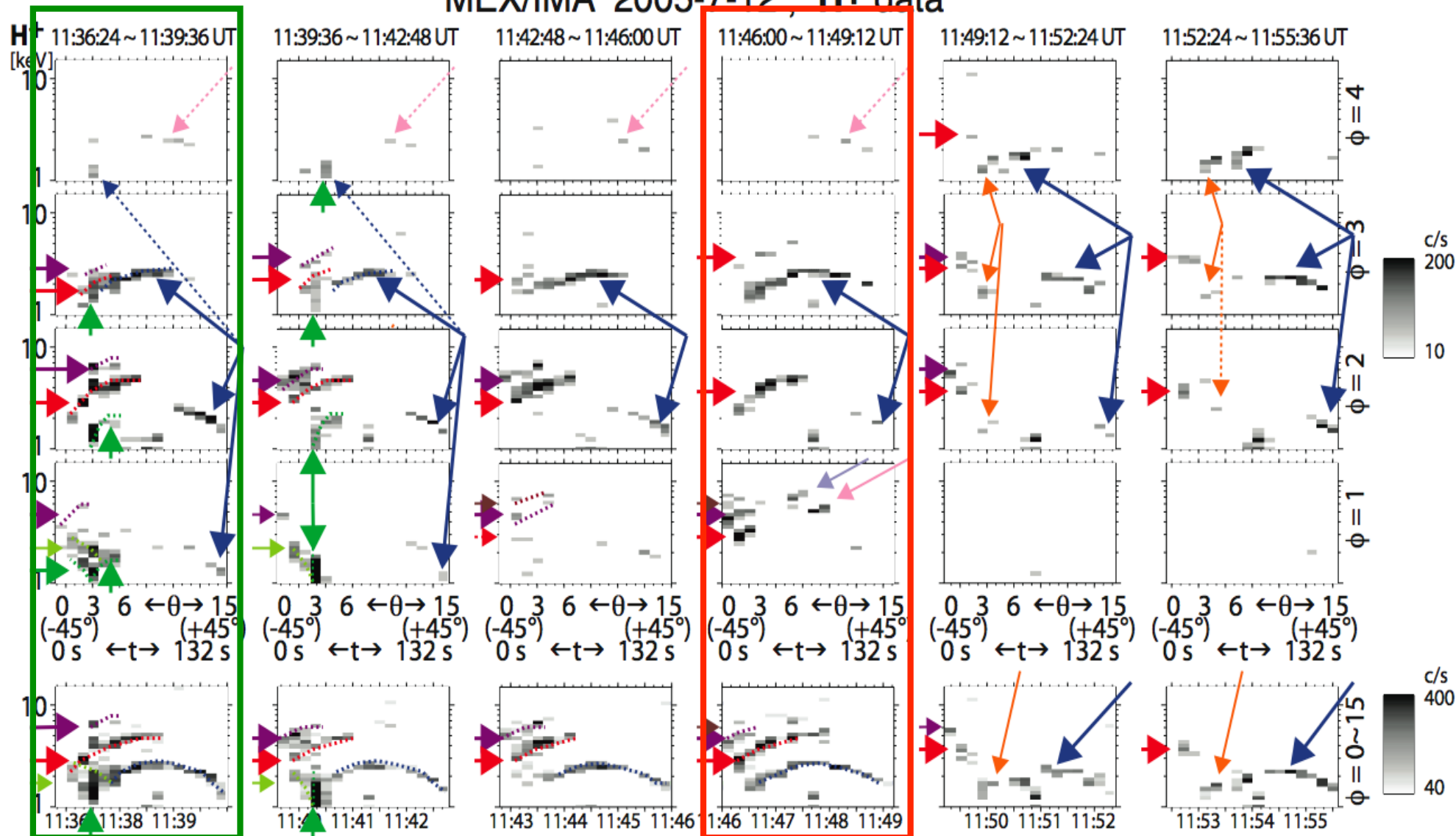
no : Finite bow shock size compared to r_g .

yes: Cold ion in the bow shock

$\Rightarrow$ This may explain “non-specular reflection” at subsolar.

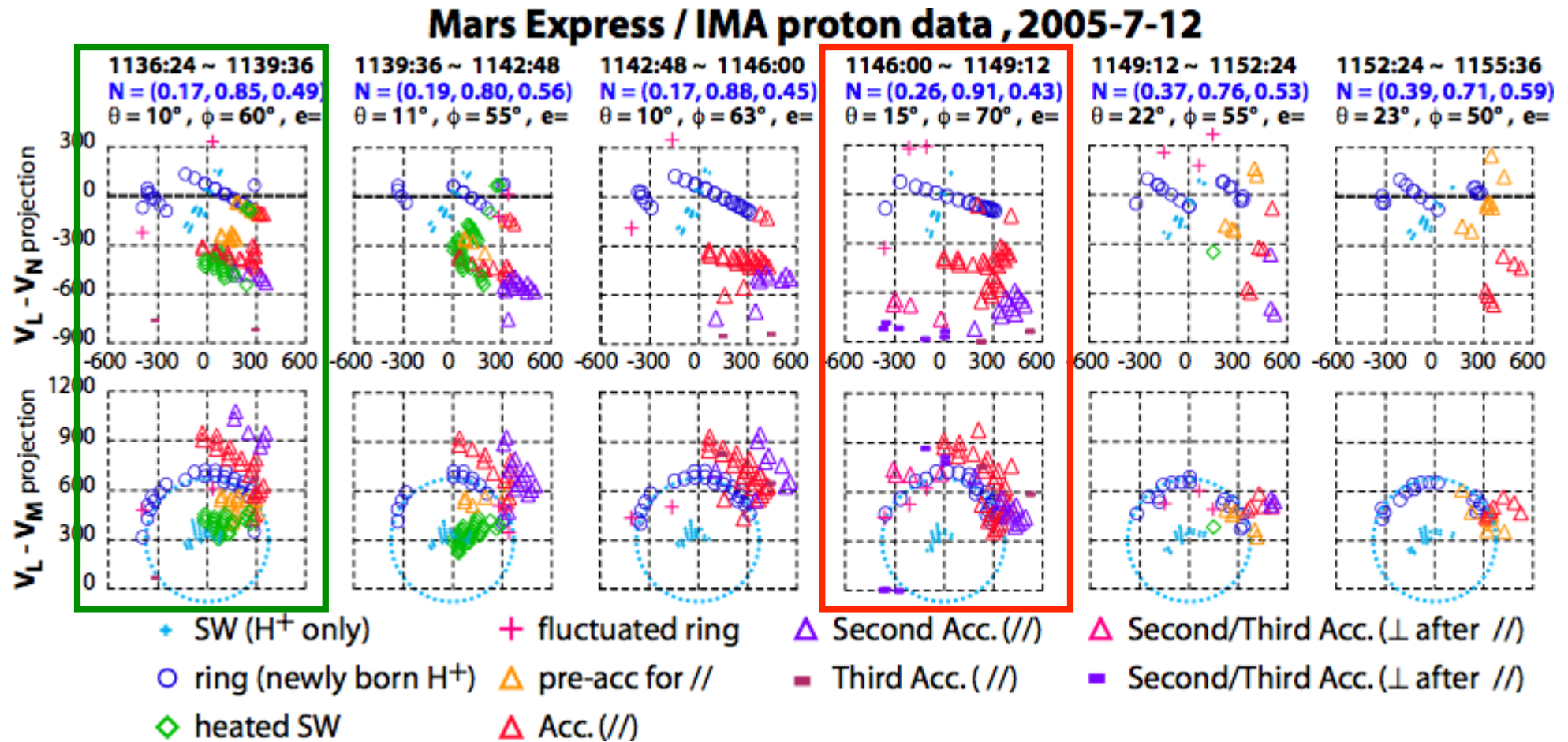


MEX/IMA 2005-7-12, H⁺ data



Time = Spatial variation

Classifying counts in // and $\perp$ directions

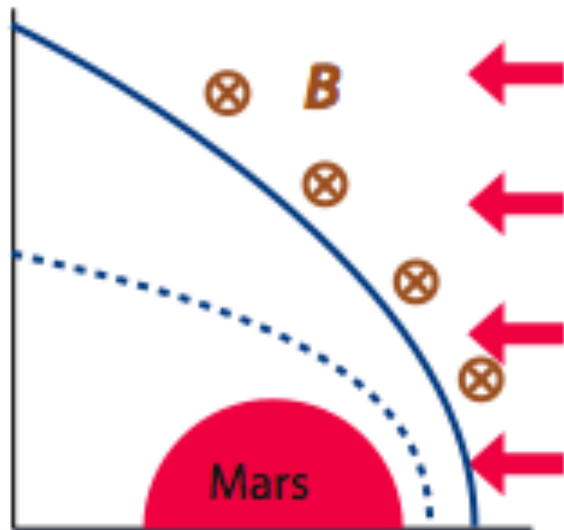


B (N-direction) is estimated from minimum variance method applied to the ring distribution

Time = Spatial variation

Three configurations (on-going work)

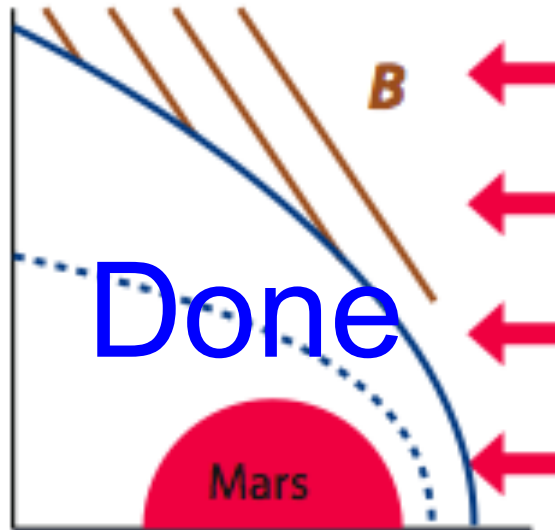
(1) QP outside foreshock



multiple bouncing?

2005-7-29
2005-8-3

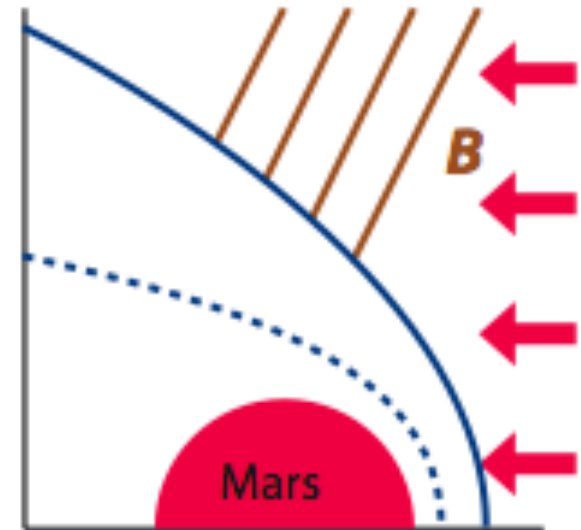
(2) QP inside foreshock



field-aligned beam?

2005-7-12

(3) QL



field-aligned beam?

2005-8-5

Done

Special features for Mars

- Energy is stepping (due to reflection?)
- Gyro-bunching effect (due to short distance?) with gradual $\perp$ acceleration (why?)
- Two different scale length
- No specular reflection near the bow shock (need to confirm)

Summary

Venus Express / ASPERA-4 often observes back-streaming H^+ in the foreshock region of Venus, in a similar ways as the Terrestrial foreshock, i.e., field-aligned component, and intermediate (gyrating) component

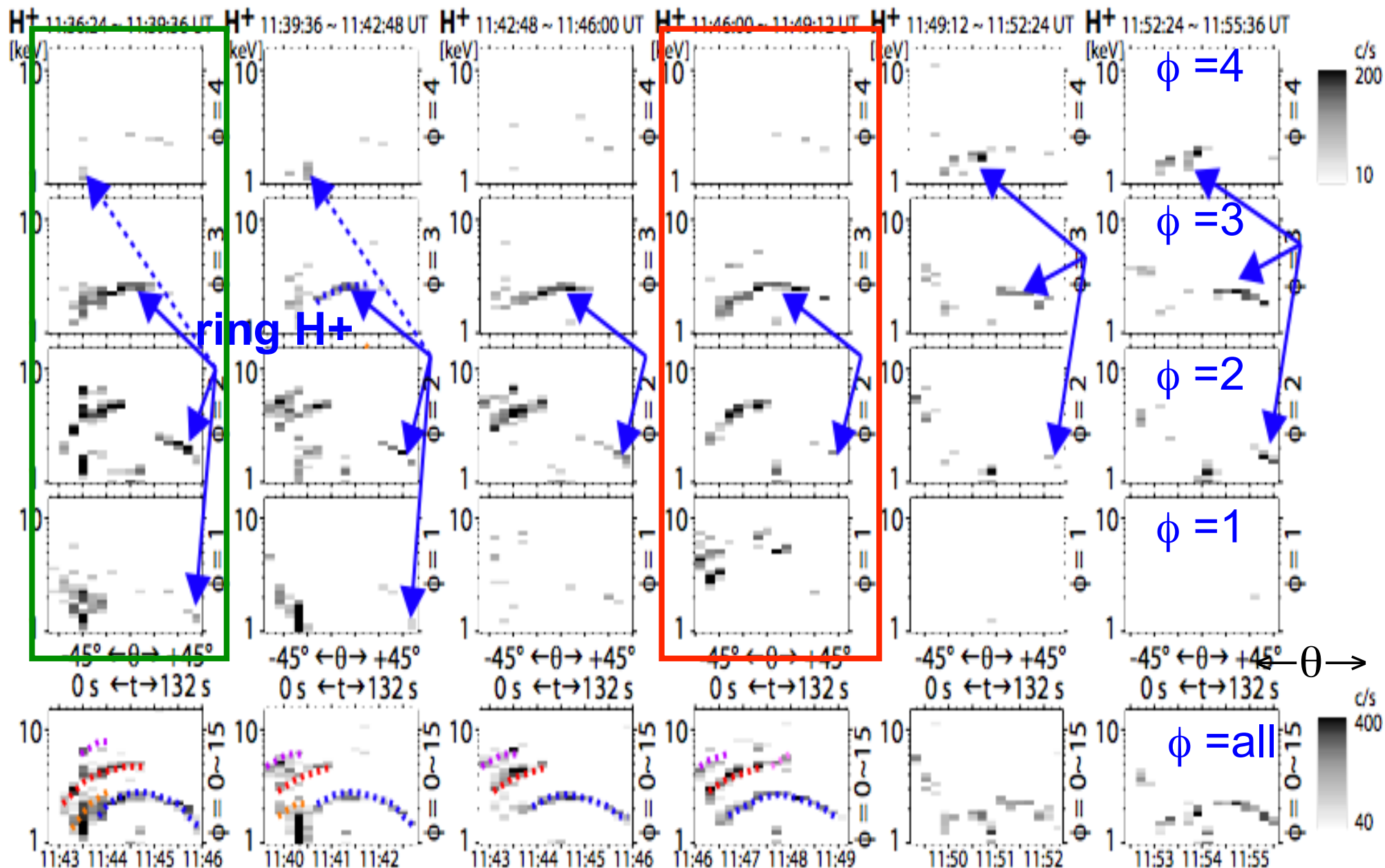
Mars Express / ASPERA-3 (same instrumentation as VEX) did not observe similar ions in the Martian foreshock region beyond the foot region. Instead, it shows different type of acceleration in the foot region, indicating the ion trajectory (history) during its gyromotion.

The finite gyroradius effect makes Mars a perfect laboratory to study acceleration processes.

End

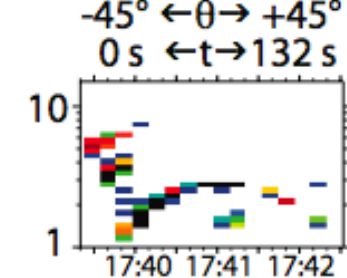
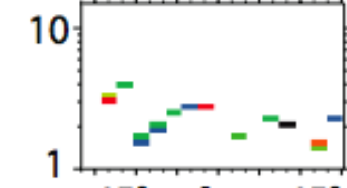
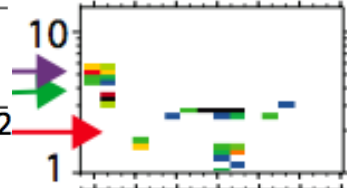
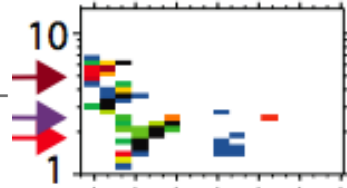
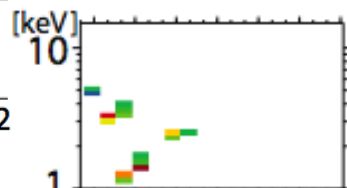
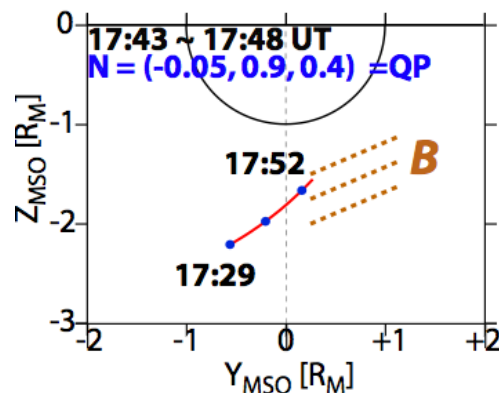
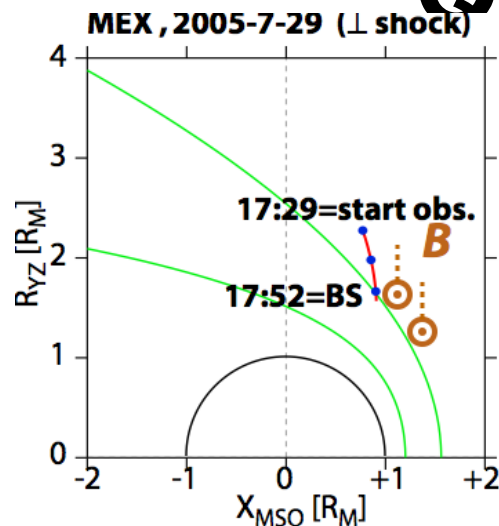
Detail of spectrogram (E=1~15 keV)

MEX/IMA 2005-7-12



Mixture of accelerated ion (H⁺) components

Quasi- $\perp$ shock (case 1)

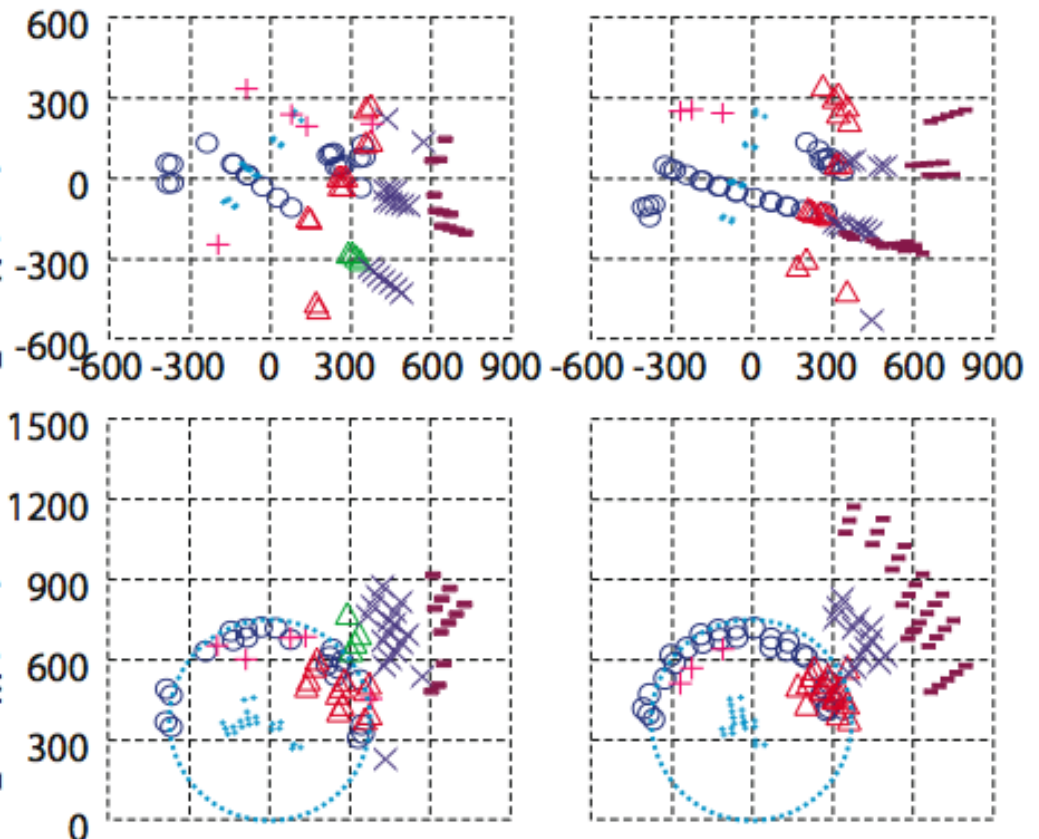


MEX / IMA H⁺, 2005-7-29

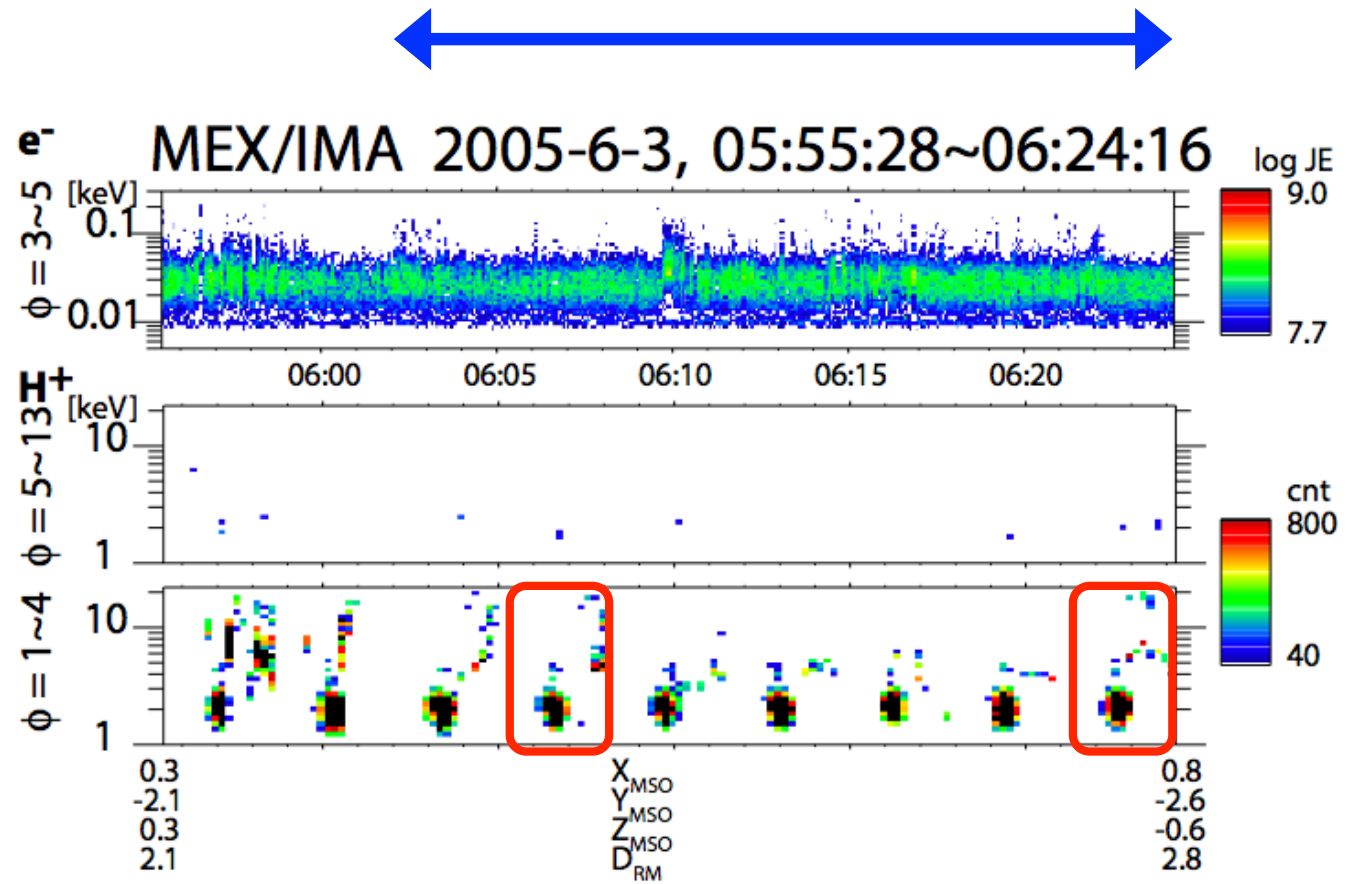
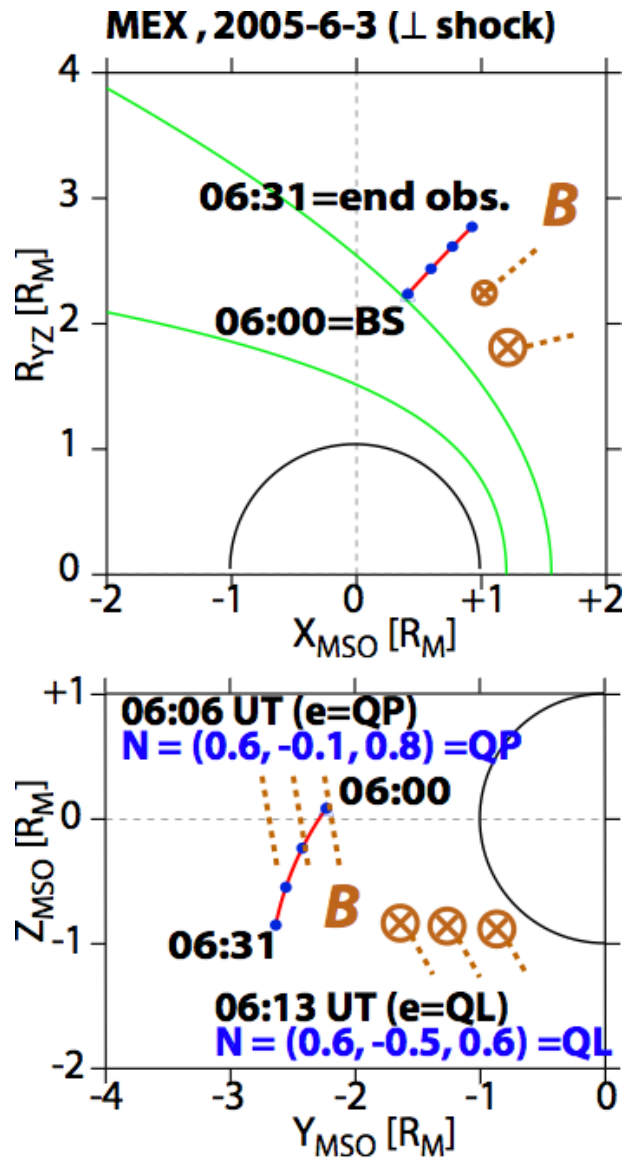
1739:12 ~ 1742:24
 $N = (-0.10, 0.76, 0.64) = QP$
 $\theta = -6^\circ, \phi = 50^\circ, e =$

1742:24 ~ 1745:36
 $N = (-0.05, 0.94, 0.34) = QP$
 $\theta = -3^\circ, \phi = 70^\circ, e =$

$V_L - V_M$ projection $V_L - V_N$ projection

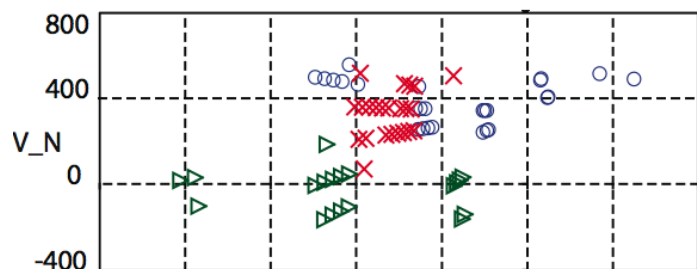


Quasi- $\perp$ and $\parallel$ (case 1+3)

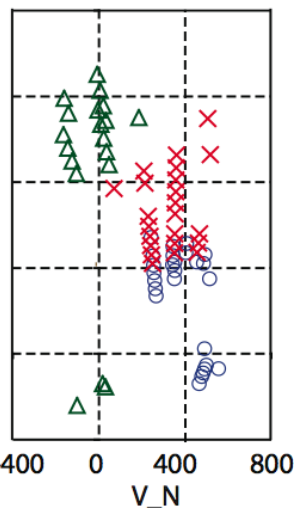
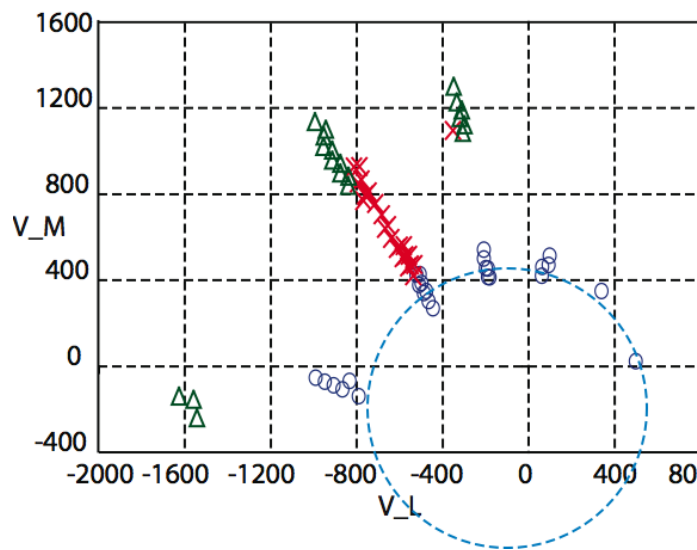


2005-6-3, 0608 UT

$N = (+0.57, +0.14, +0.81)$

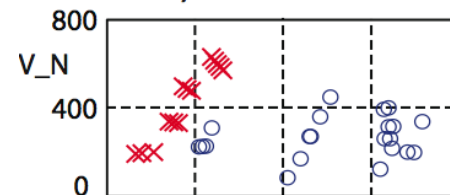


- #1=ring
- × #3=⊥ accel.
- △ #5=2nd ring

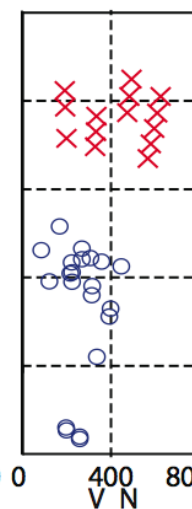
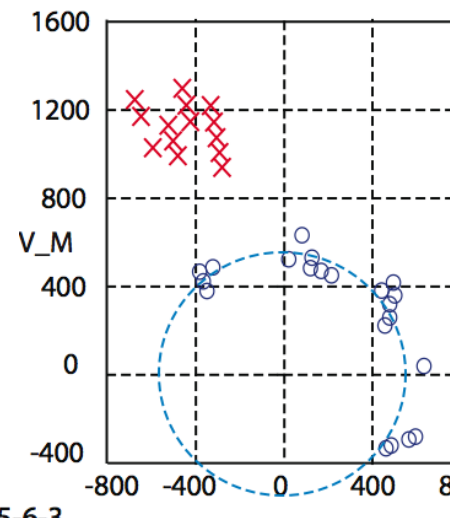


2005-6-3, 0624 UT

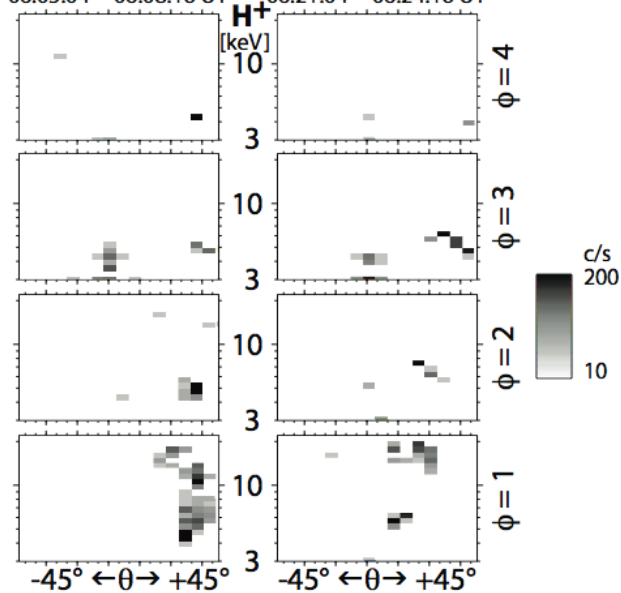
$N = (+0.42, -0.45, +0.79)$



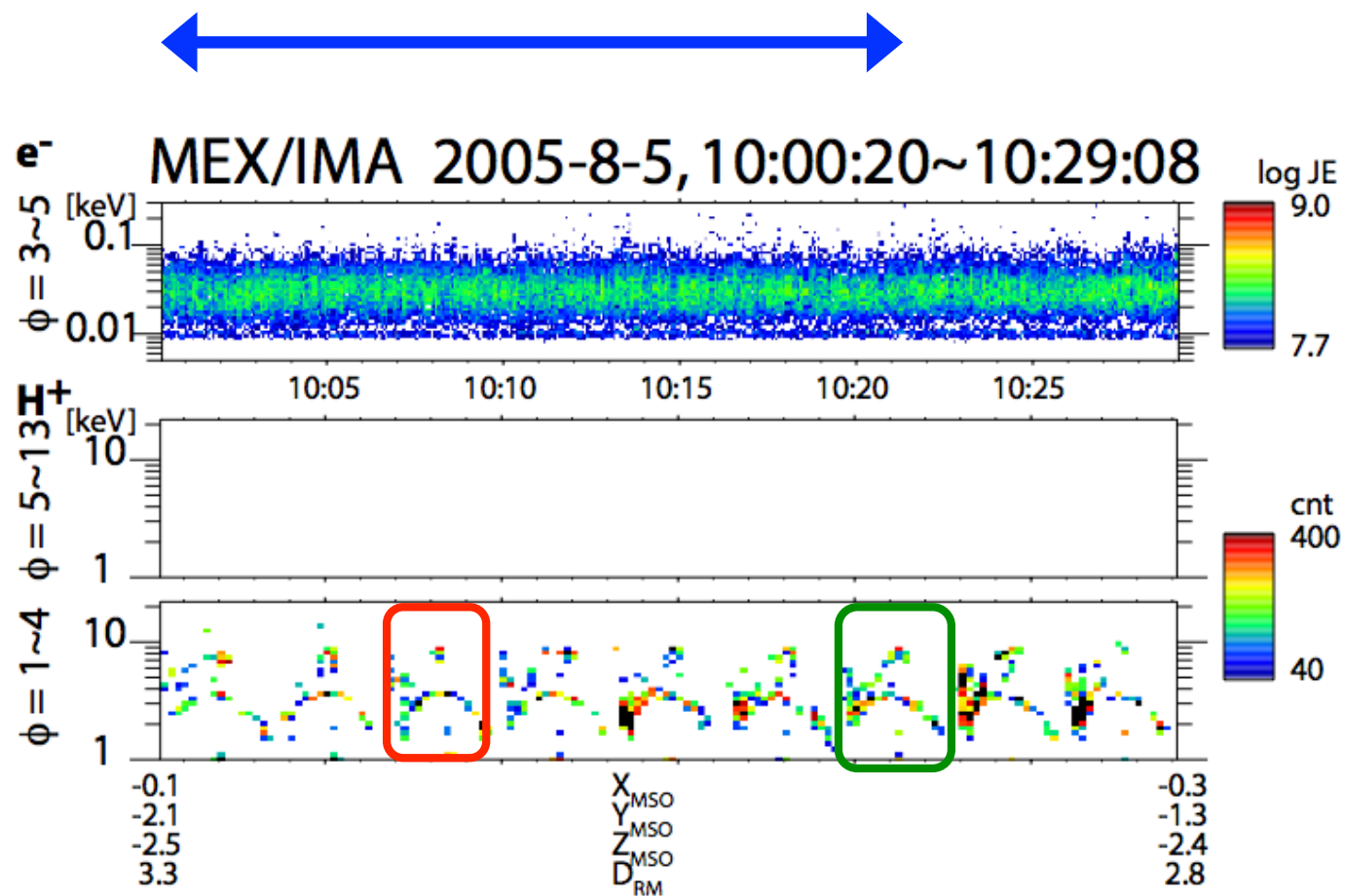
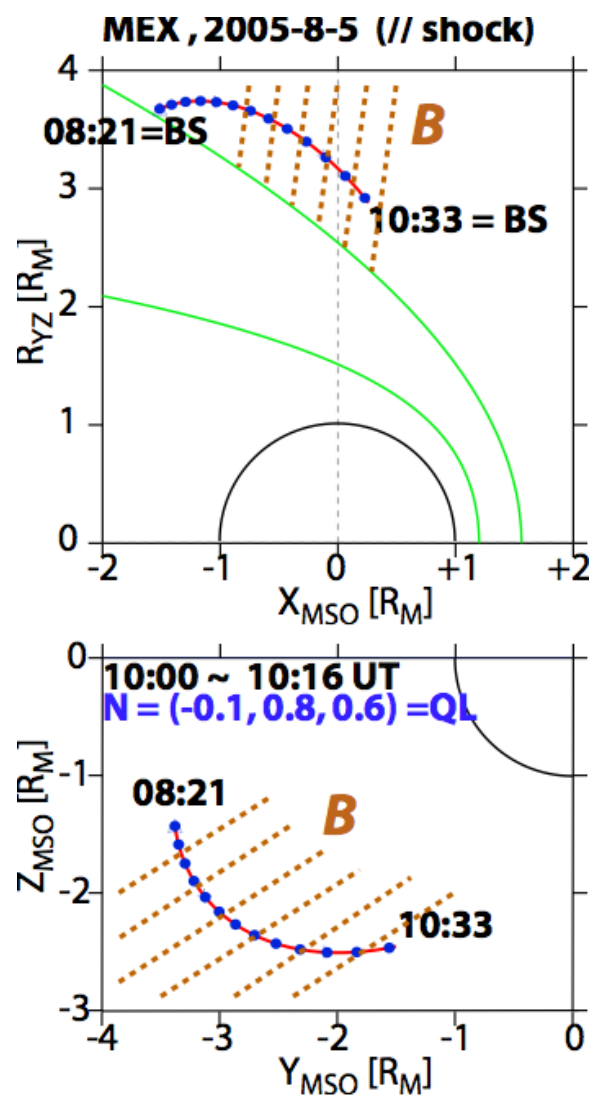
- #1=ring
- × #3=⊥ accel.



MEX/IMA 2005-6-3
06:05:04 ~ 06:08:16 UT 06:21:04 ~ 06:24:16 UT

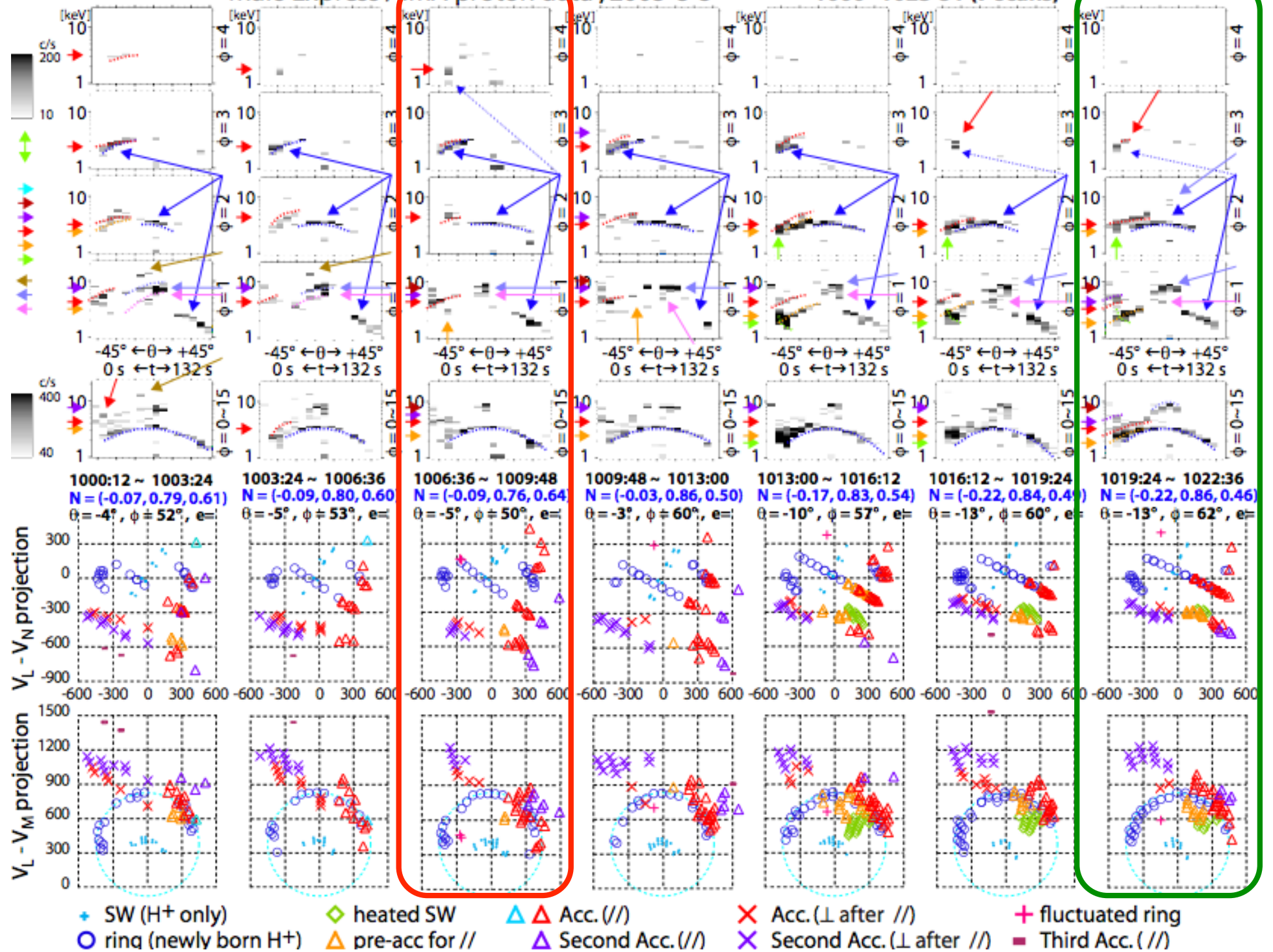


Quasi-// (case 3)



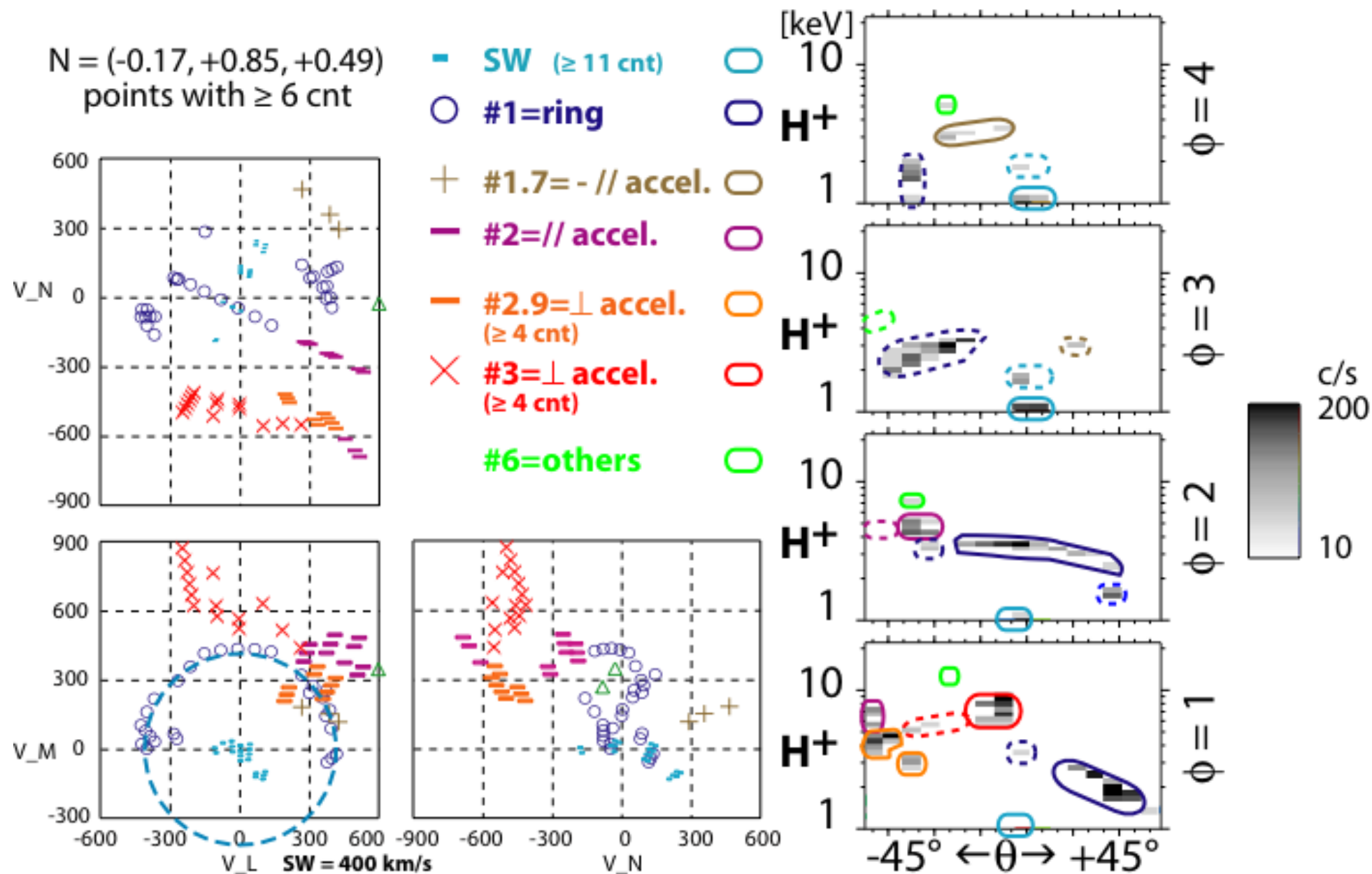
Mars Express / JMA proton data, 2005-8-5

1000~1023 UT (7 scans)



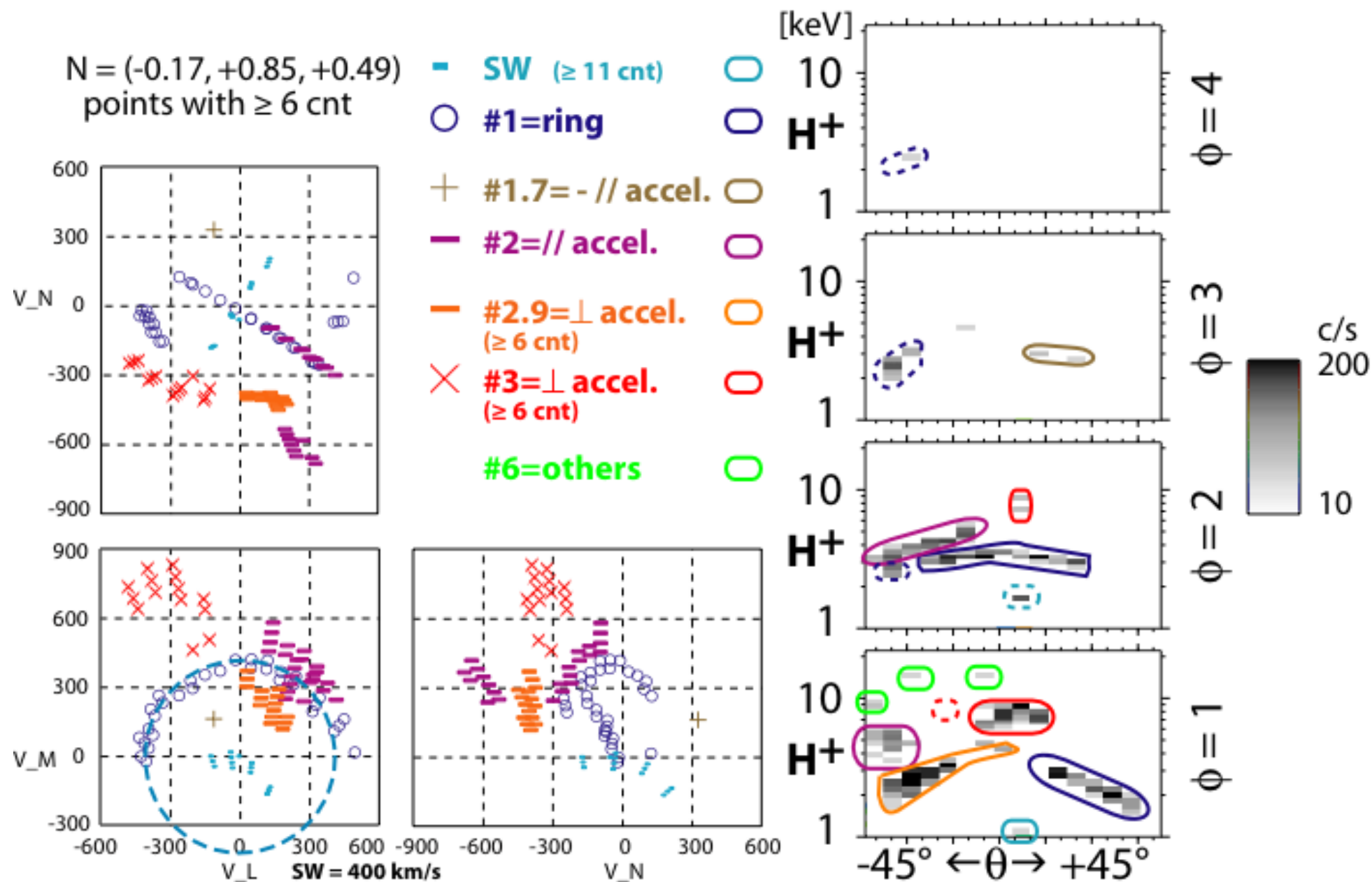
case 3a

MEX/IMA 2005-8-5, 1006:24~1009:36 UT

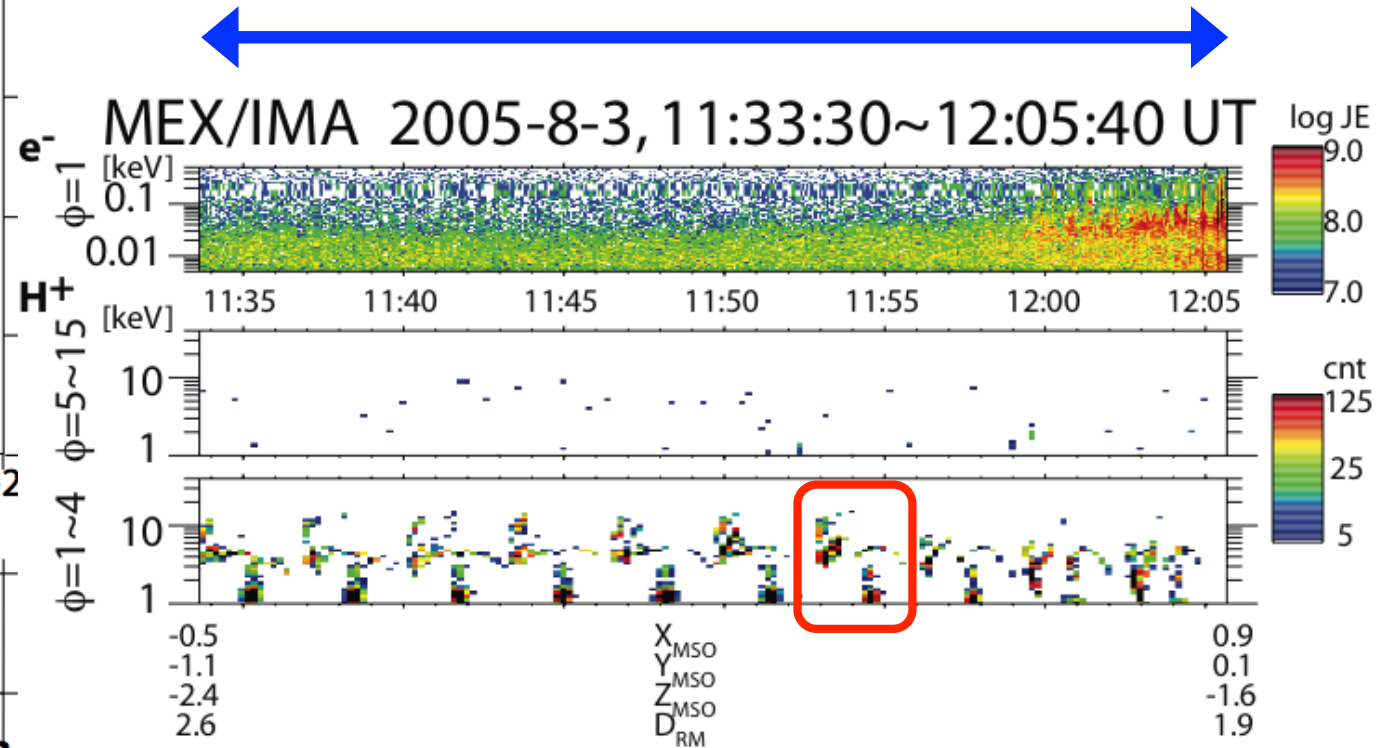
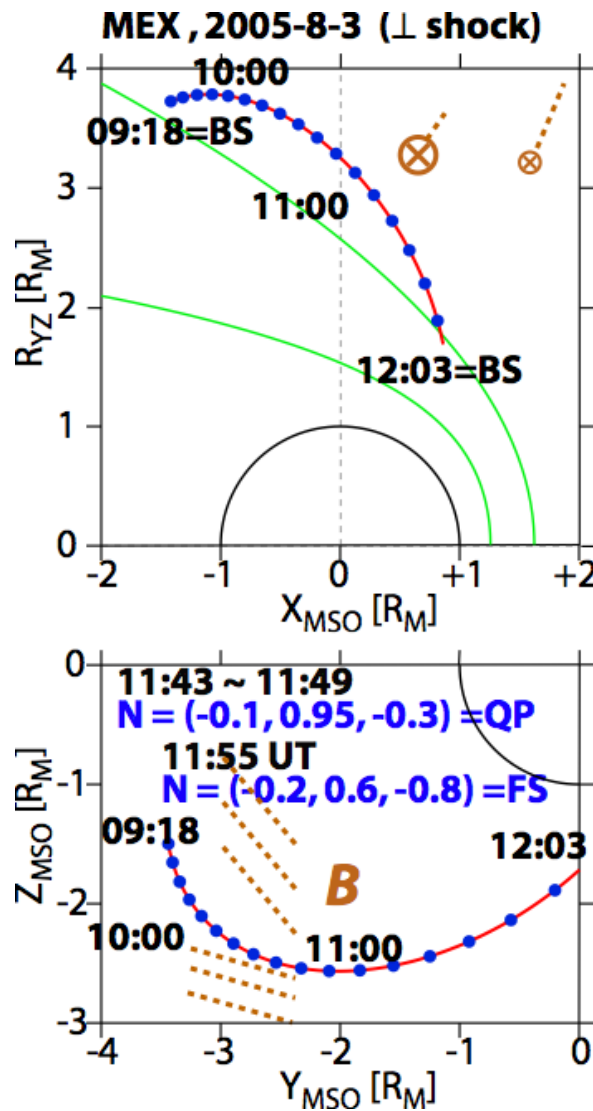


case 3b

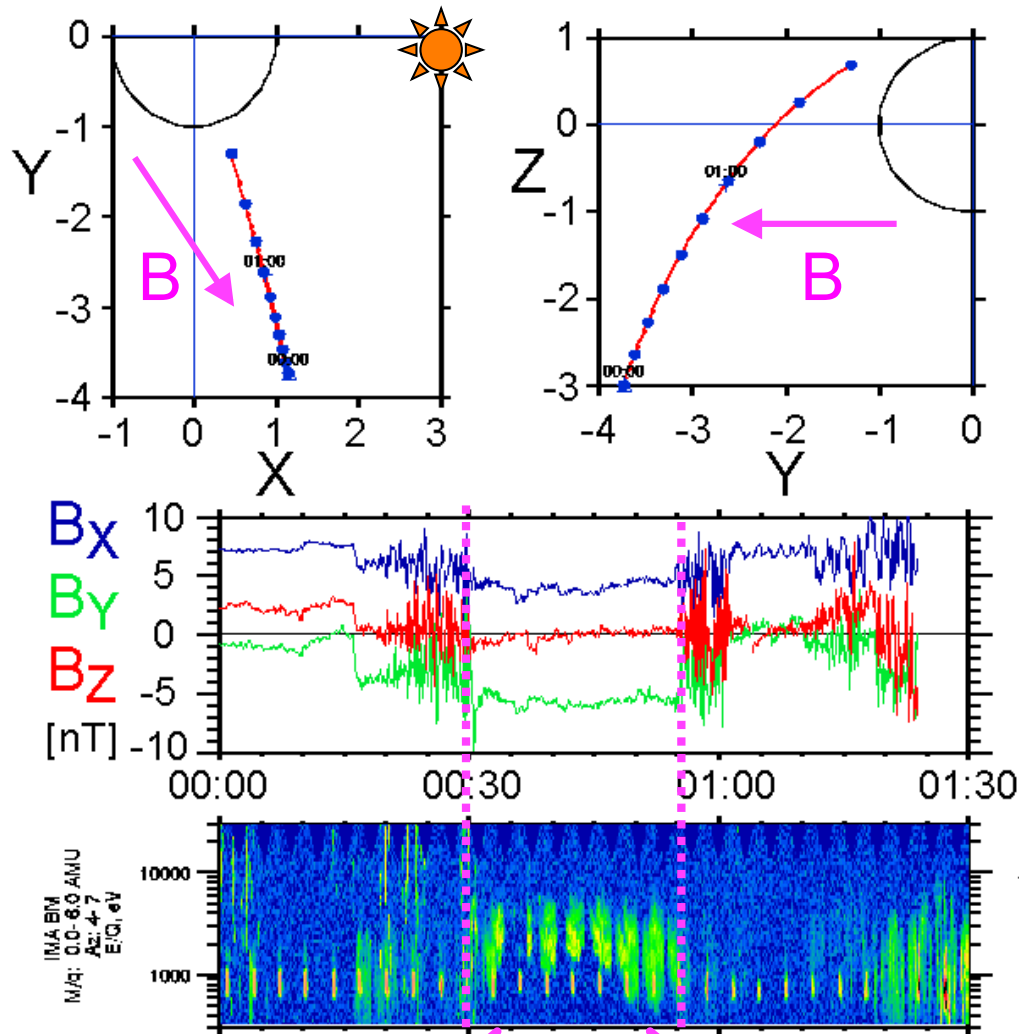
MEX/IMA 2005-8-5, 1019:12~1022:24 UT



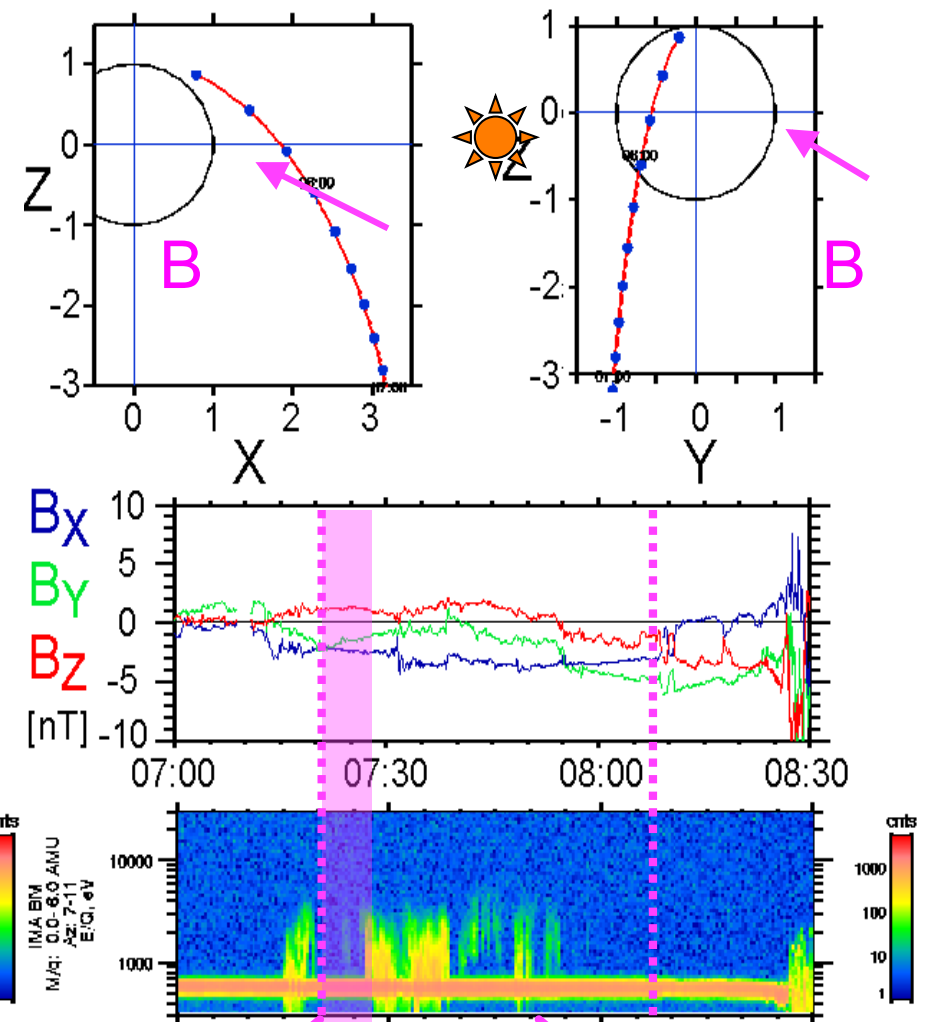
Quasi- $\perp$ (case 1+3)



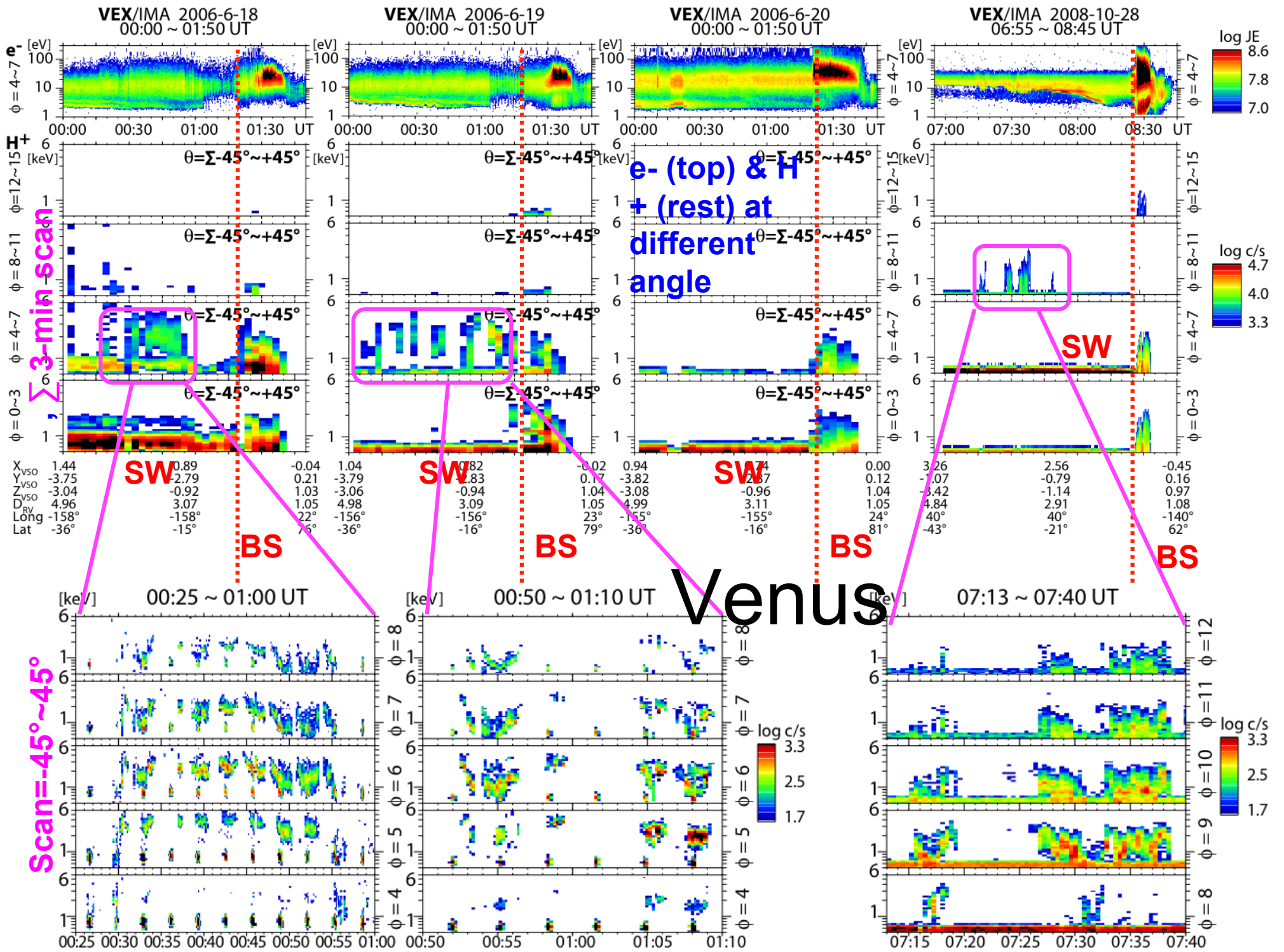
Venus $\approx$ Earth



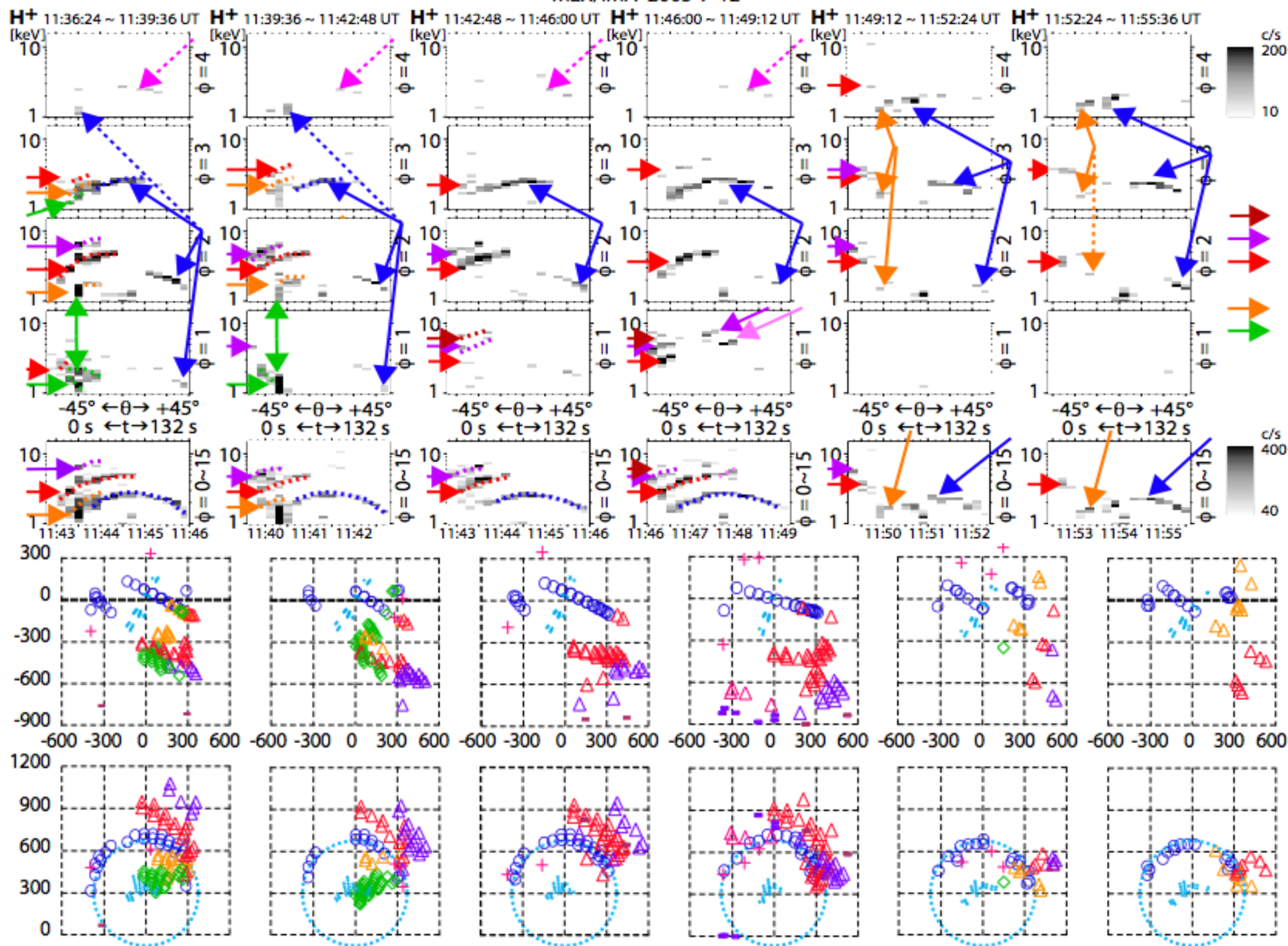
2006-6-18 connected to BS
Beam = foreshock



2008-10-28 connected to BS
Sometimes no beam in foreshock



MEX/IMA 2005-7-12



5th Alfvén Conference

on **“Plasma Interaction with Non-magnetized Planets/Moons and its Influence on Planetary Evolution”**

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SC-1 CIS/HIA 2007-1-18

